

Connecting Large Scale Structures to galactic spin

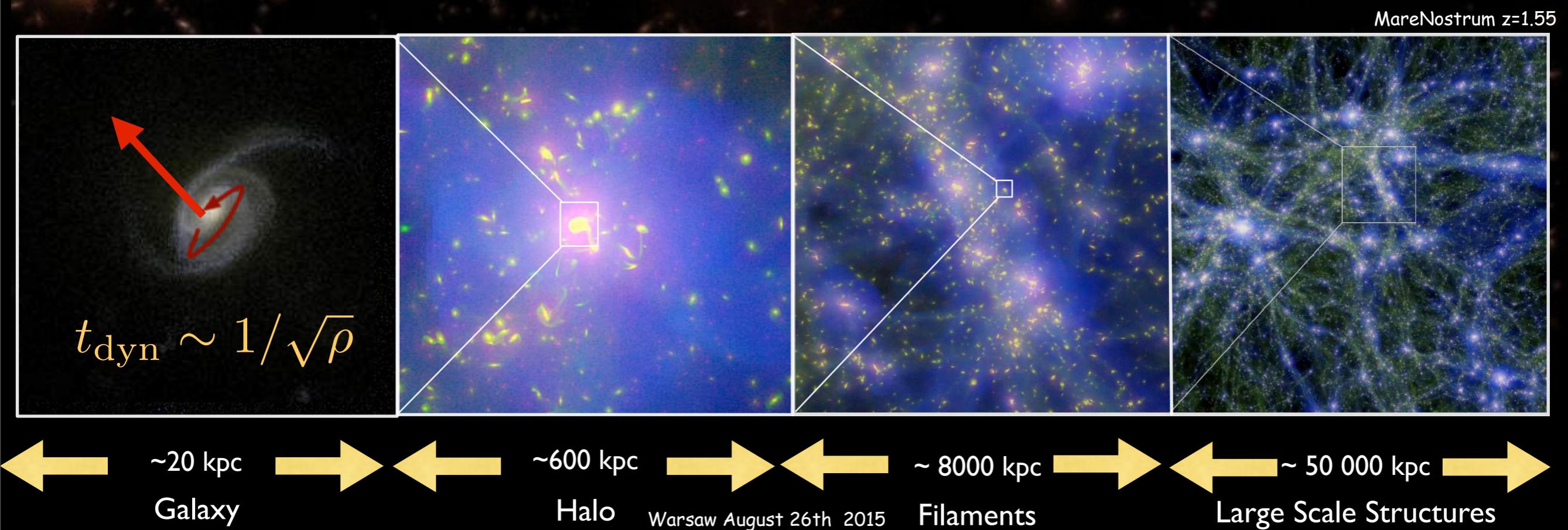
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Institut d'astrophysique de Paris

S. Codis, C. Laigle, C. Welker T. Kimm D. Pogosyan, J. Devriendt, Y Dubois+ Horizon/Spin(e) Collaboration



Can we predict the spin of galaxies **on** the cosmic web from first principles?



Outline

- What is the geometry of spin near saddle?
- How do dark halo's spin flip relative to filament
- Why does it induce a transition mass:
Eulerian & Lagrangian theory?

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Why do we care?

- Weak lensing
- AM stratification drives **morphology**
- Galaxy formation is not a 1D manifold
- Because we can understand something !

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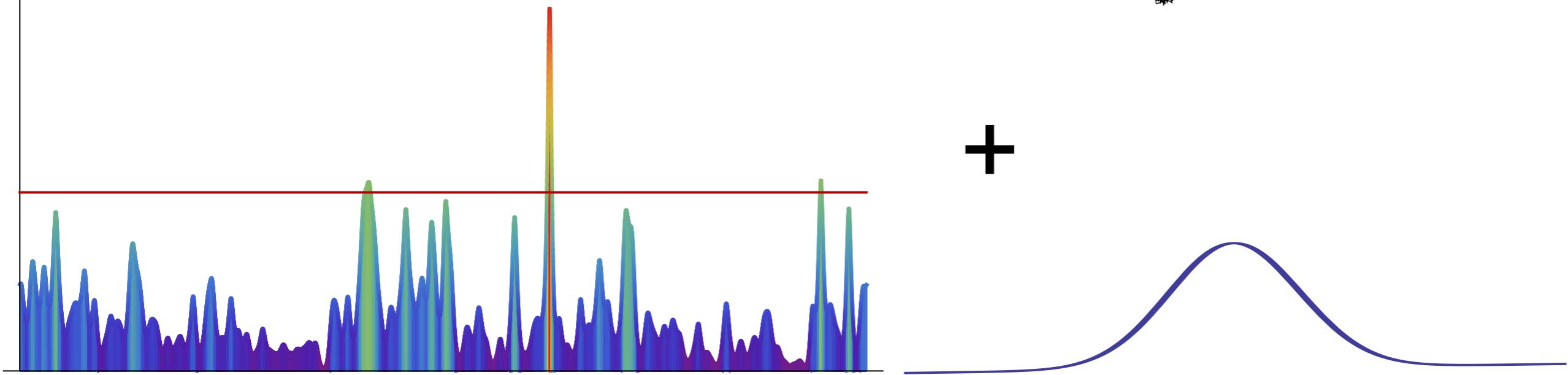
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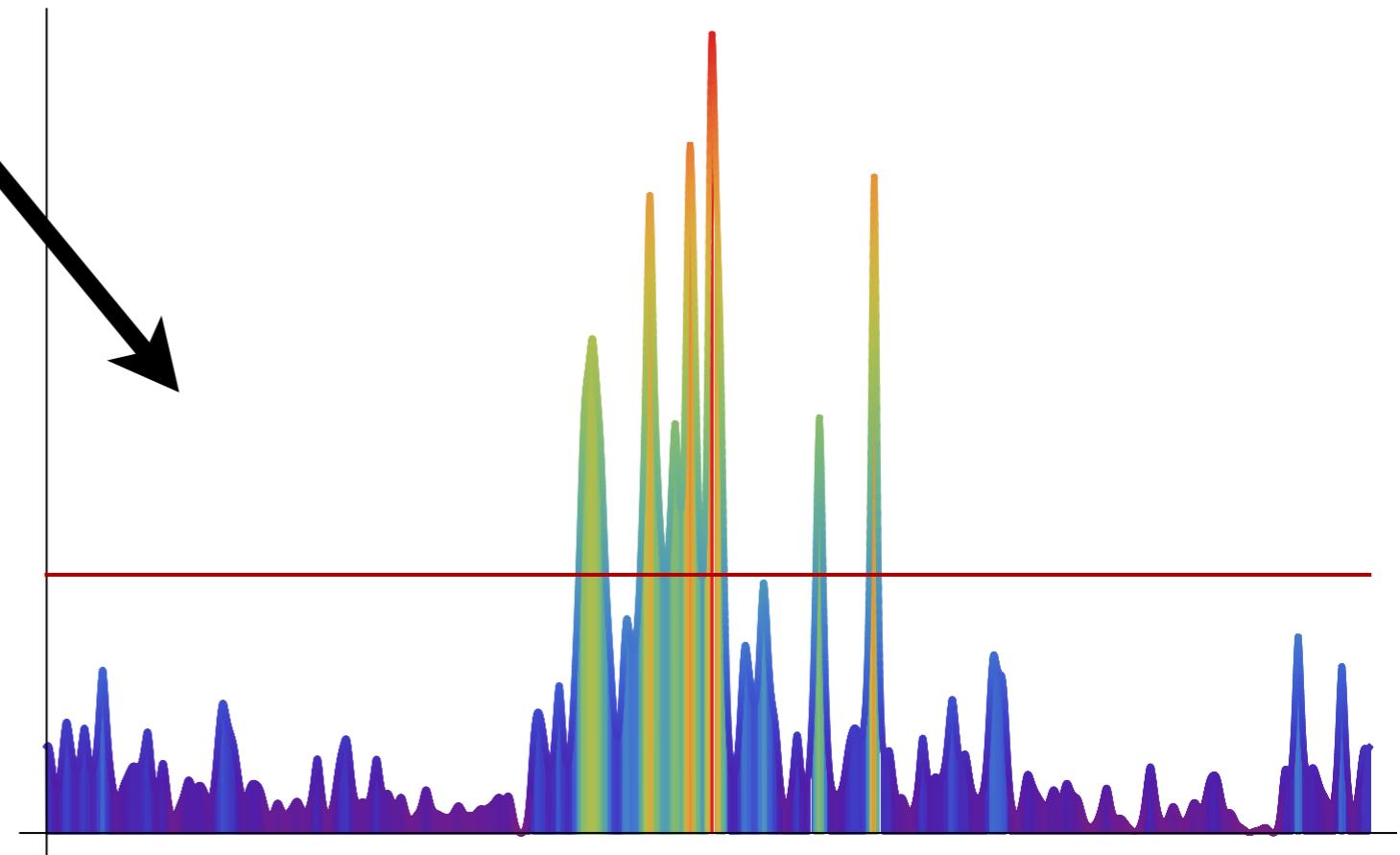
Where **galaxies** form does matter, and can be traced back to ICs.

*Flattened filaments generate point-reflection-symmetric AM/vorticity distribution:
they induce the observed spin transition mass*

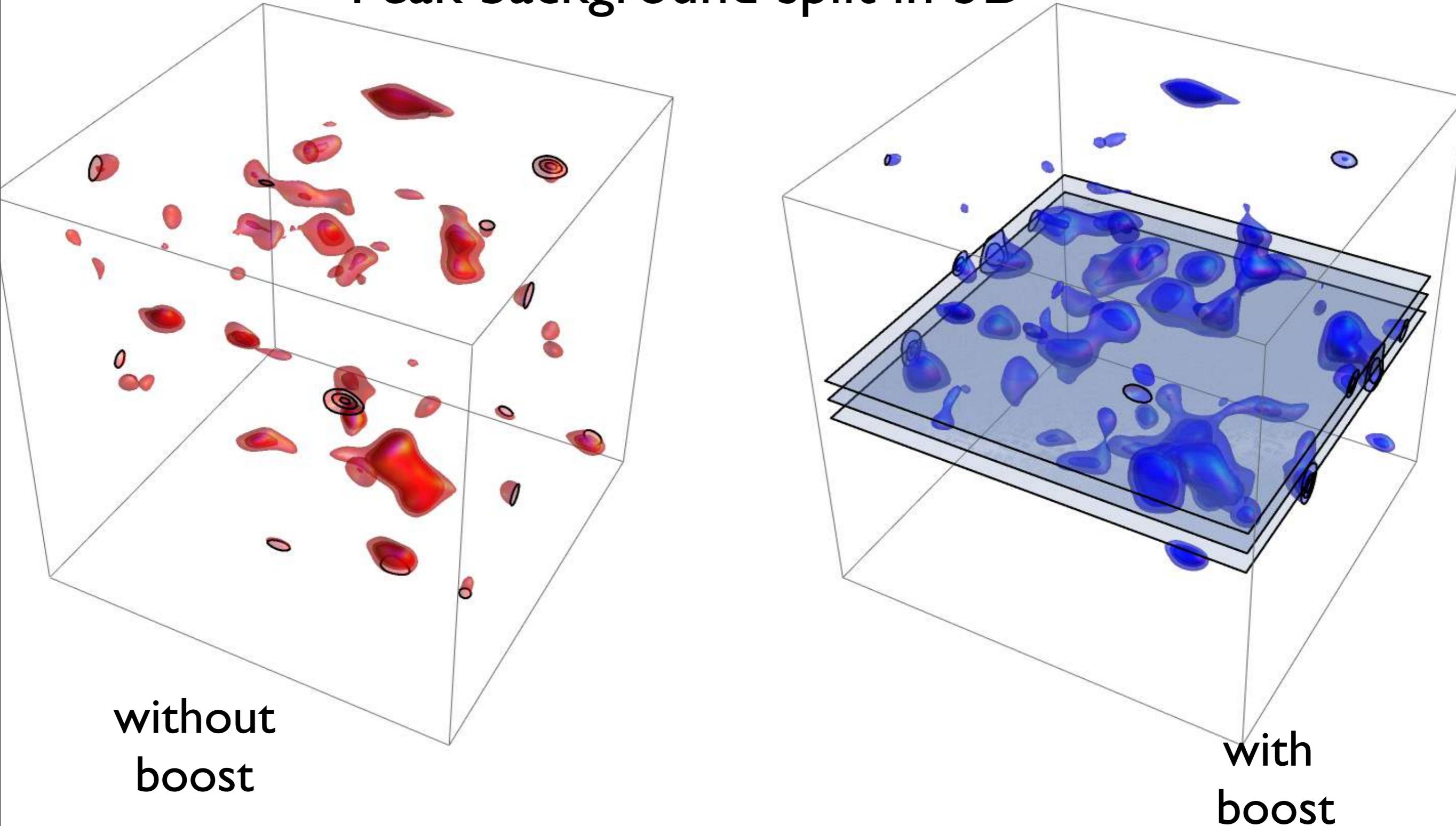
dark halos don't form anywhere



Peak background split
(PBS) in 1D

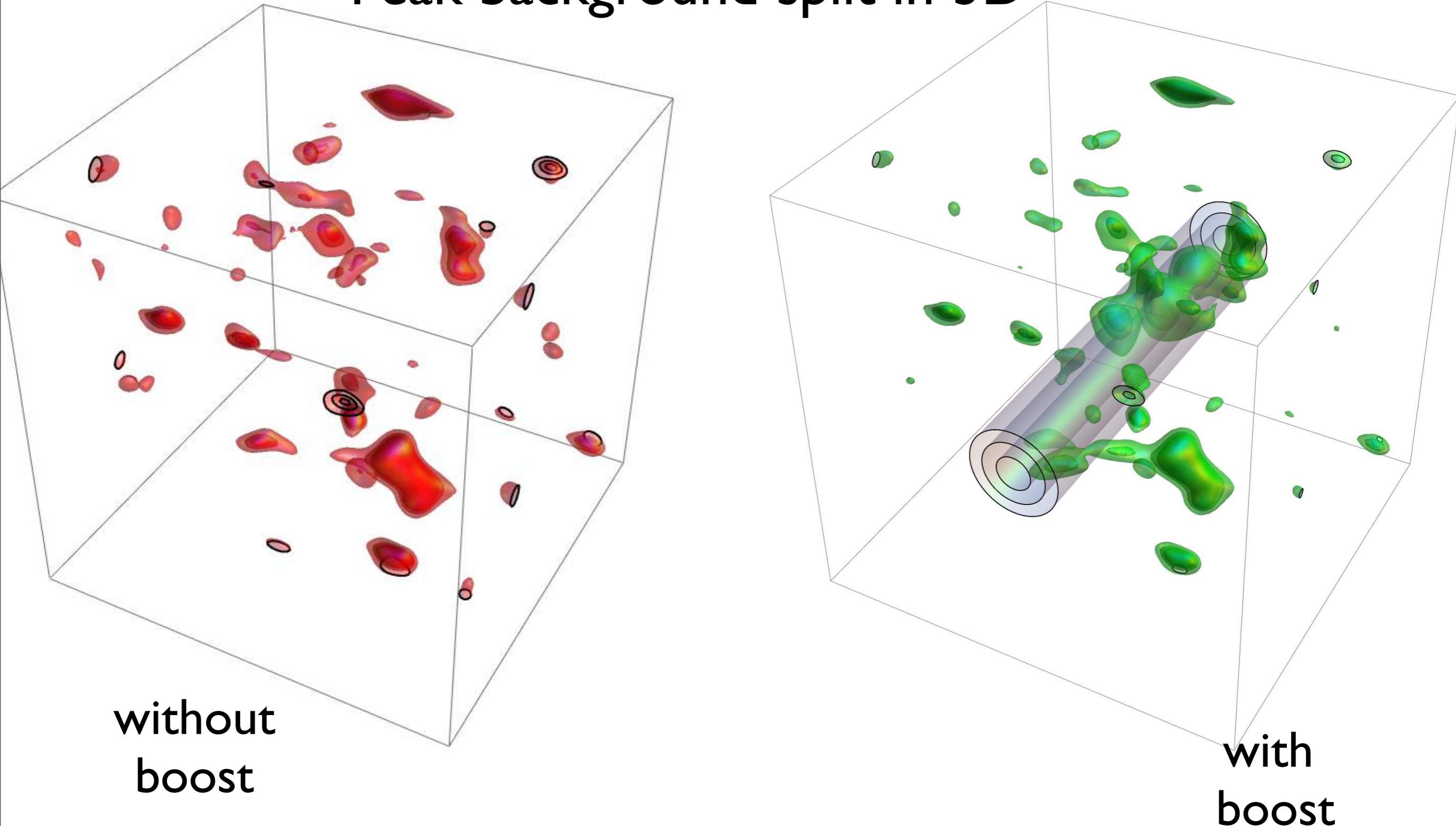


Peak background split in 3D



**Does this anisotropic biassing have
a dynamical signature? yes! in term of spin!**

Peak background split in 3D



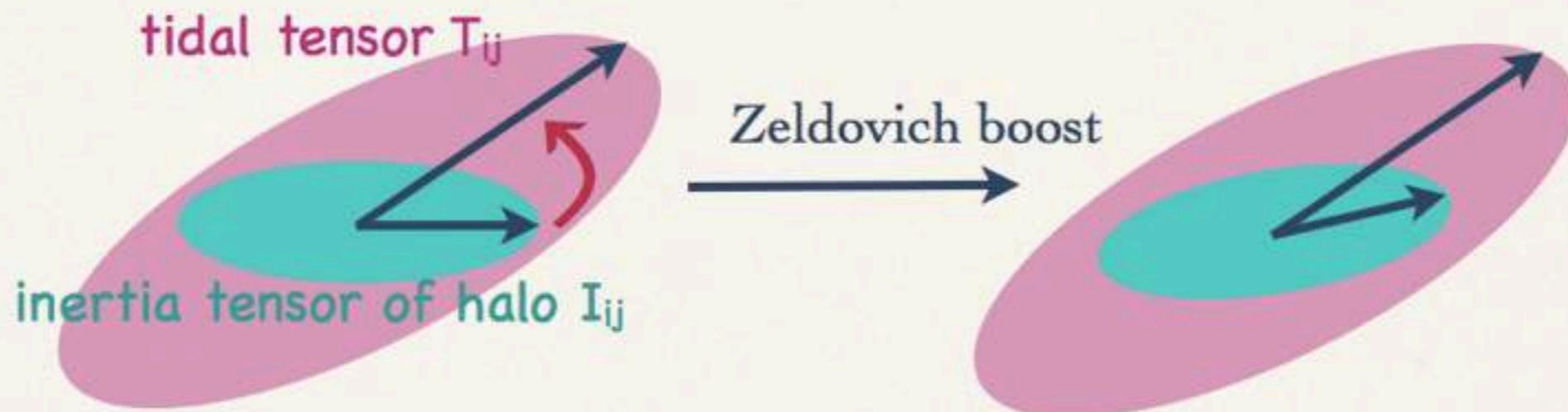
**Does this anisotropic biassing have
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Tidal Torque Theory in one cartoon

Can we understand where spin and vorticity alignments come from?

-usual tidal torque theory

$$L_k = \varepsilon_{ijk} I_{li} T_{lj}$$



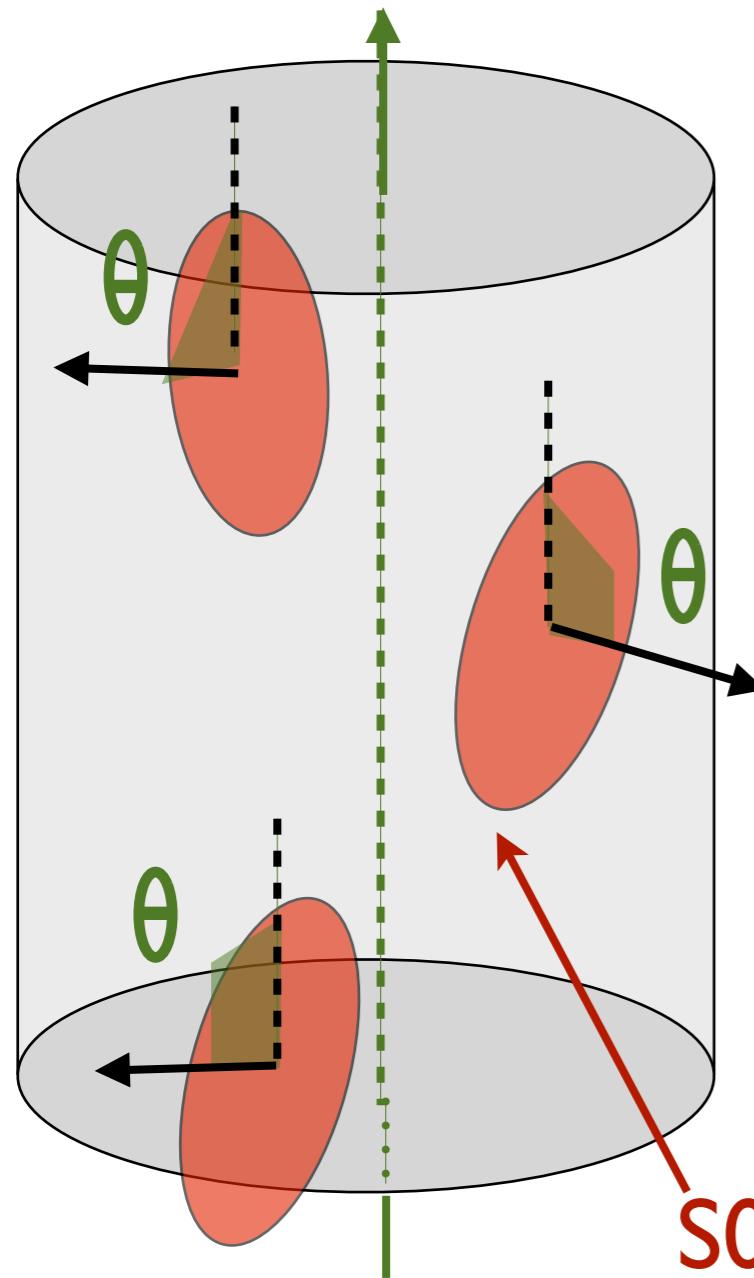
YES! via conditional TTT subject to PBS

"Cond' prob, good stuff!"

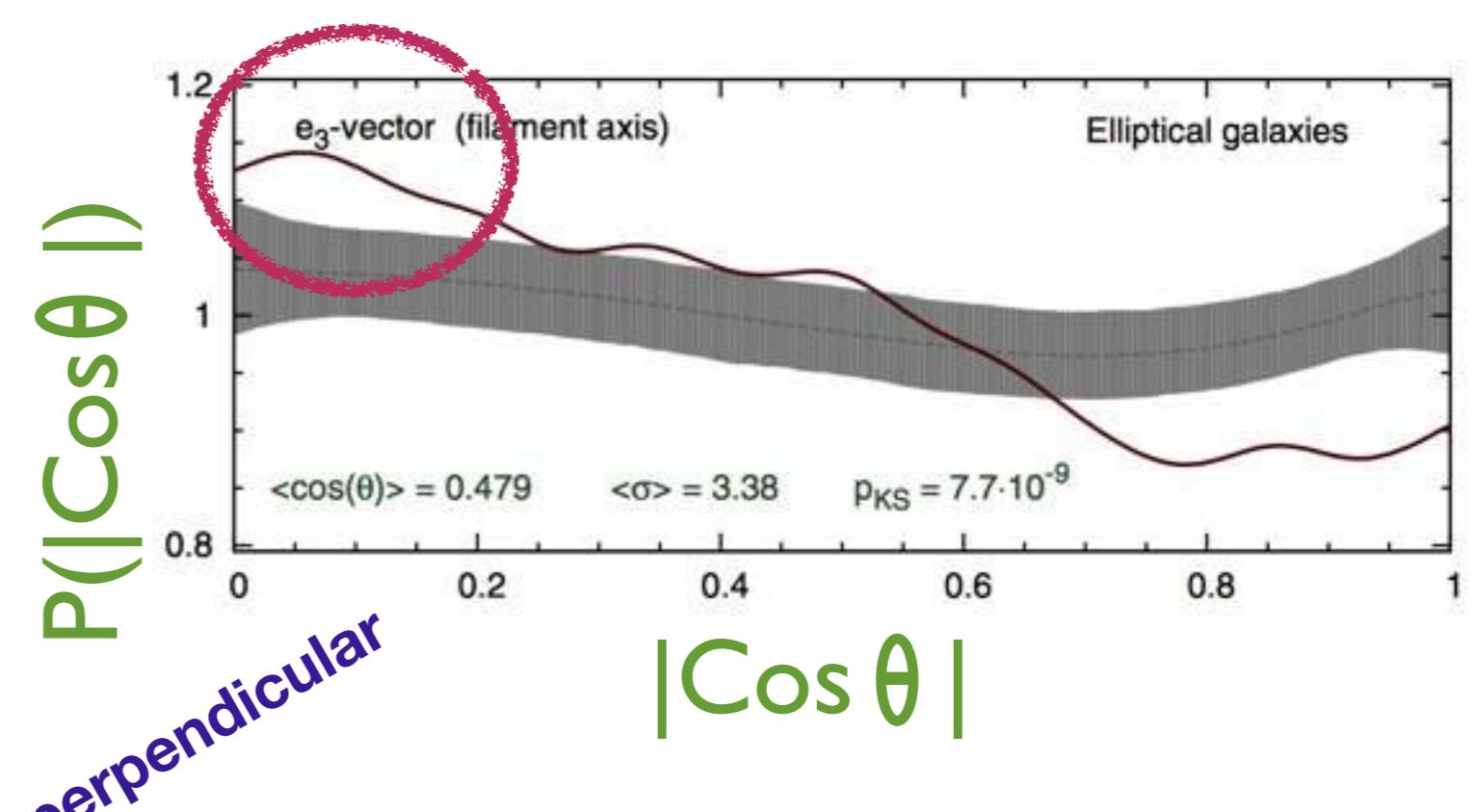
Et Voilà !

Evidences of galaxy spin - filament alignment

Cosmic Filament



Tempel+ (2013) in the SDSS



perpendicular

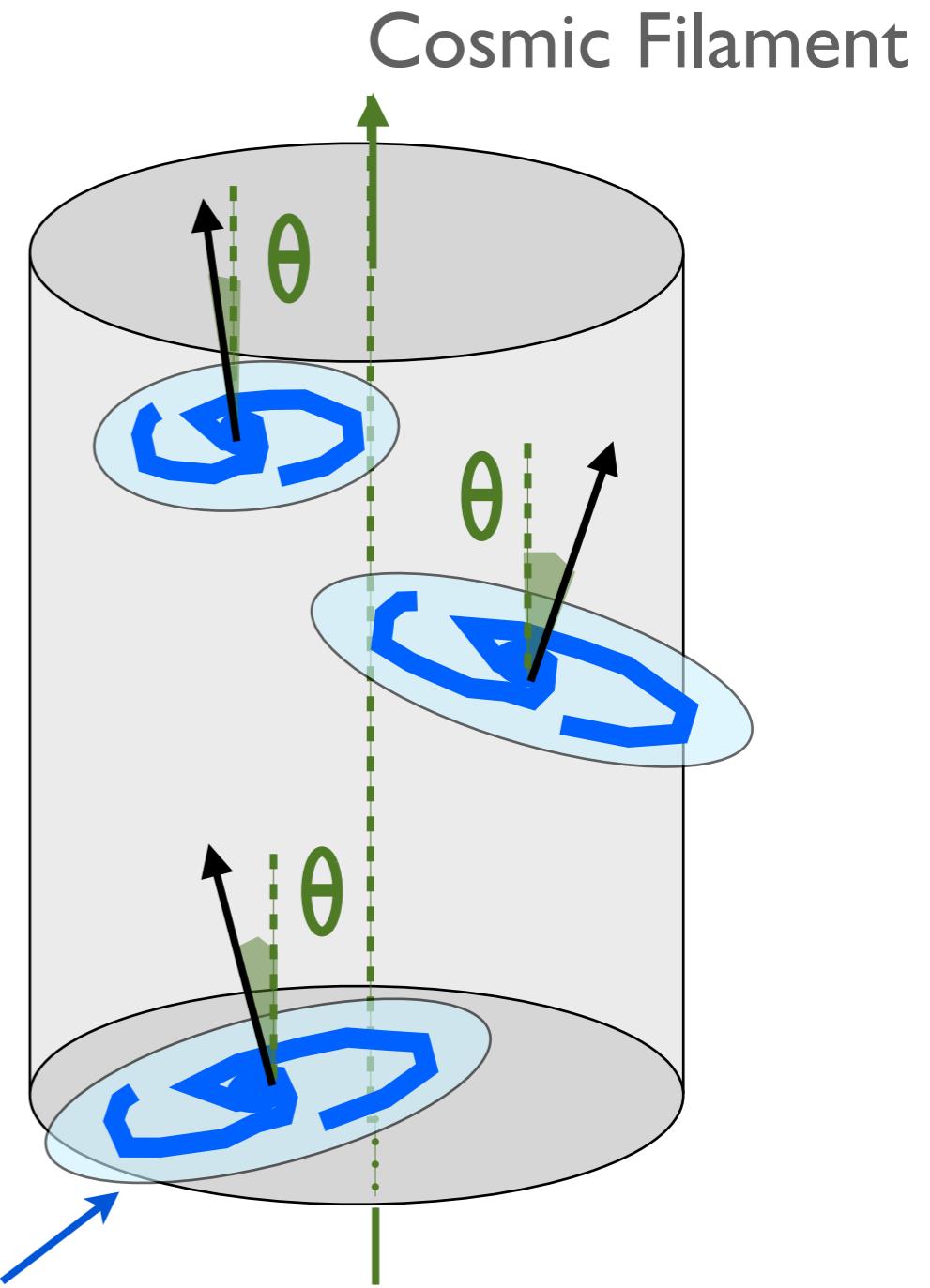
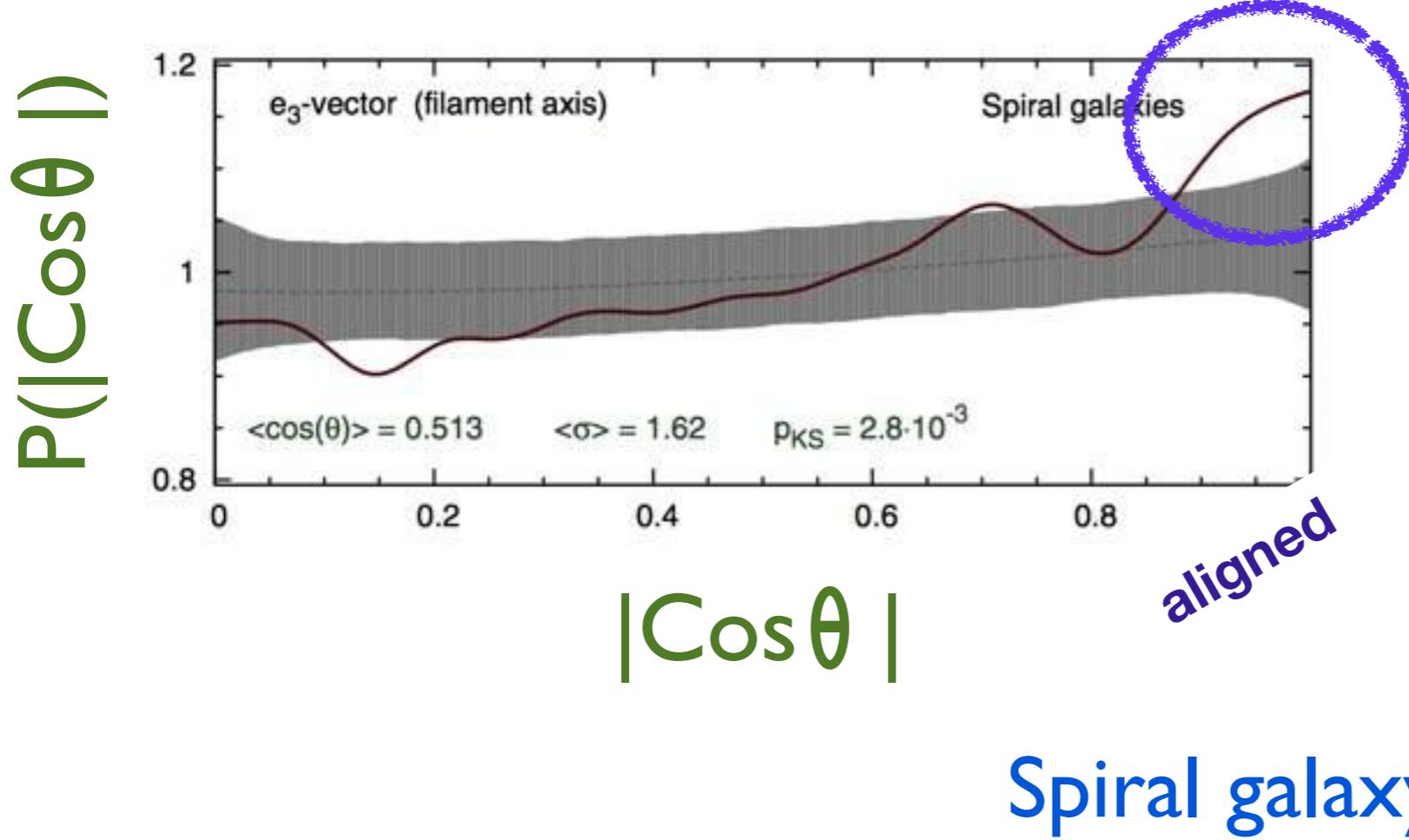
S0 galaxy

See also:

Aragon-Calvo+ 2007, Hahn+ 2007, Paz+ 2008, Zhang+ 2009, Codis+ 2012, Libeskind+ 2013, Aragon-Calvo 2013, Dubois+ 2014

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Orientation of the spins w.r.t the filaments

Horizon 4Pi:

DM only

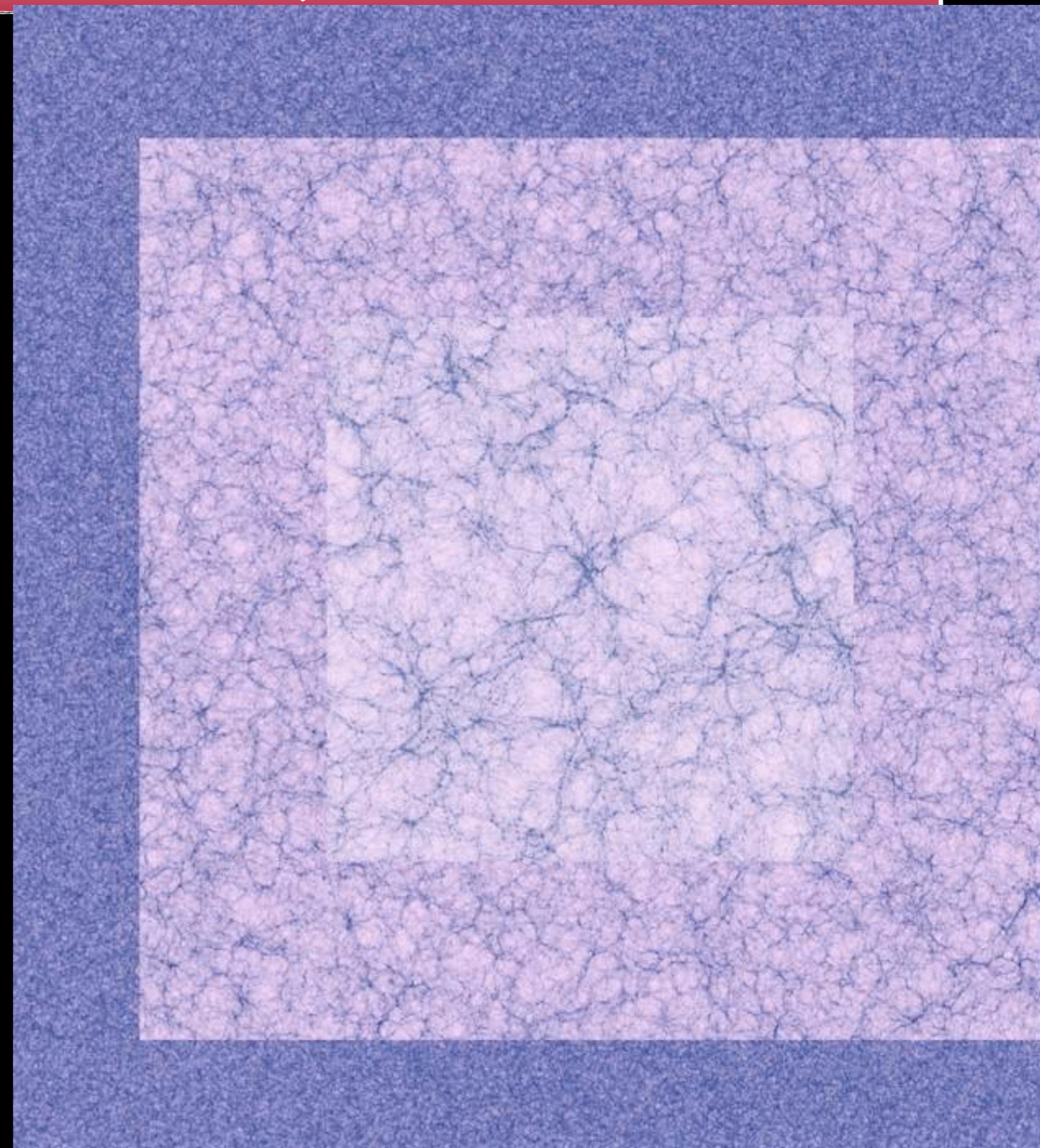
2 Gpc/h periodic box

4096^3 DM part.

43 million dark halos at
 $z=0$

(Teyssier et al, 2009)

10 000 000 hrs CPU



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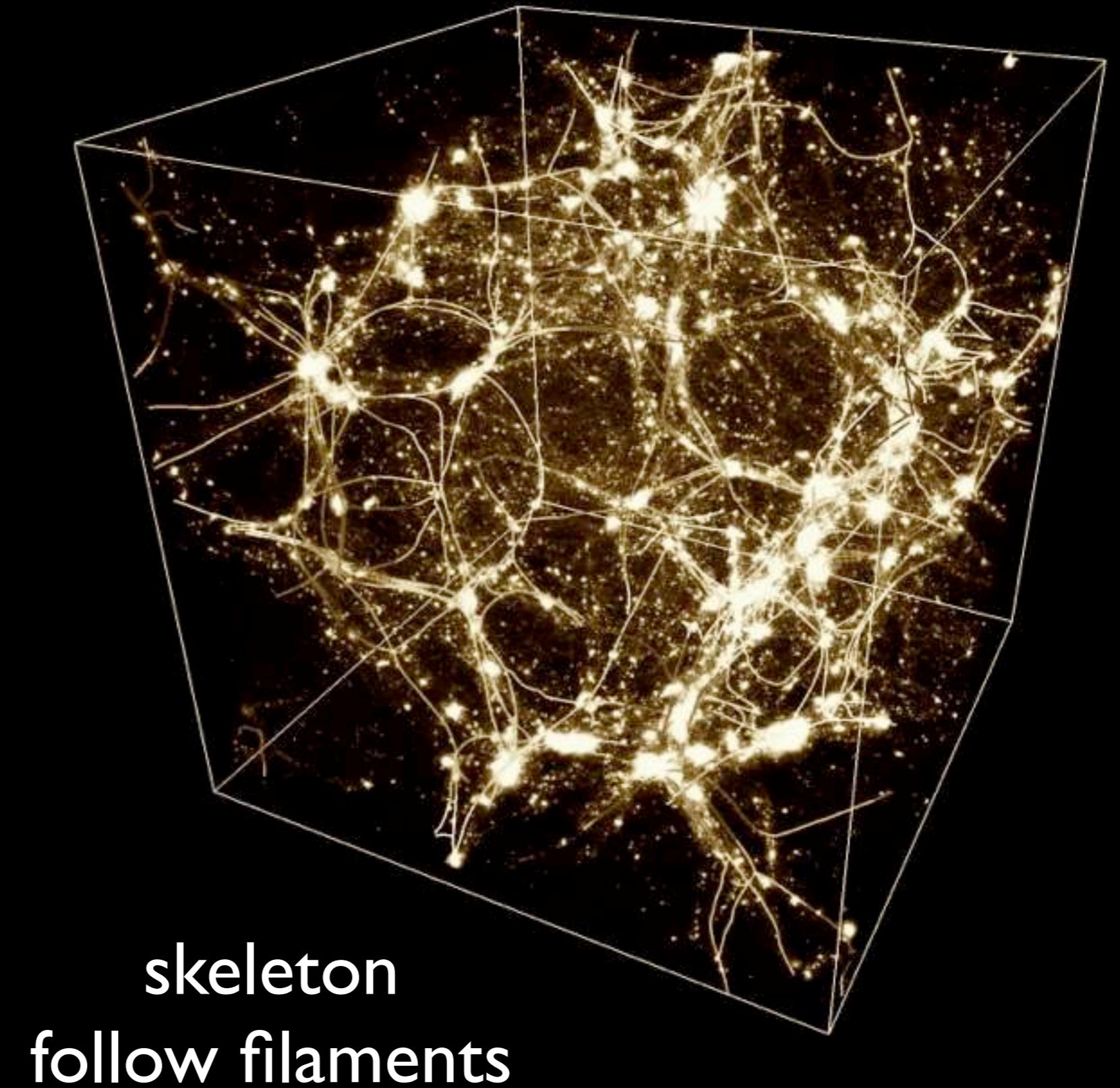
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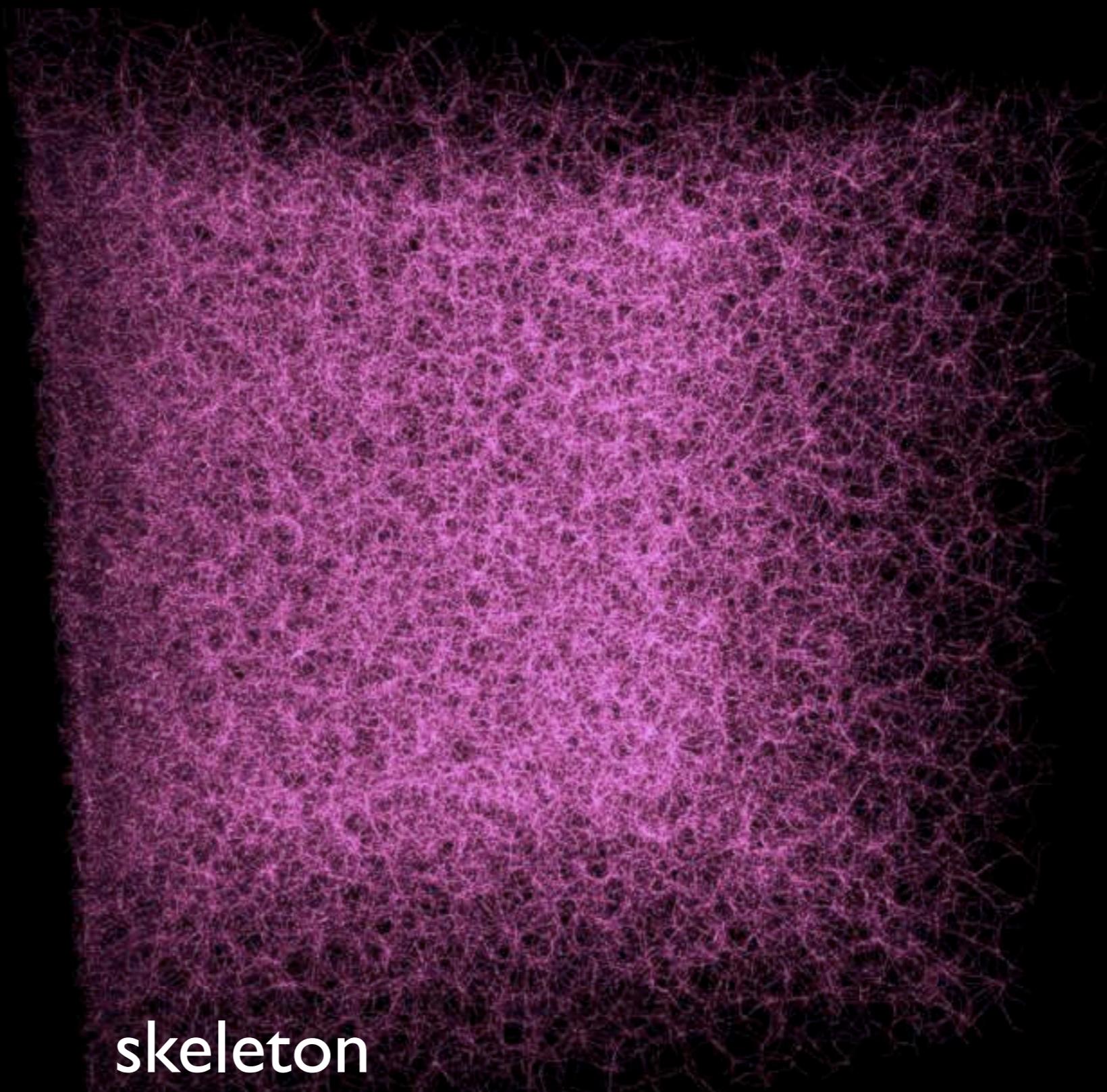
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skeleton
follow filaments



Excess probability of alignment between the spins and their host filament

mass transition:

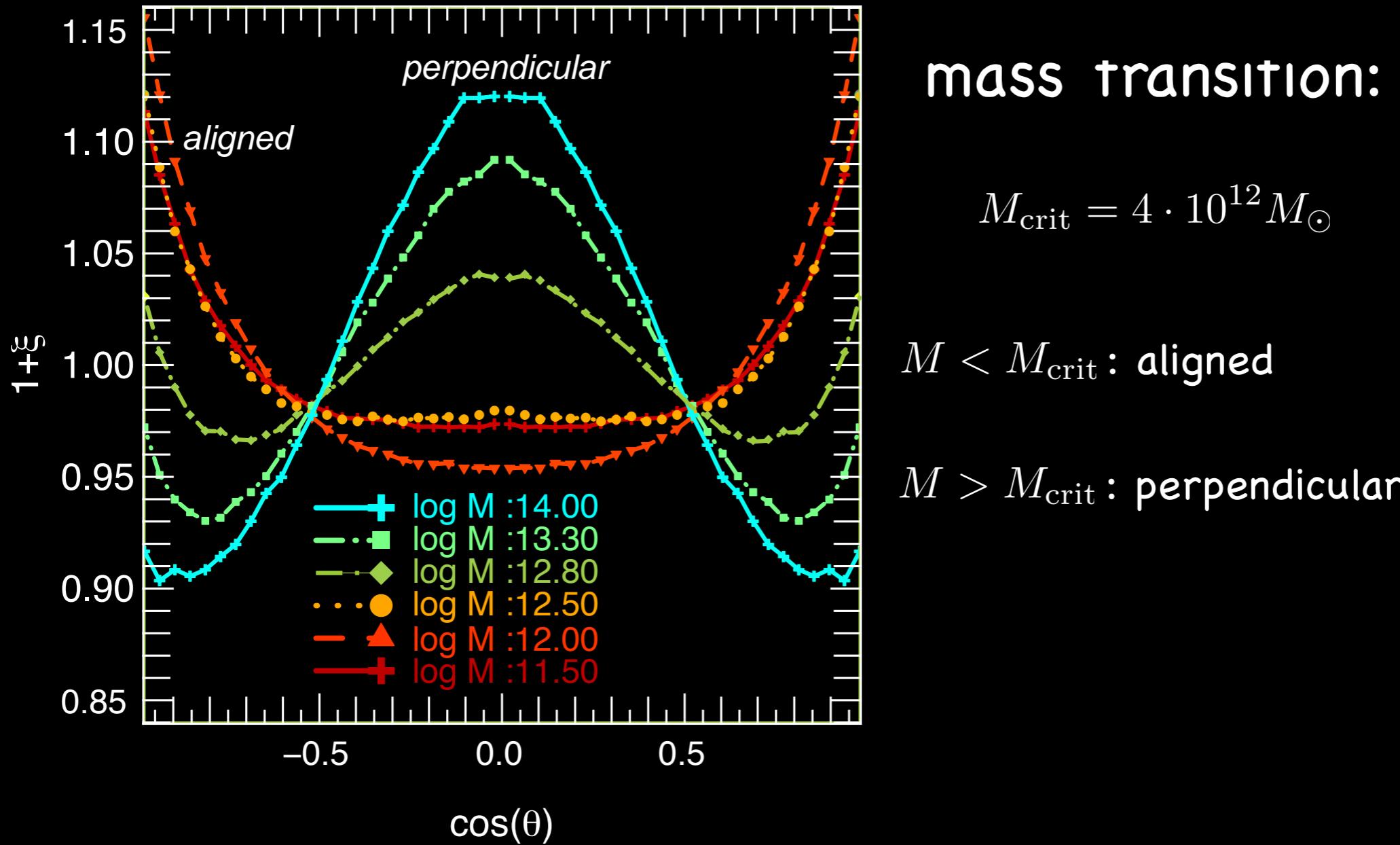
$$M_{\text{crit}} = 4 \cdot 10^{12} M_{\odot}$$

$M < M_{\text{crit}}$: aligned

$M > M_{\text{crit}}$: perpendicular

(Codis et al, 2012)

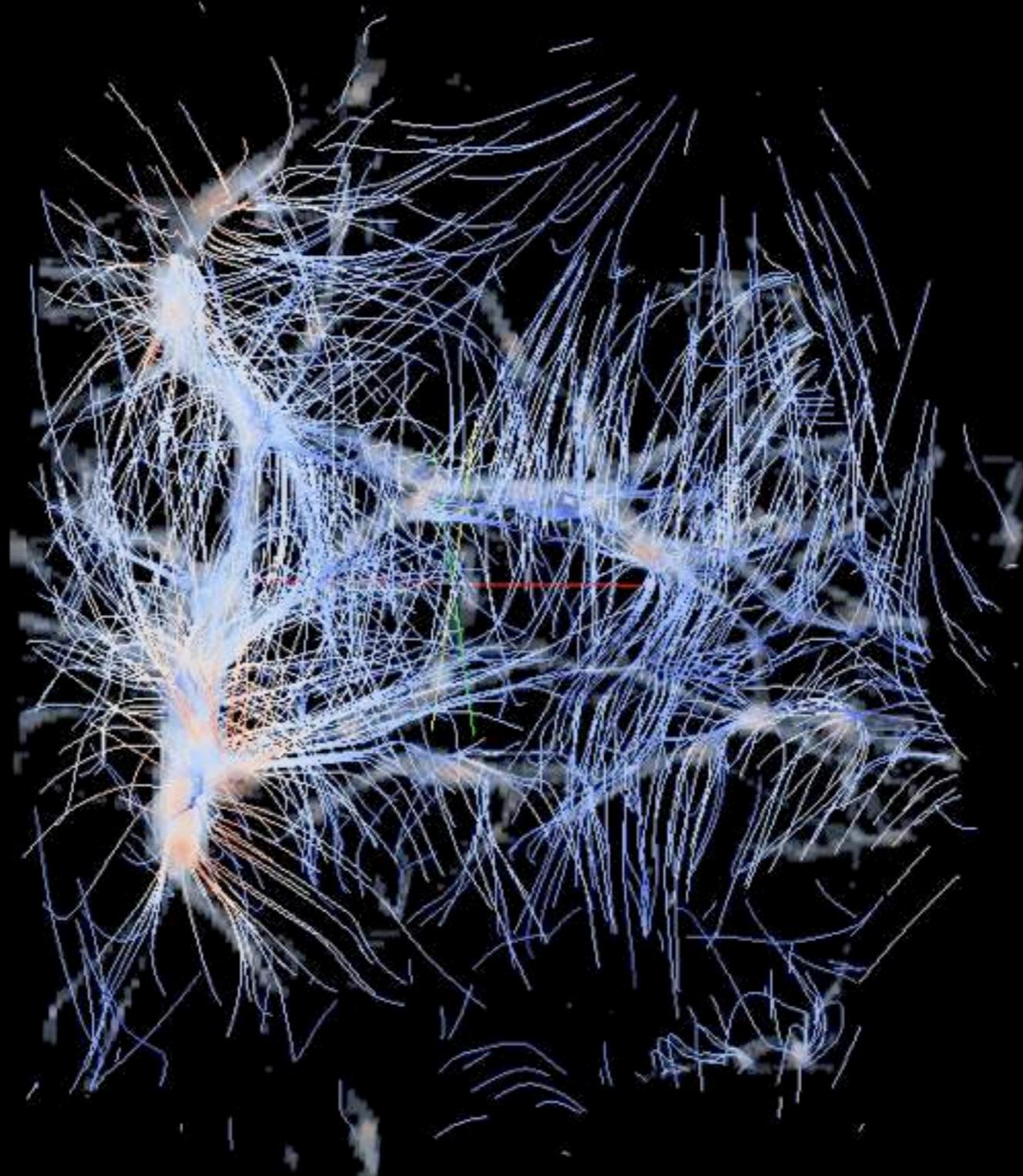
Excess probability of alignment between the spins and their host filament



(Codis et al, 2012)

How does the formation of the filaments generate spin parallel to them?

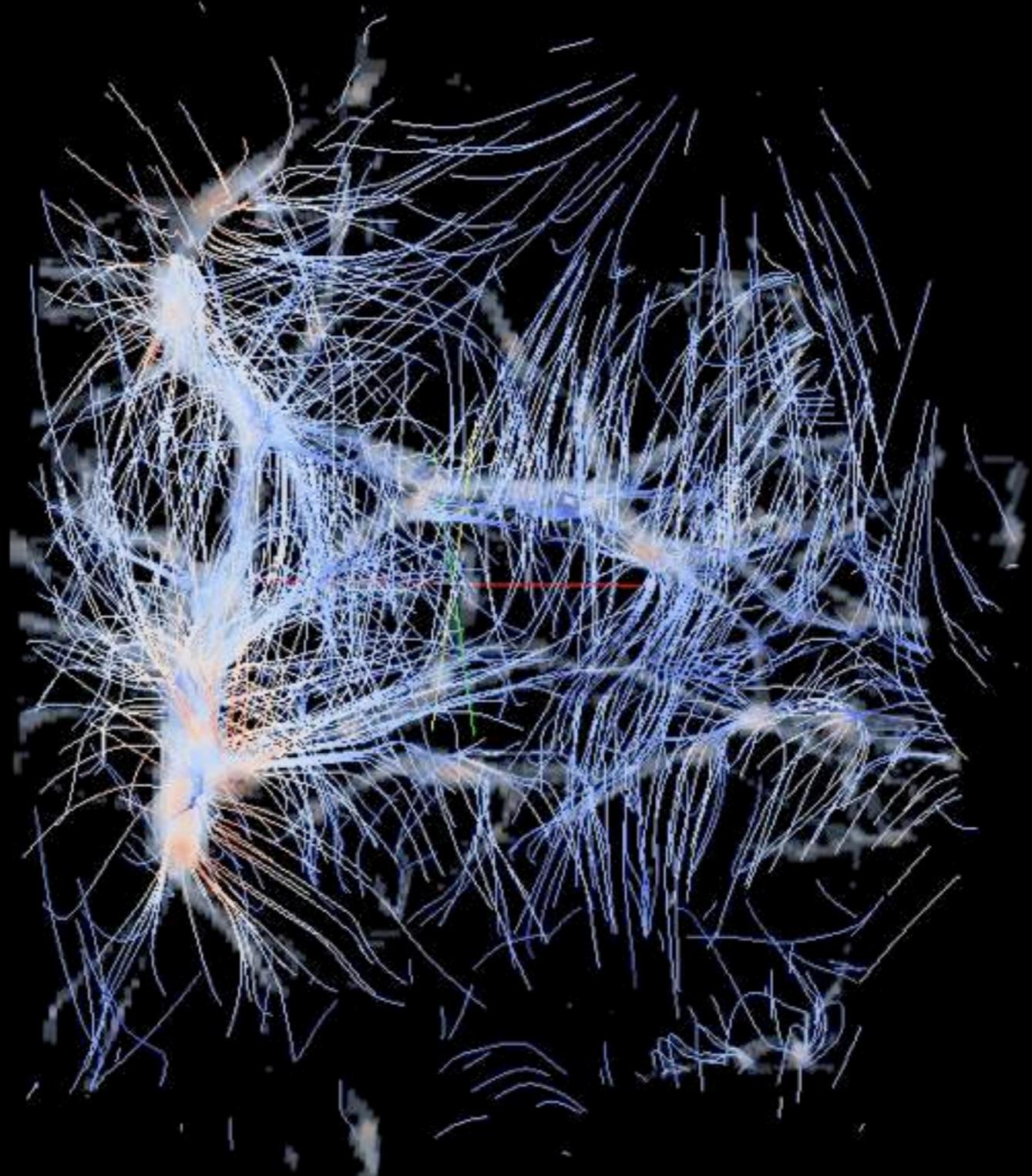
Voids/wall saddle
repel...



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Voids/wall saddle
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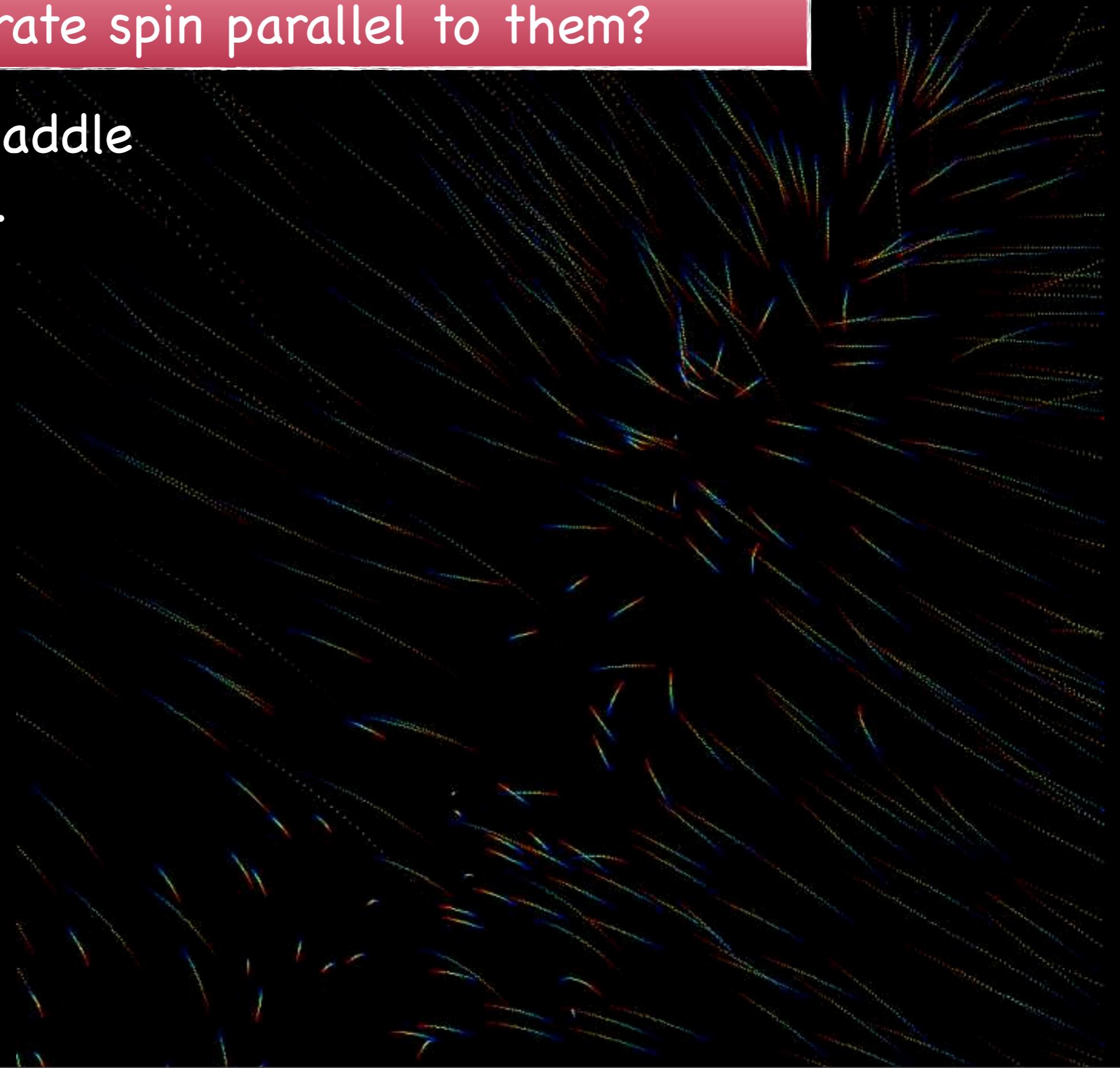
winding of walls
into filaments



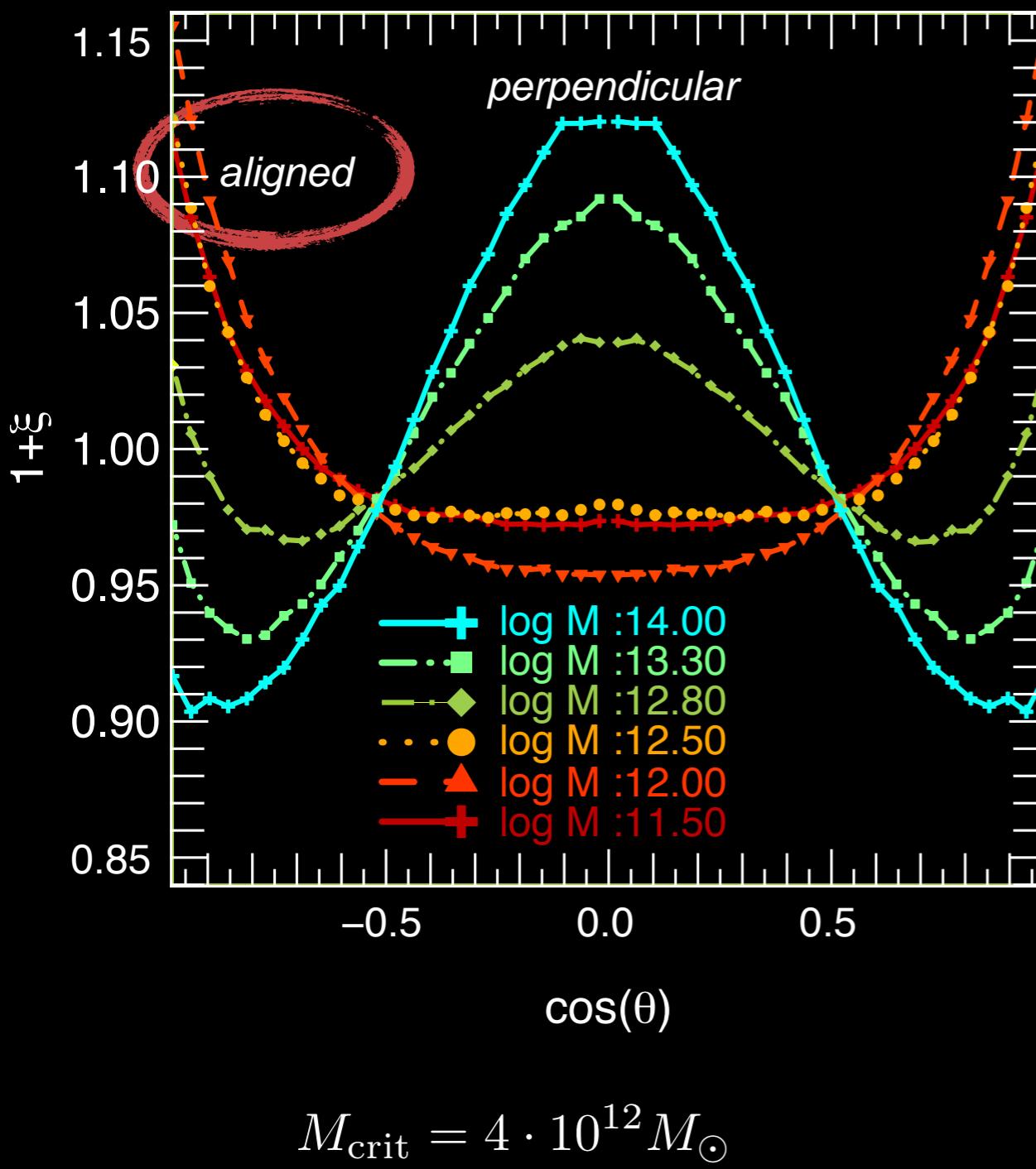
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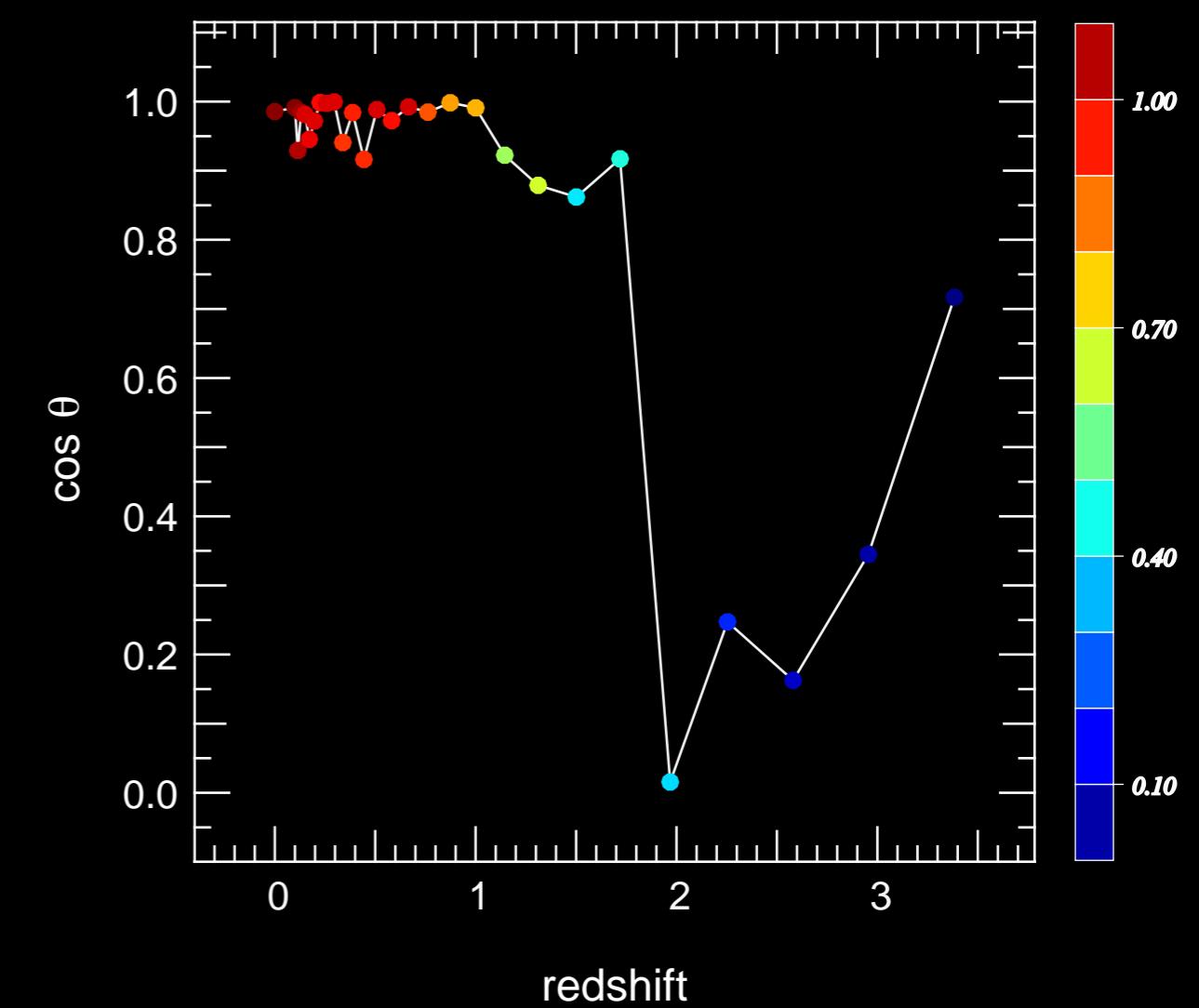


Low-mass haloes: $M < M_{\text{crit}}$

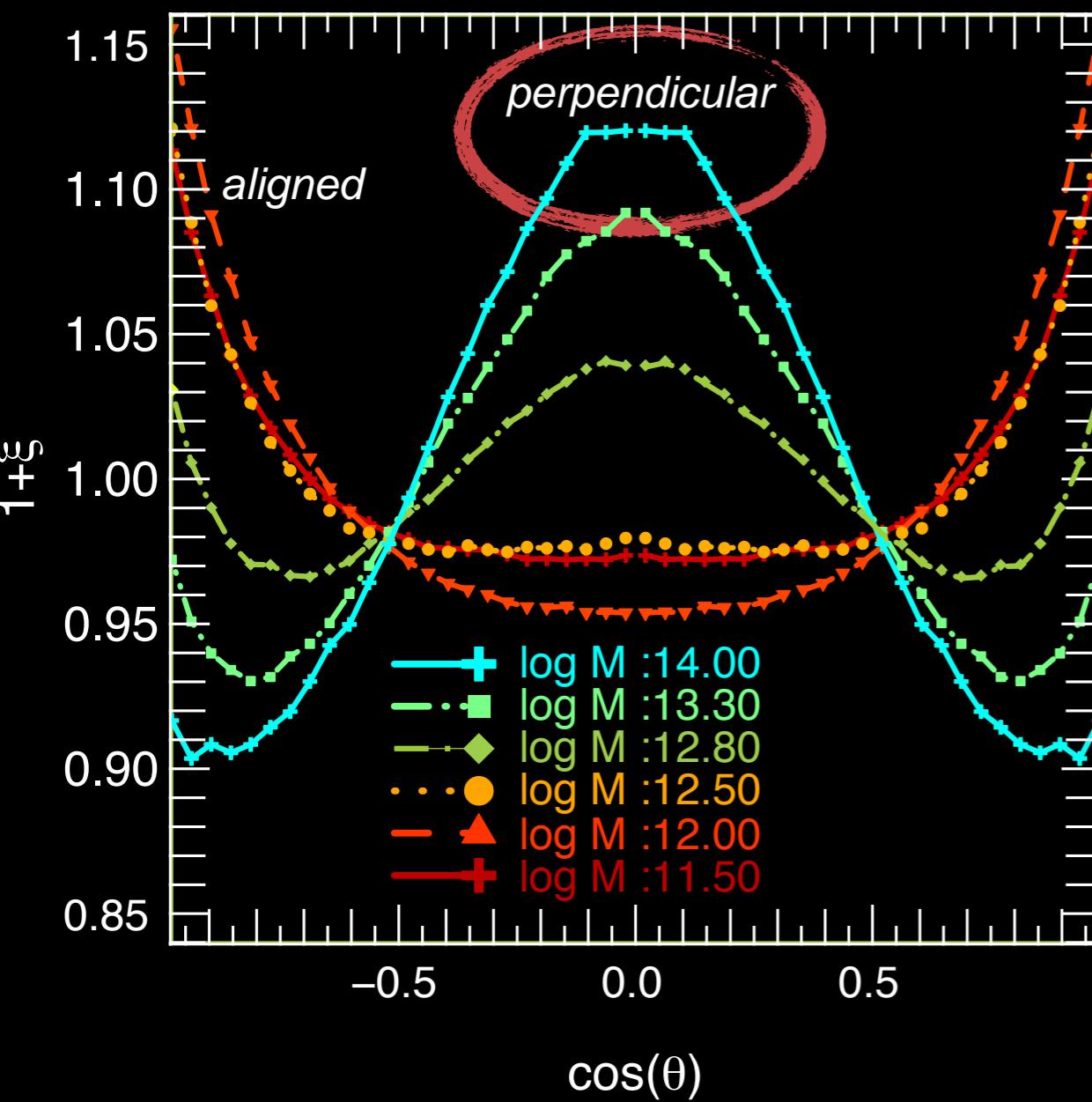


-formed at high z during the formation within filaments

-no major merger but smooth accretion until $z=0$

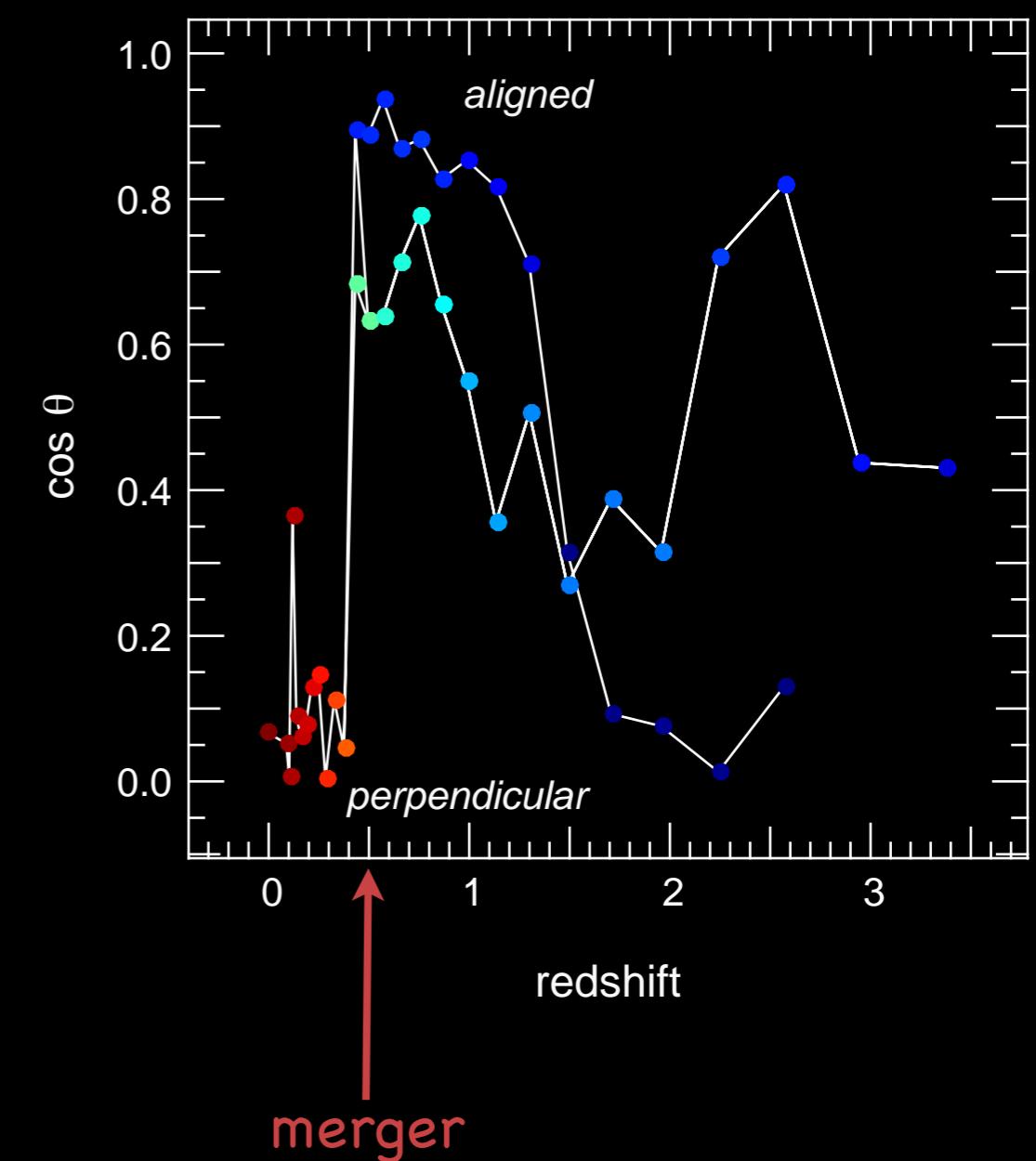


High-mass haloes: $M > M_{\text{crit}}$



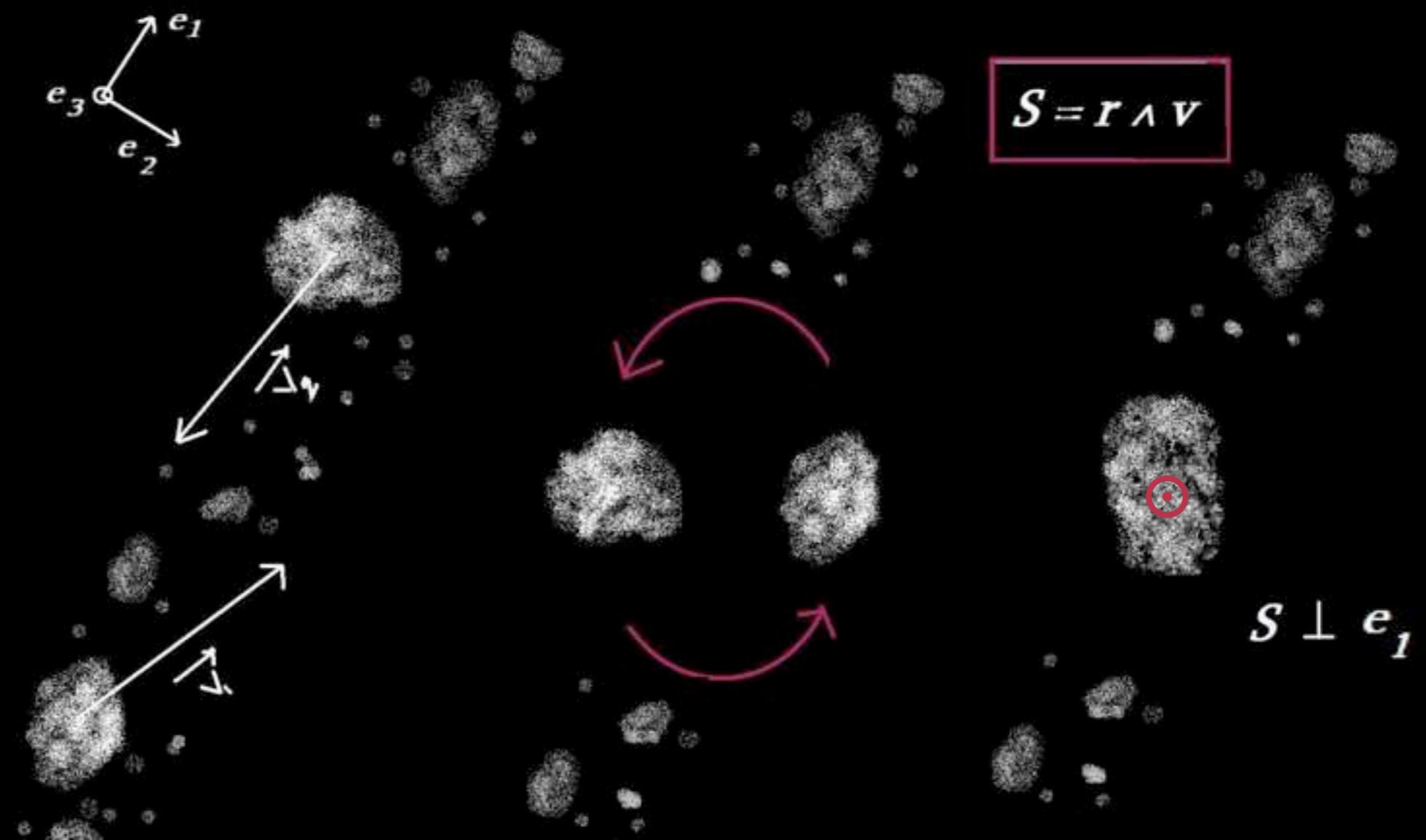
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formed at low z by mergers inside the filaments

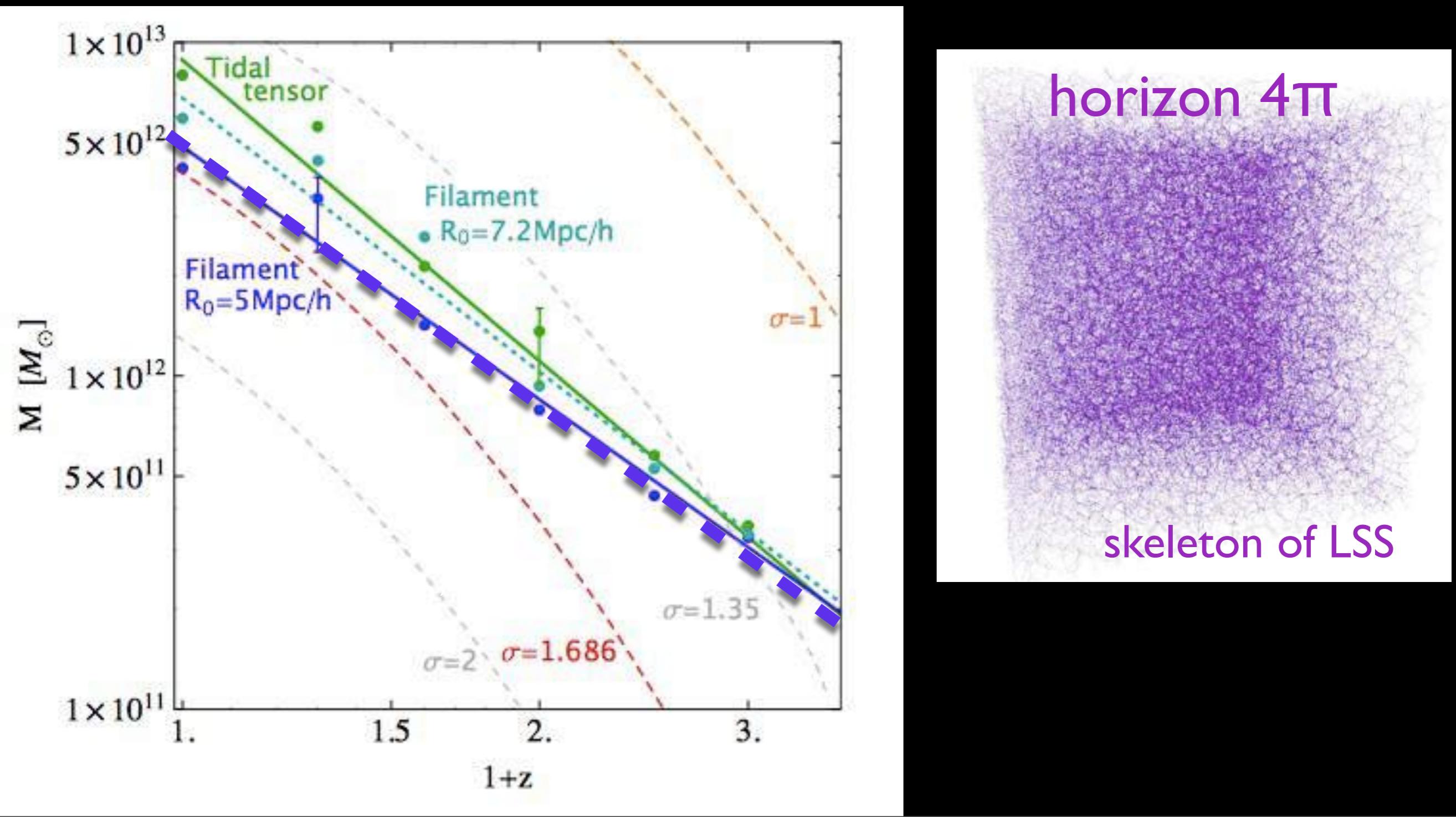


How do mergers along the filaments create spin perpendicular to them?

Halos catch up
with each other
along the filaments

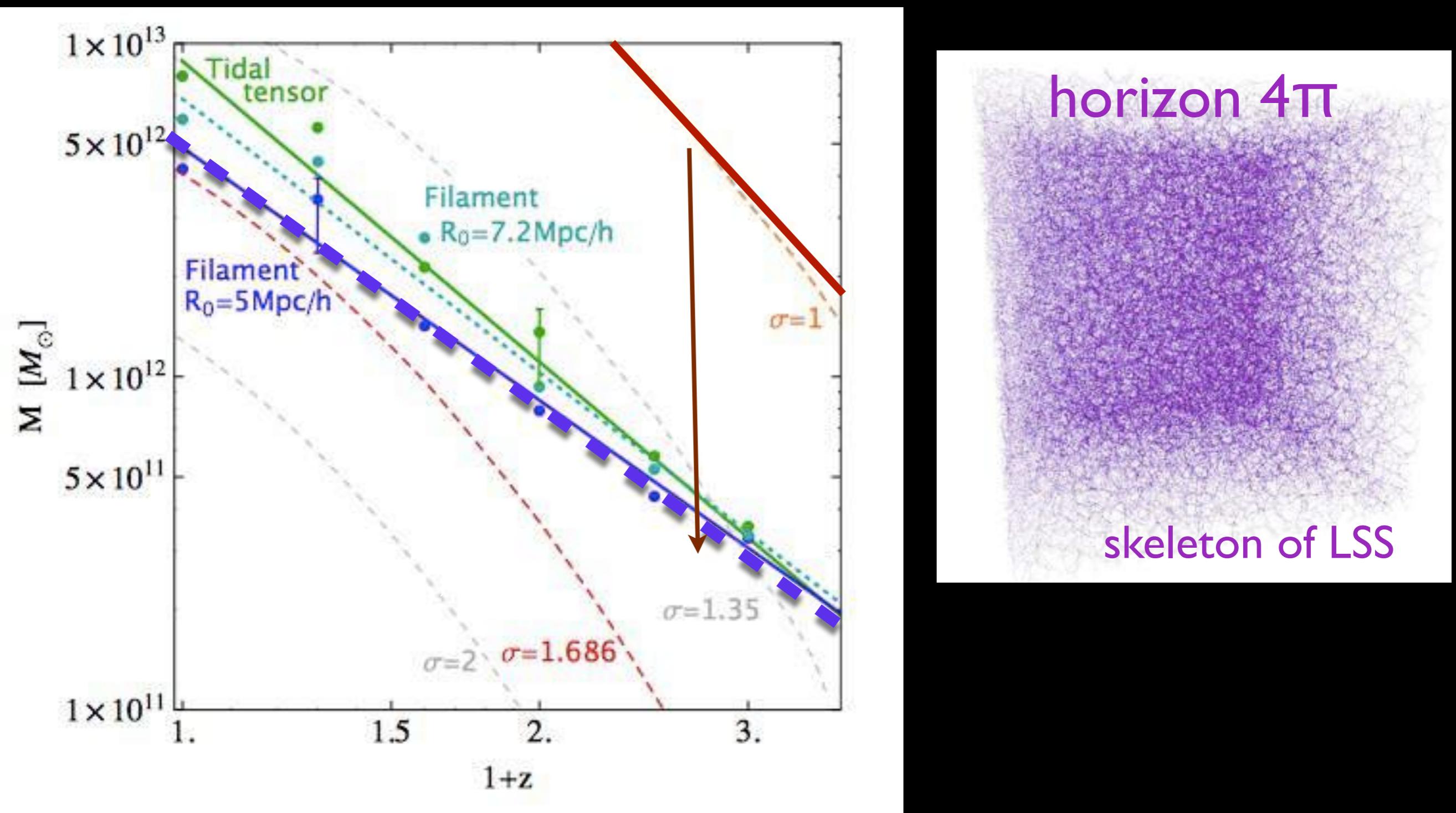


Explain transition mass?

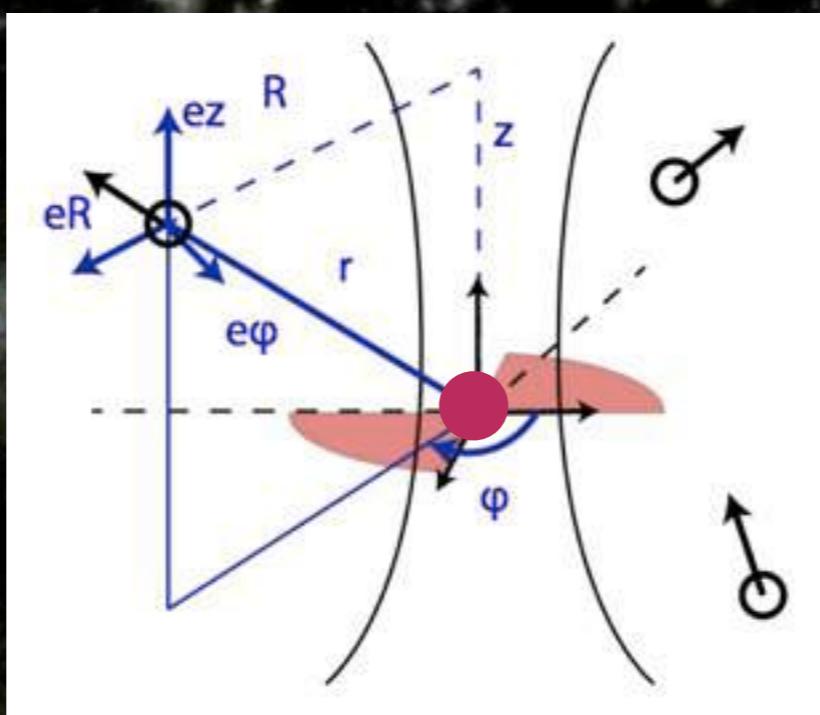


Explain transition mass?

Transition mass versus redshift: **what's wrong???**



Tidal torque theory with a peak background split near a saddle



Warsaw August 26th 2015

⦿ The Idea

- ⦿ walls/filament/peak locally bias differentially tidal and inertia tensor: spin alignment reflect this in TTT

⦿ The picture

- ⦿ Geometry of spin near saddle: point reflection symmetric distribution, 1/10 of 'naive size'

⦿ The Maths

- ⦿ Very simple **ab initio** prediction for **mass** transition

The Lagrangian view of spin/LSS connection

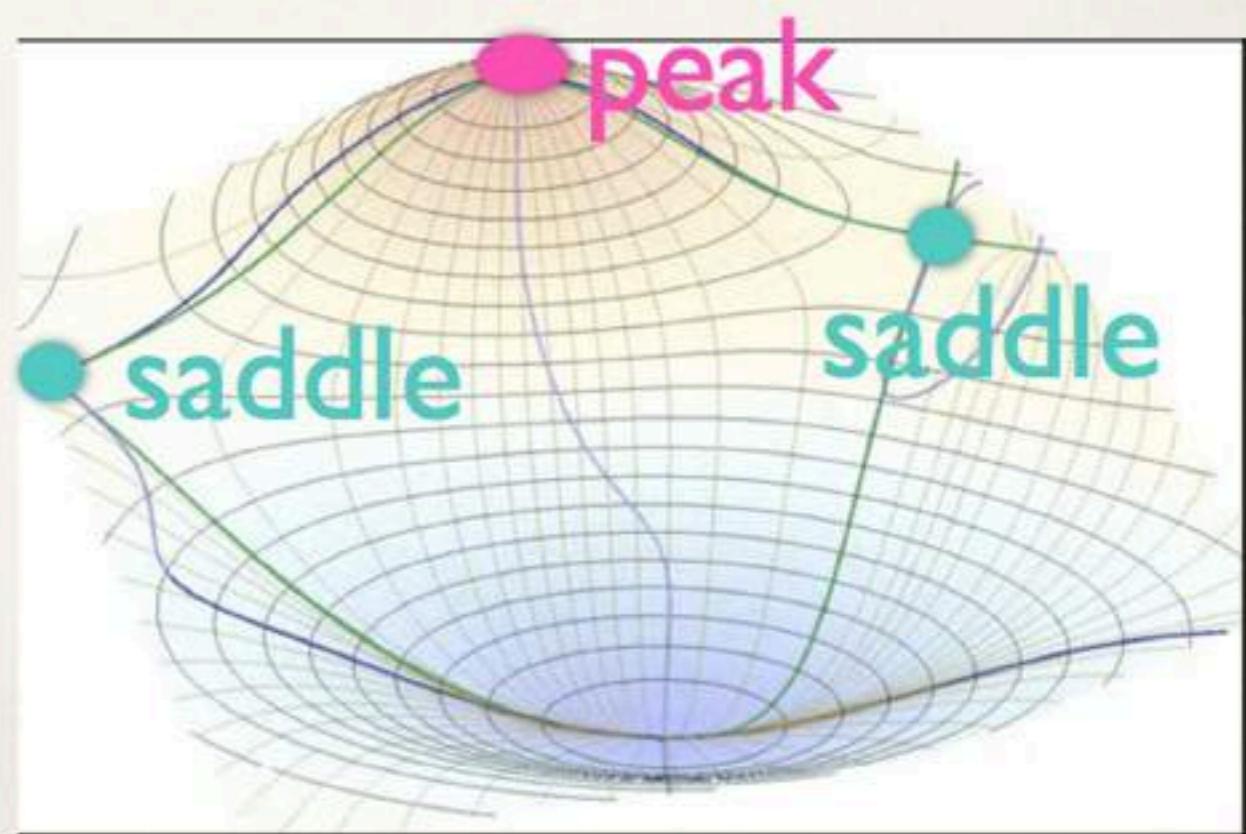
Can we understand where spin and vorticity alignments come from?

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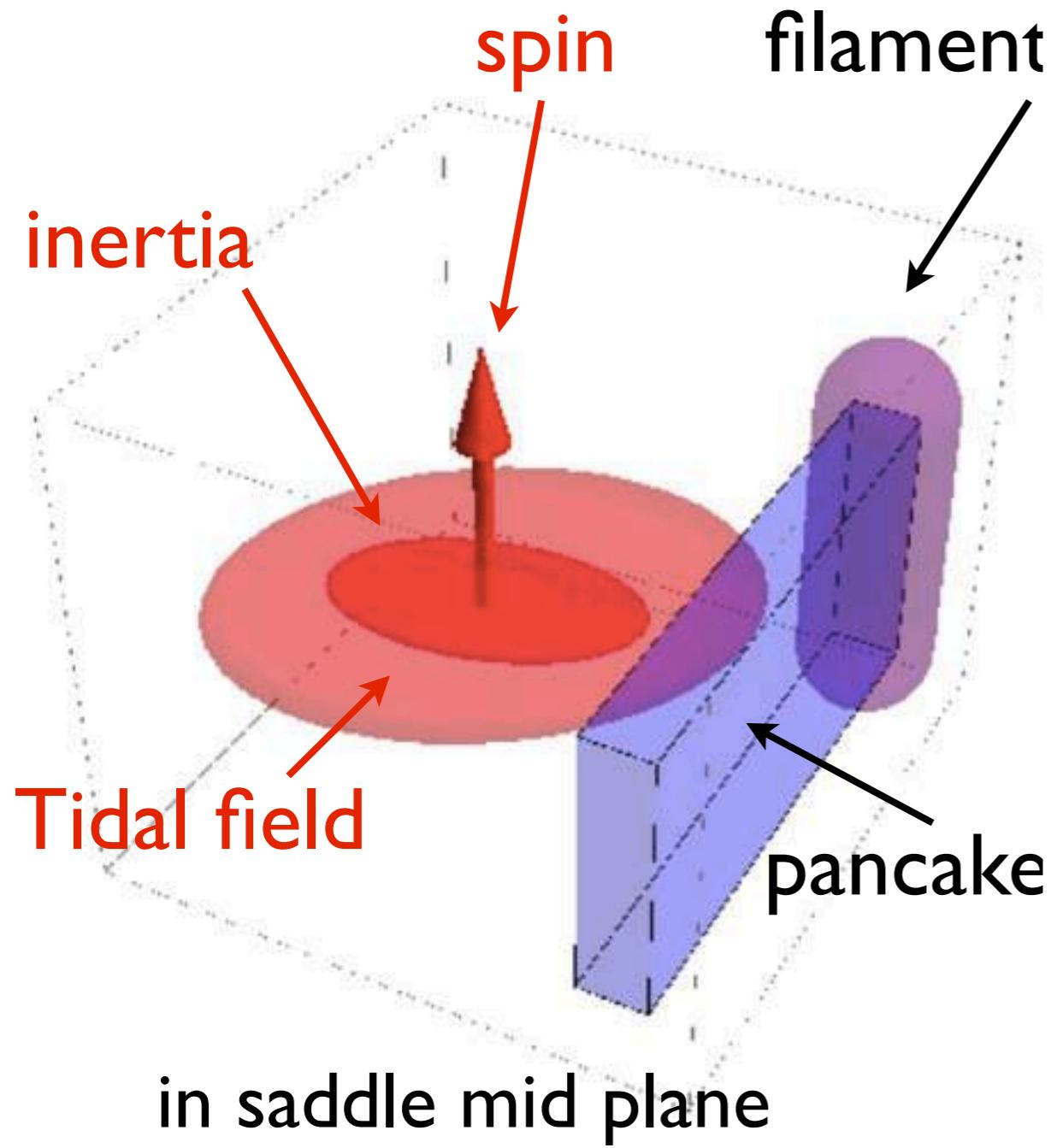
$$L_k = \varepsilon_{ijk} I_{li} T_{lj}$$



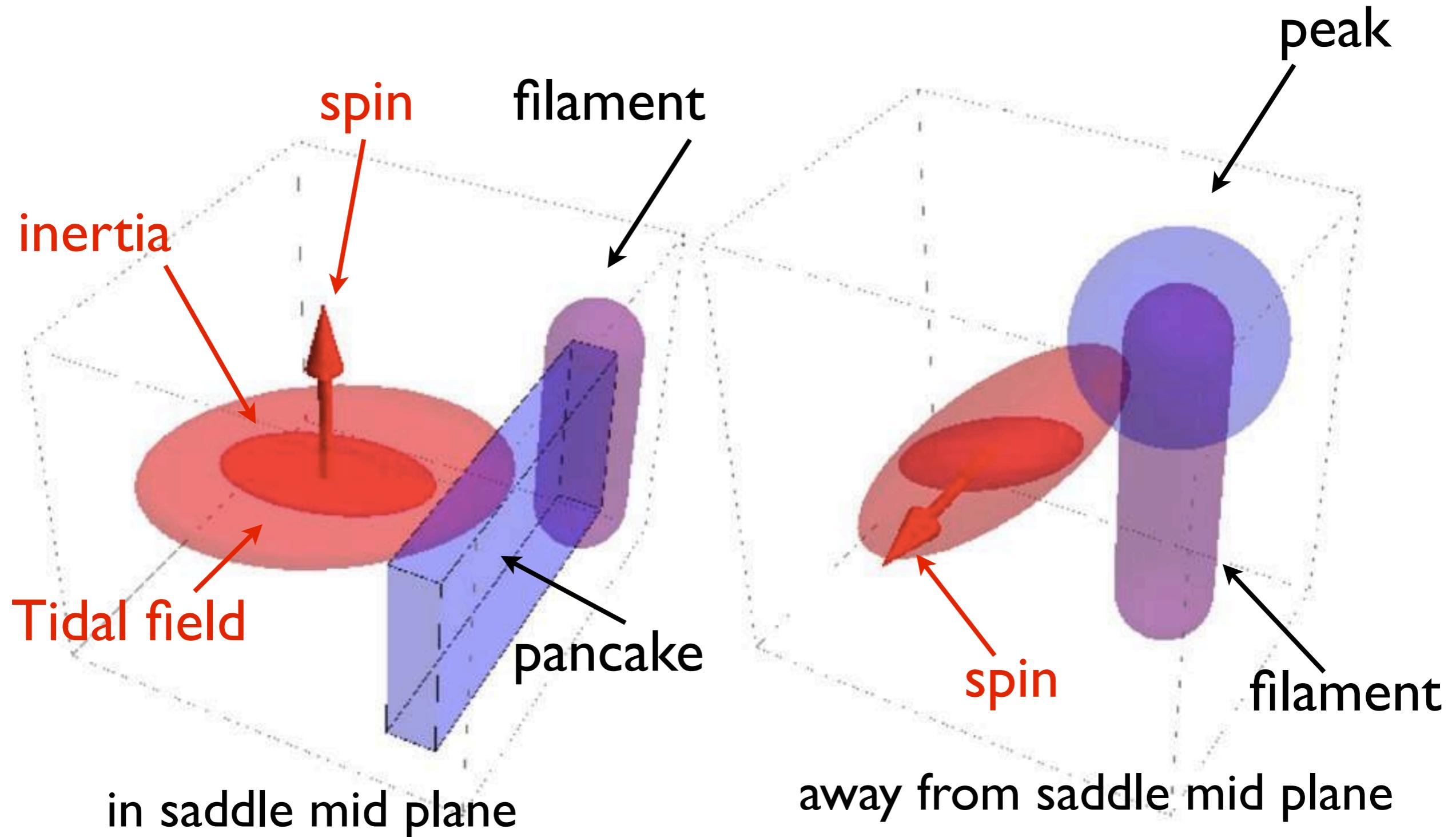
-anisotropy of the cosmic web:
surrounding of a saddle point
with typical geometry



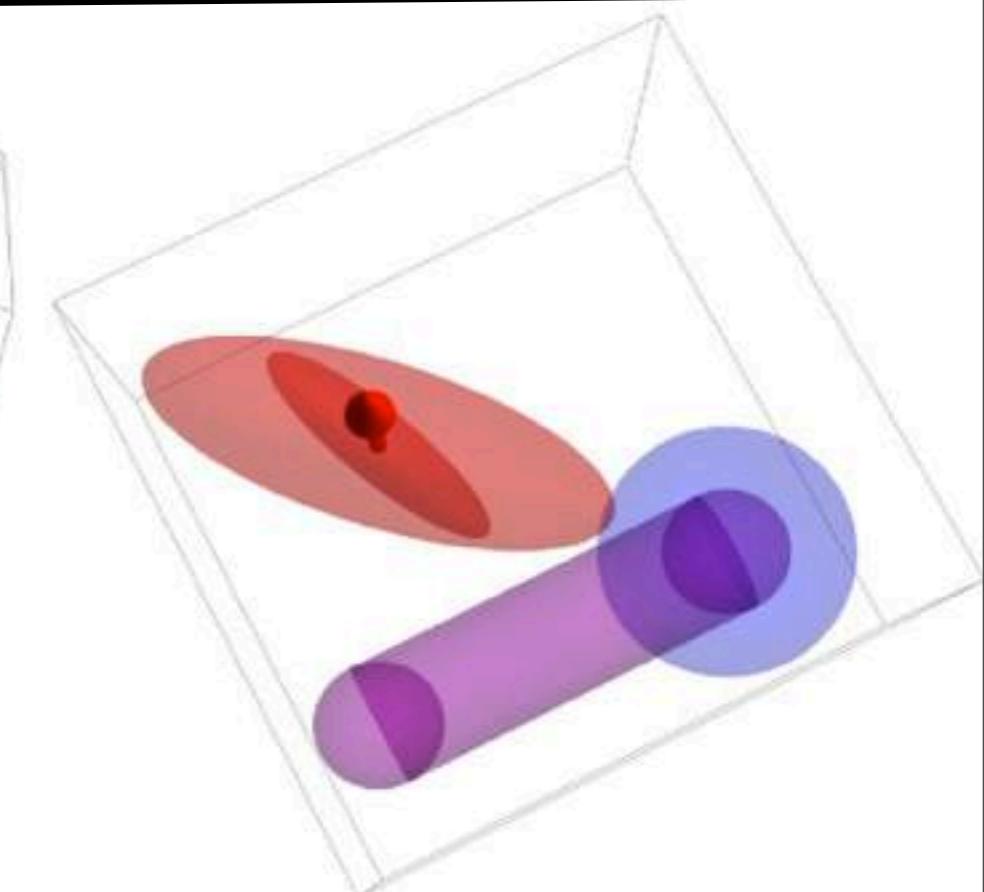
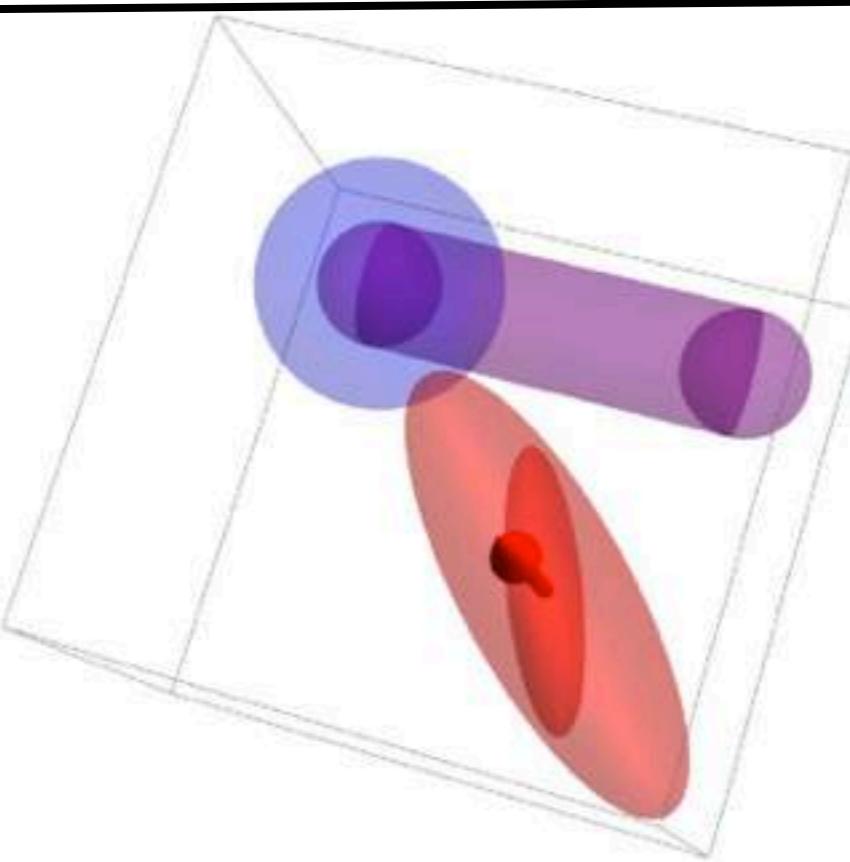
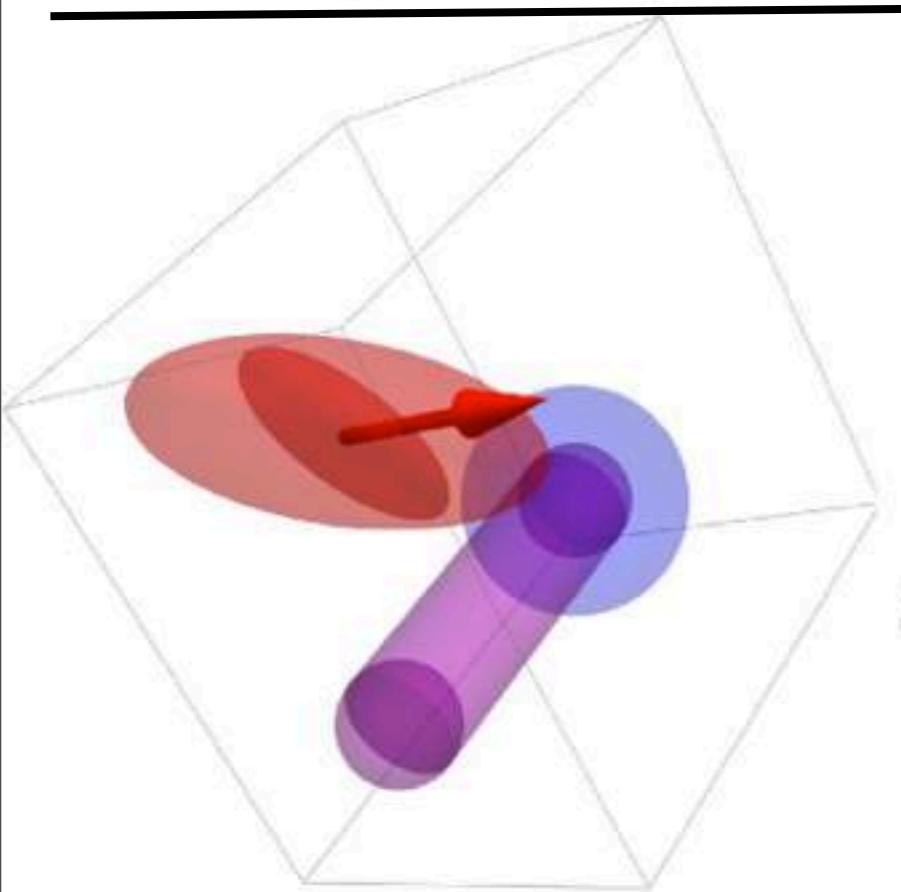
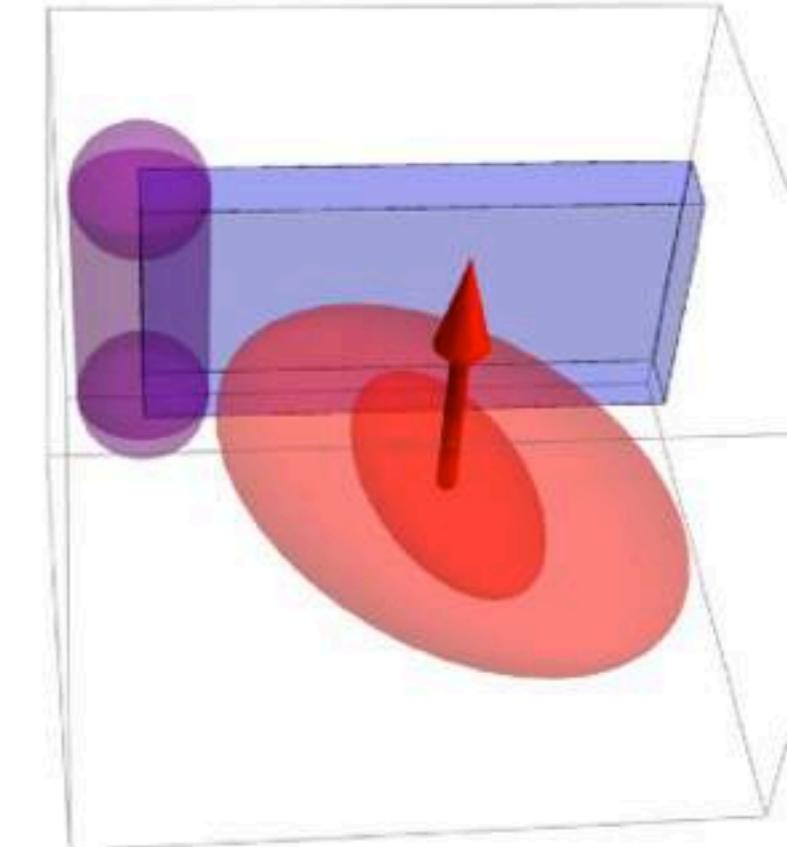
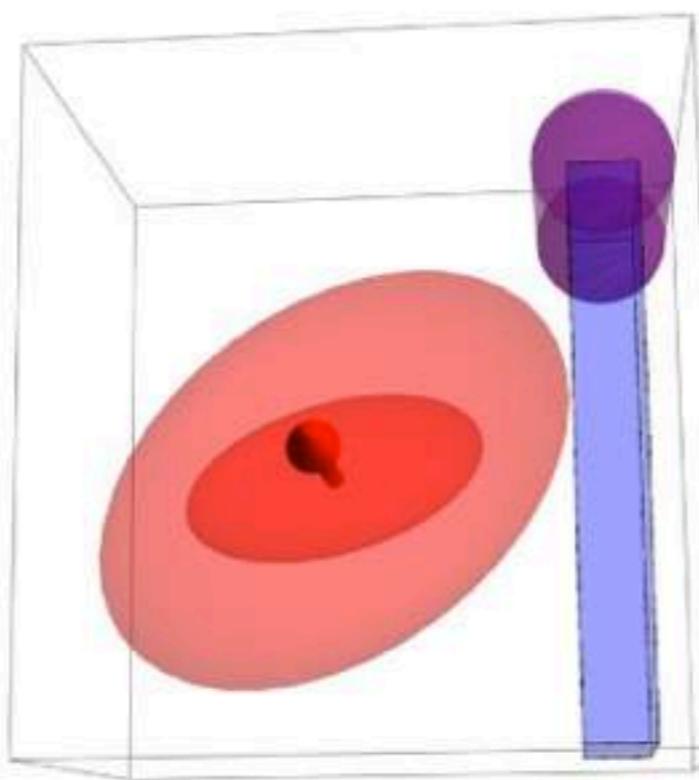
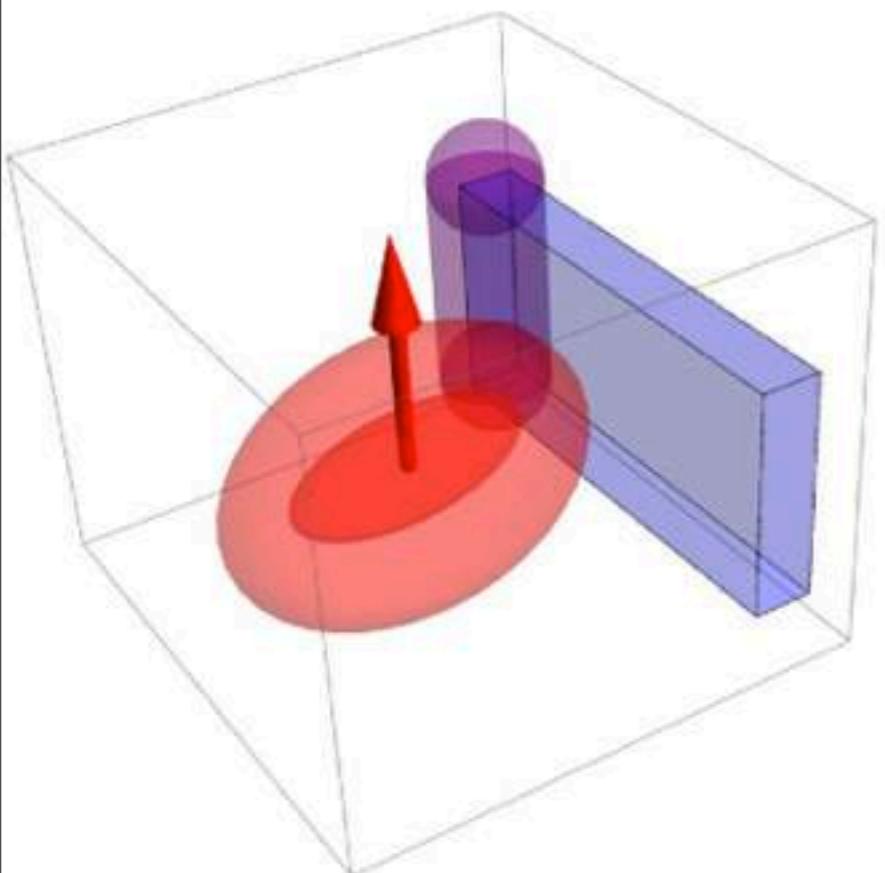
Tidal/Inertia mis-alignment



Tidal/Inertia mis-alignment



spin wall -filament



spin filament-cluster

animation?

Spin structure near Saddle

$$L_k = \varepsilon_{ijk} I_{li} T_{lj}$$

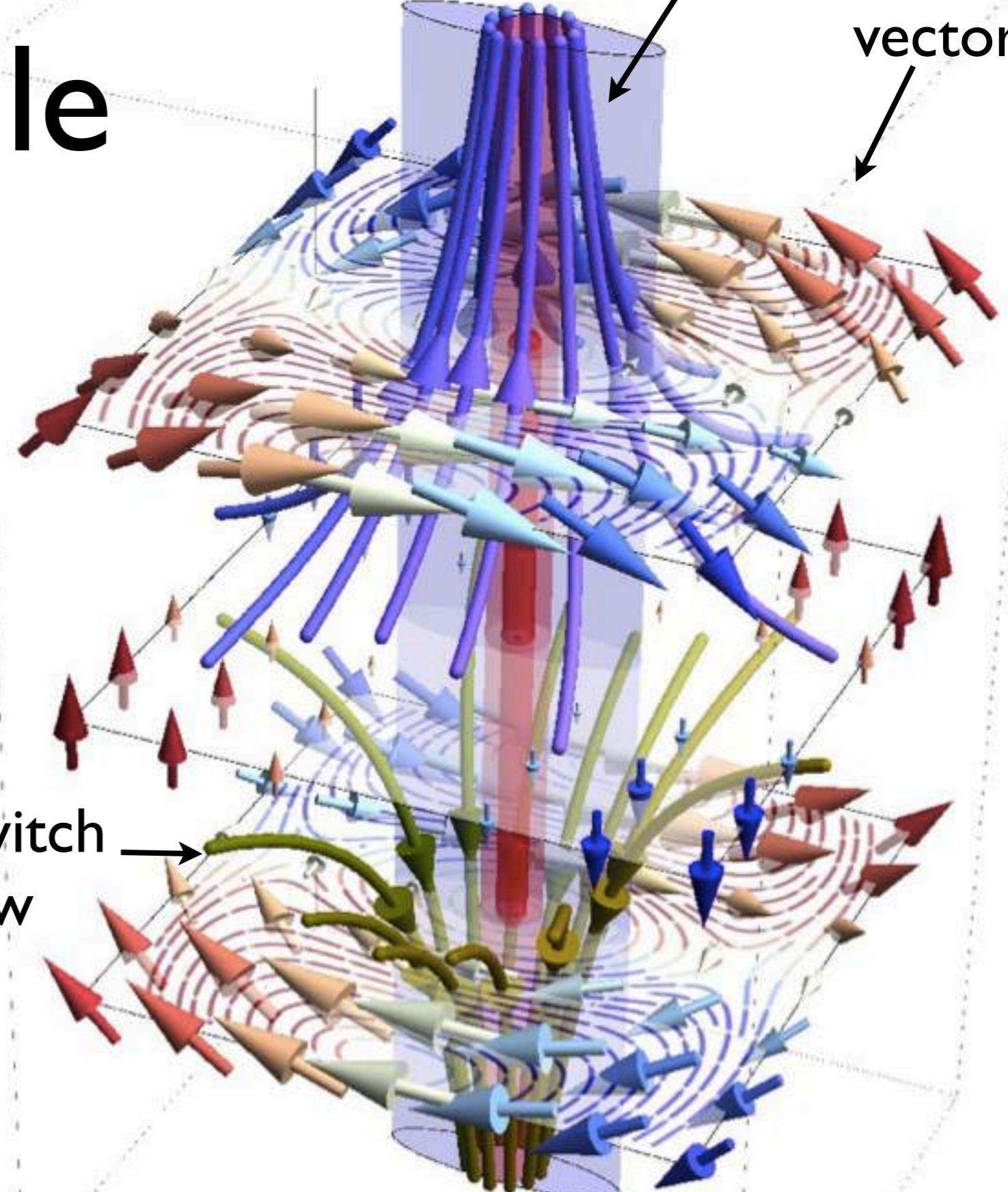
$$\approx \varepsilon_{ijk} H_{li} T_{lj}$$

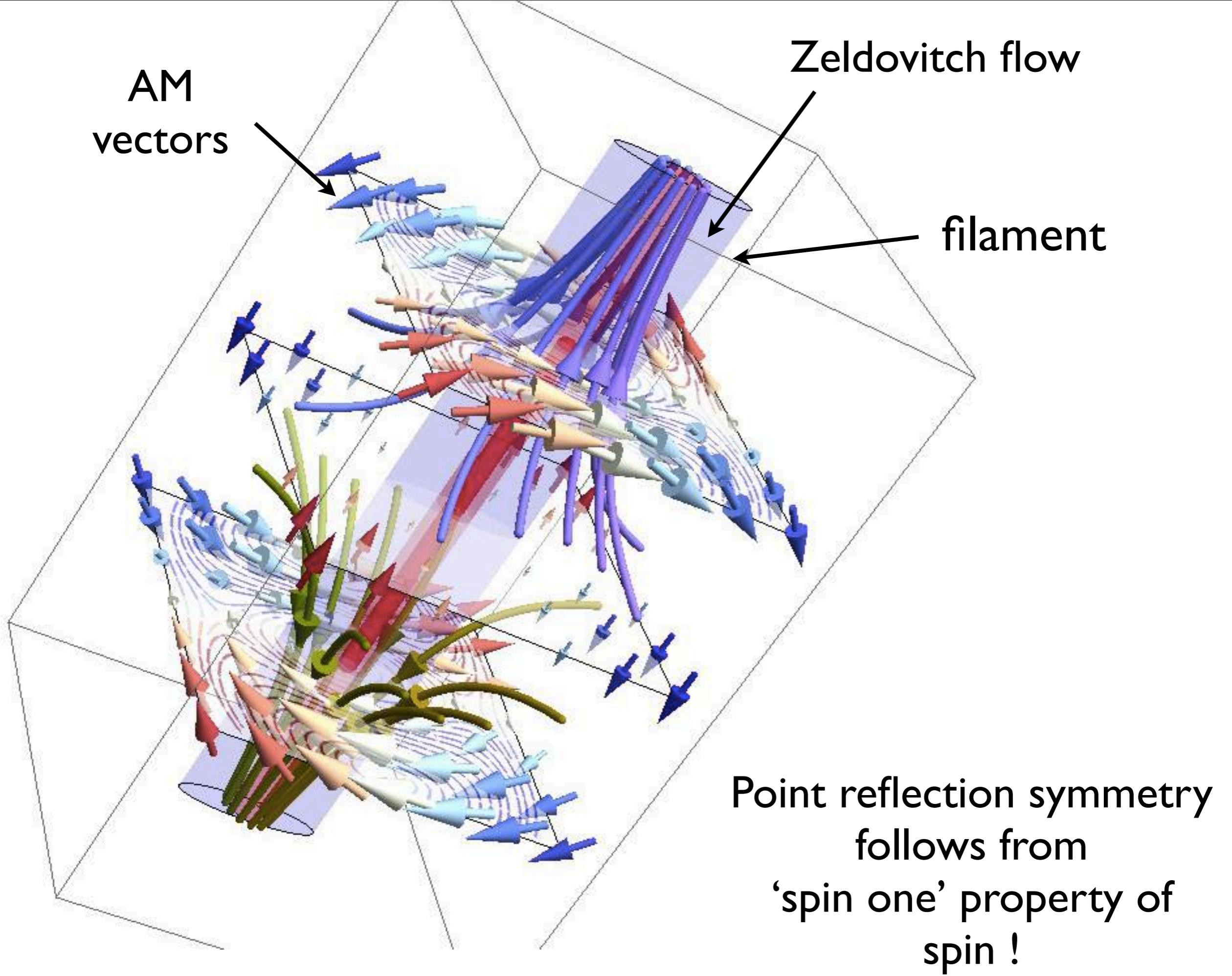
Hessian

Tidal

Zeldovitch
flow

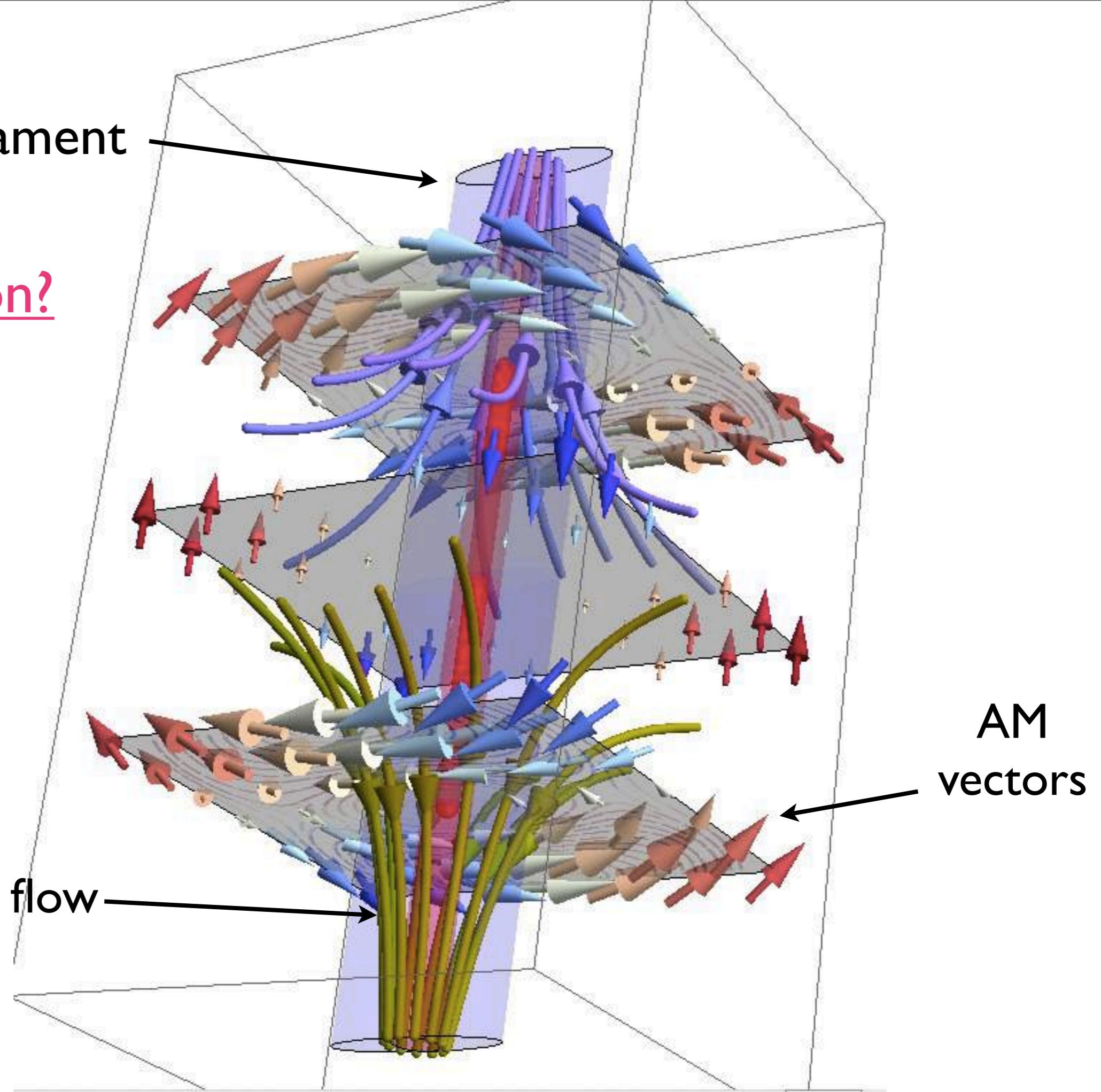
Flattened filament
AM
vectors





filament

animation?



3D TTT@ saddle?

- point reflection symmetric $\mathbf{r} \rightarrow -\mathbf{r}$
- vanish if no a-symmetry

perp. along e_φ



spin //
to filament

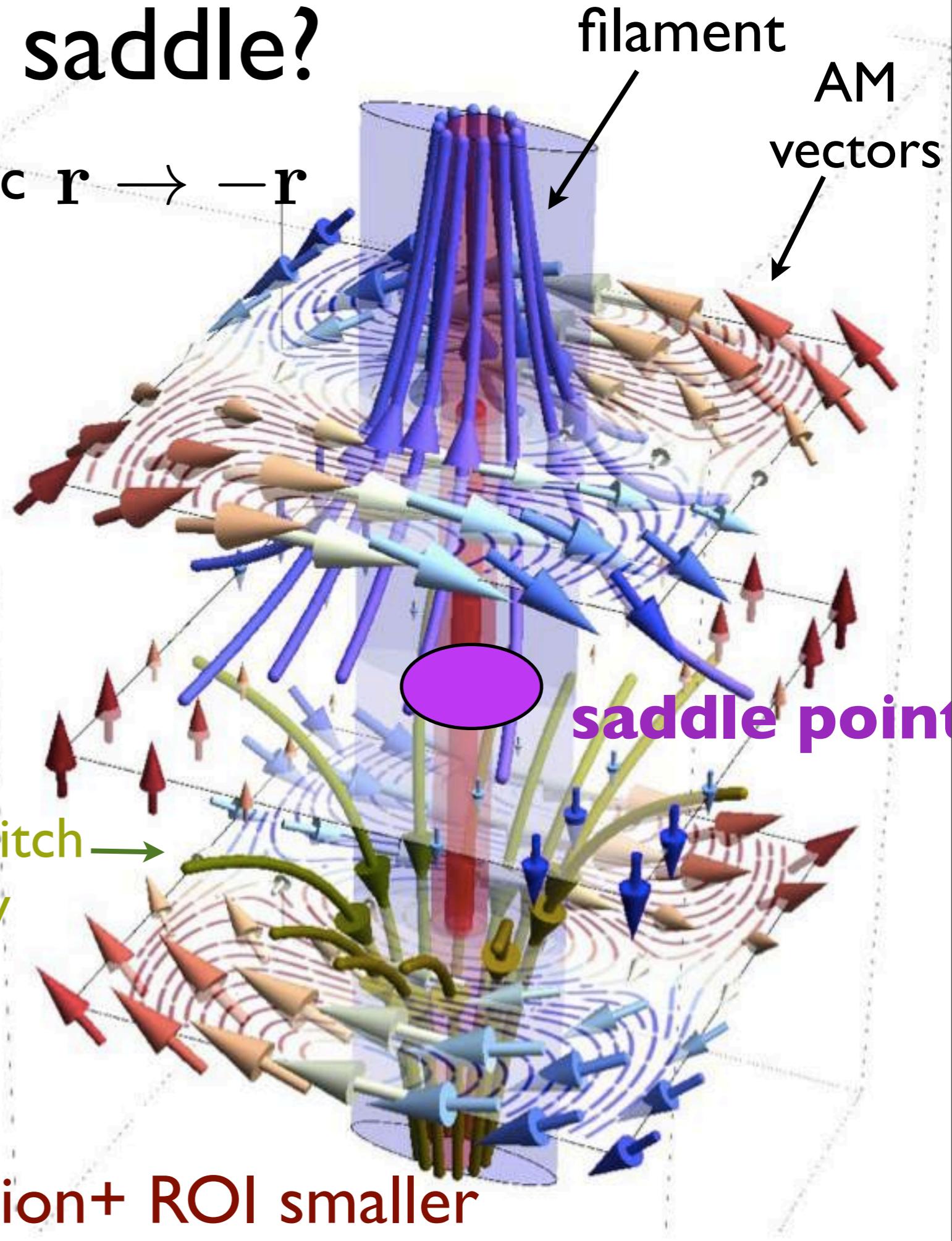


perp =
along e_φ

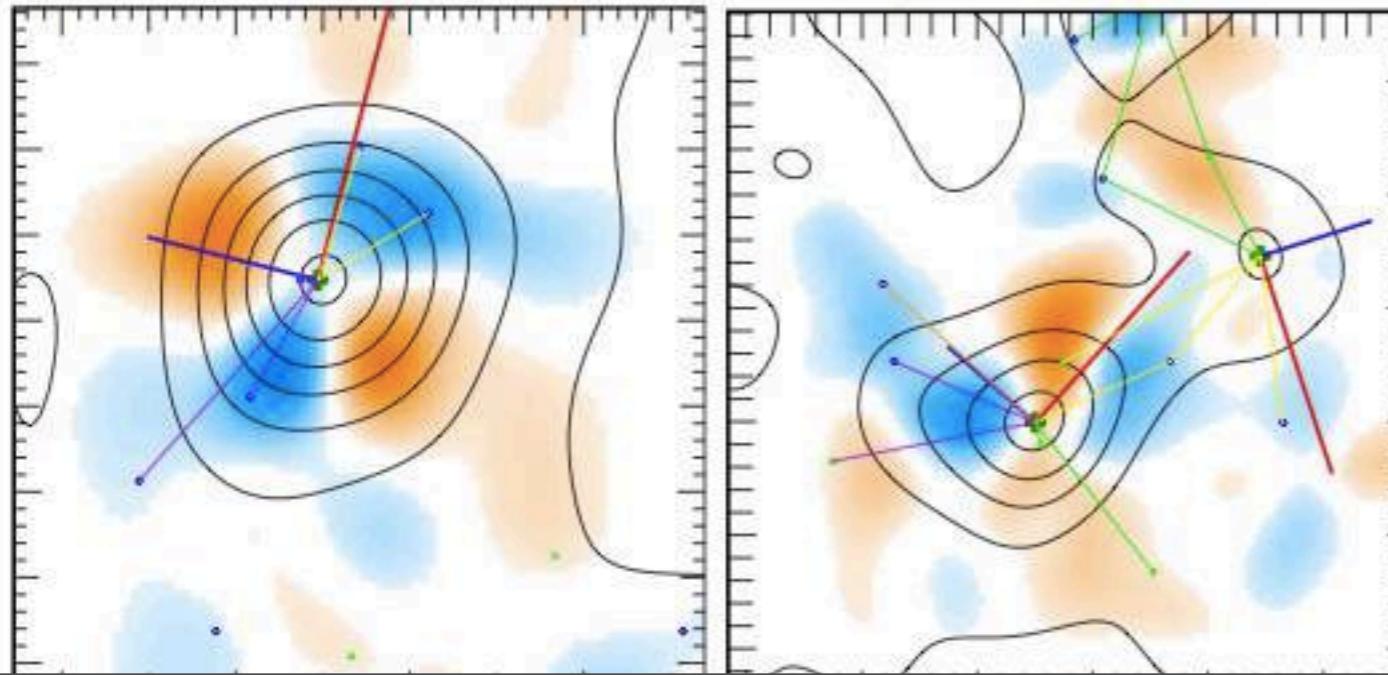
Zeldovitch
flow

saddle point

spatial transition+ ROI smaller



Does it work with log-Gaussian Random Fields?



point reflection symmetry
for realistic sets of saddles
from log GRF

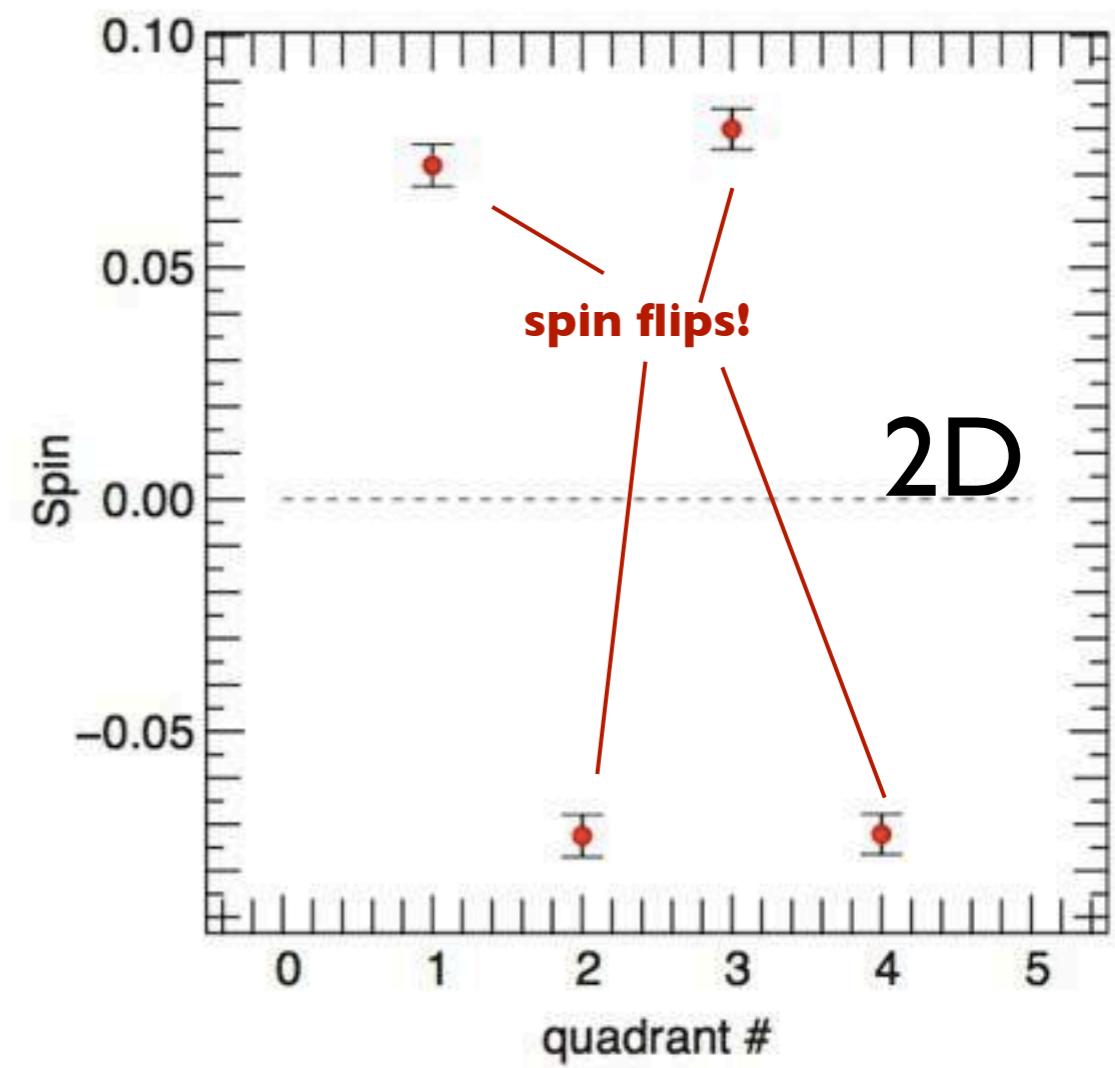
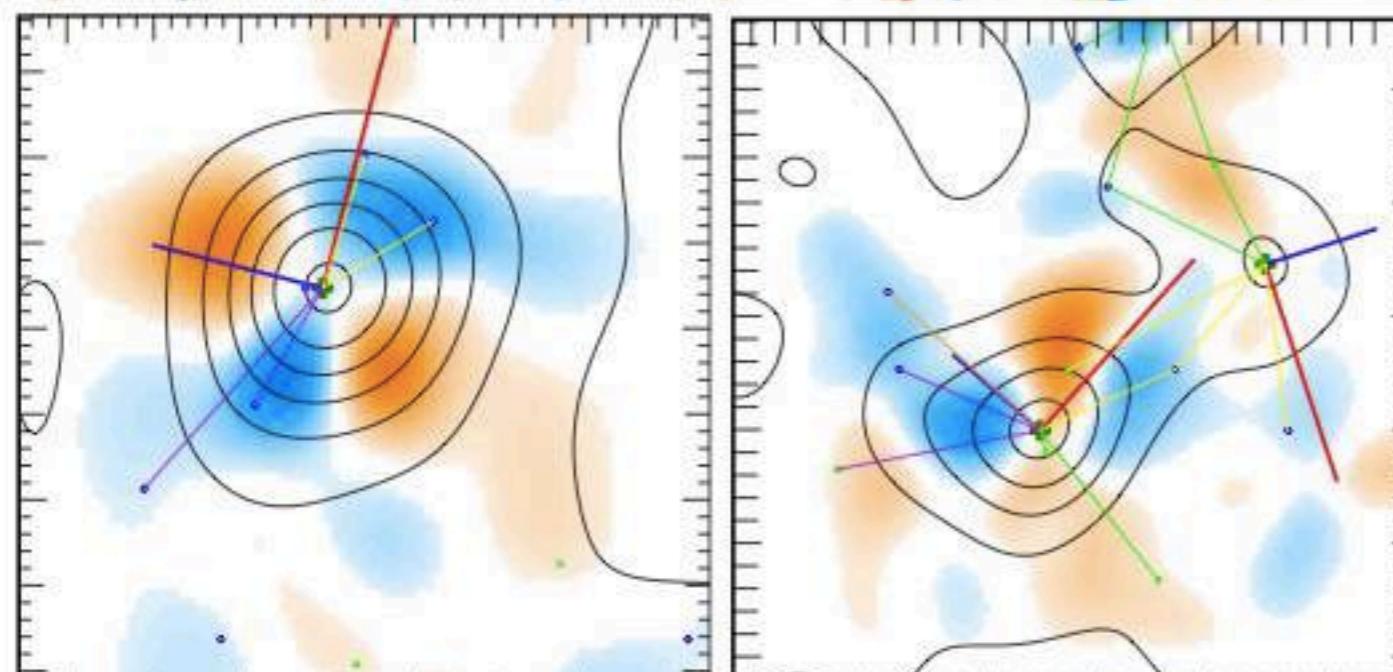
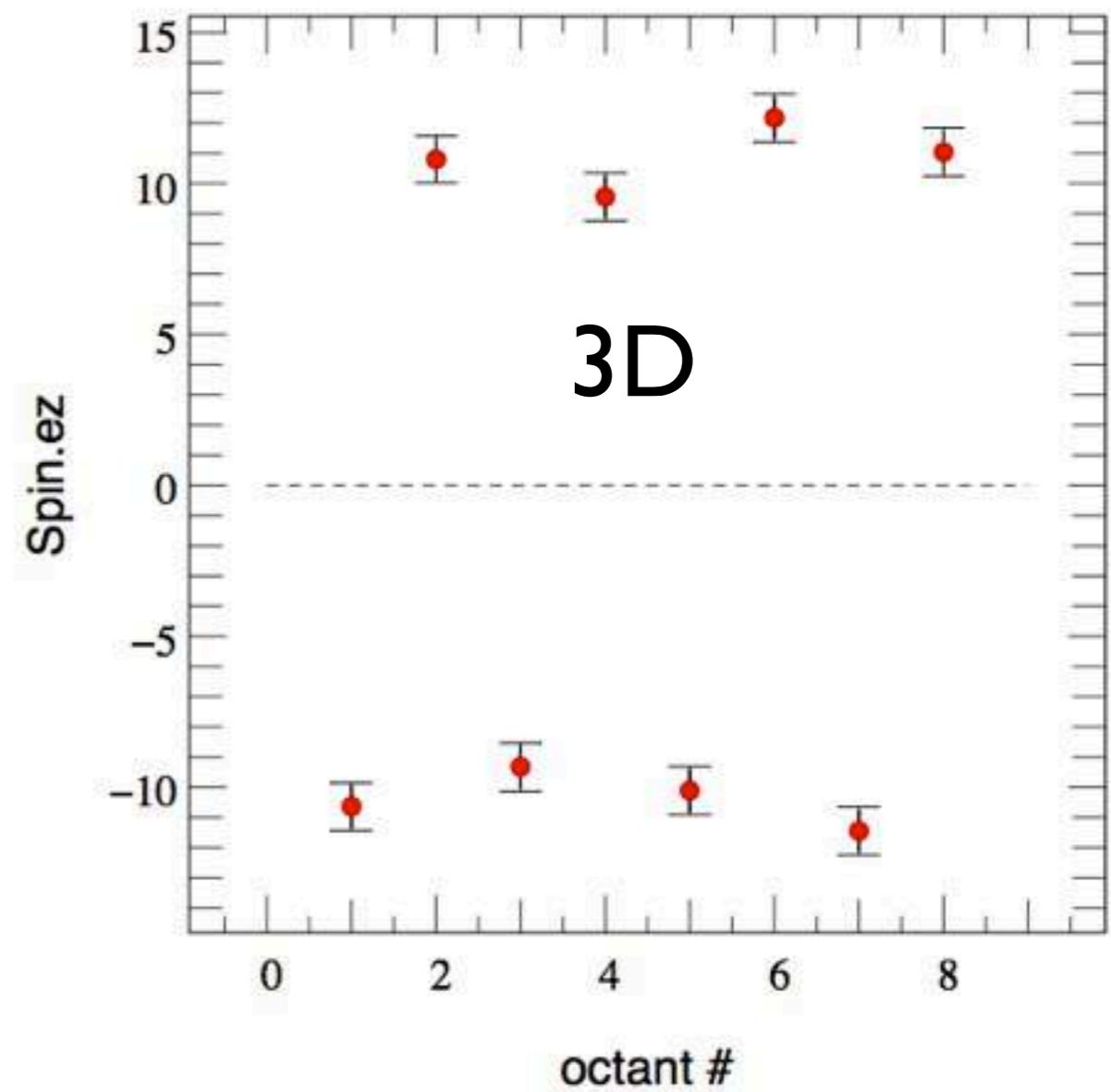


Figure 11. Alignment of ‘spin’ along e_z in 2D as a function of quadrant rank, clockwise. As expected, from one quadrant to the next, the spin is flipping sign.

Does it work with log-Gaussian Random Fields?



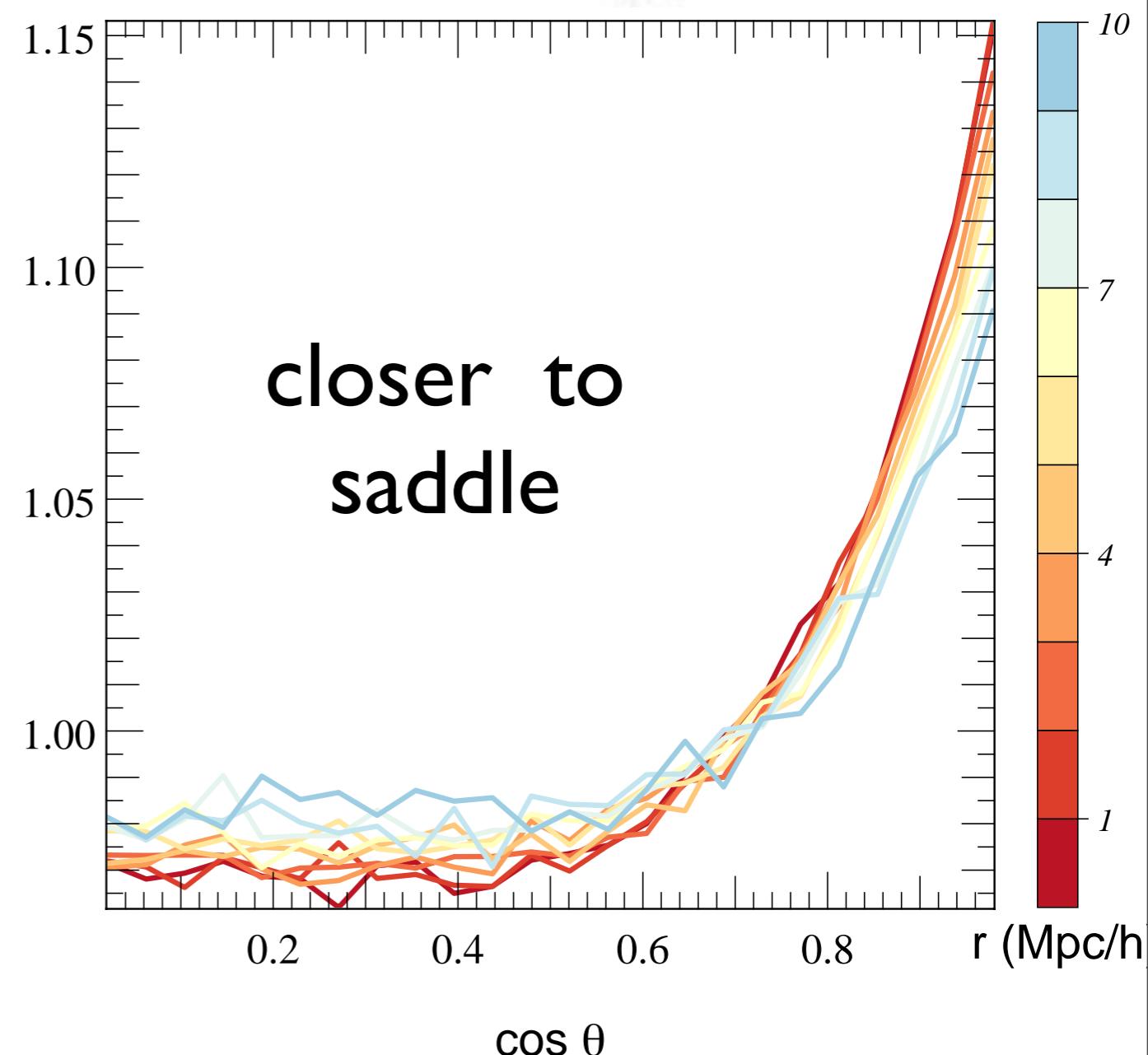
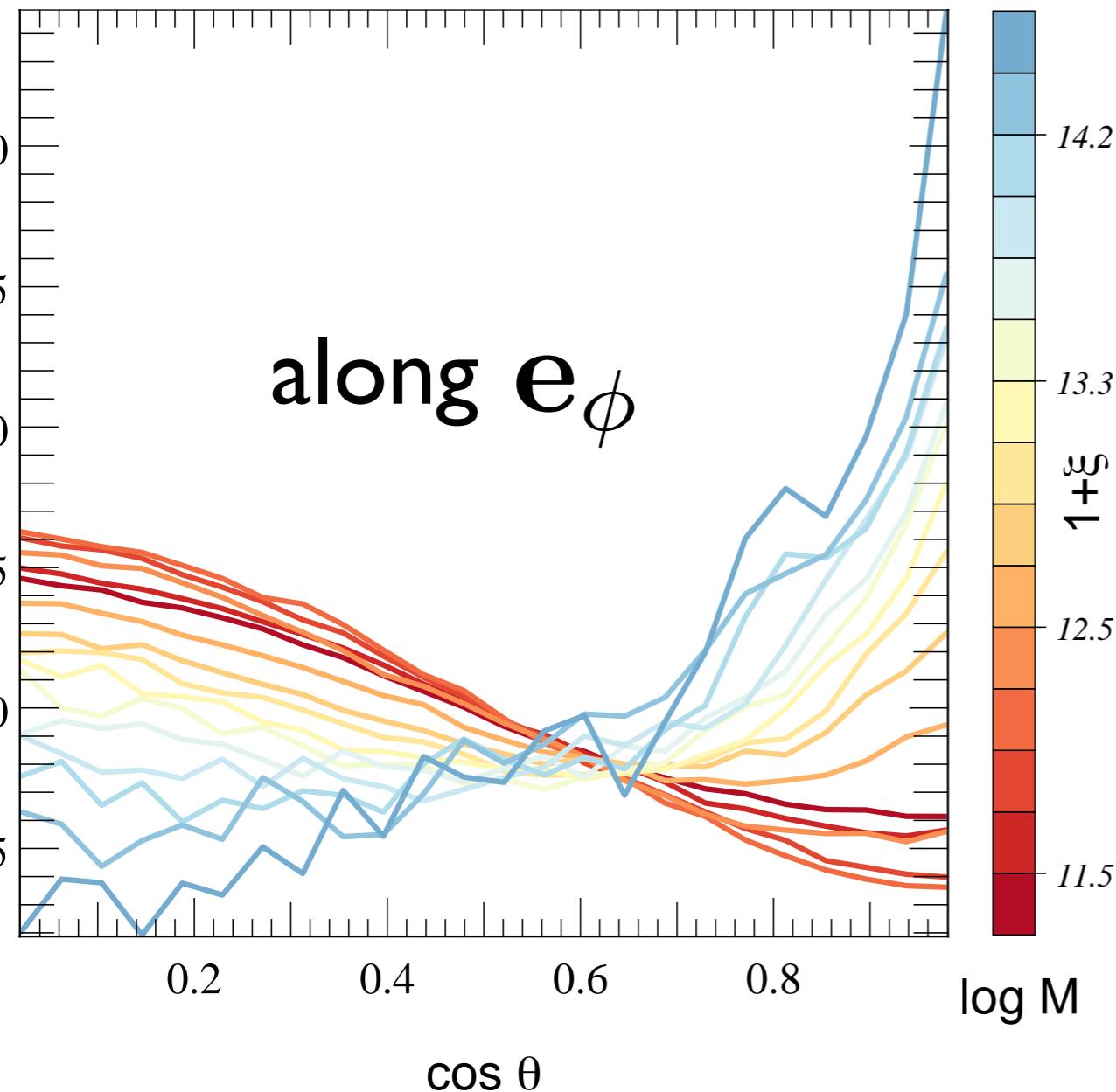
point reflection symmetry
for realistic sets of saddles
from log GRF



Does it work with Dark matter @ z=0?



Clear predictions of aTTT



2D Spin acquisition near peaks

$$L_k = \varepsilon_{ijk} I_{li} T_{lj}$$

Tidal

$$\approx \varepsilon_{ijk} H_{li} T_{lj}$$

Hessian

$\langle L | \text{peak} \rangle_{\text{2D}}?$

Theory will involve 2pt correlation of field AND 2nd derivatives

TTT@ saddle?

the Gaussain joint PDF of the derivatives of the field, $\mathbf{X} = \{x_{ij}, x_{ijk}, x_{ijkl}\}$ and $\mathbf{Y} = \{y_{ij}, y_{ijk}, y_{ijkl}\}$ in two given locations (\mathbf{r}_x and \mathbf{r}_y separated by a distance $r = |\mathbf{r}_x - \mathbf{r}_y|$) obeys

$$\text{PDF}(\mathbf{X}, \mathbf{Y}) = \frac{1}{\det[2\pi\mathbf{C}]^{1/2}} \times$$

$$\exp\left(-\frac{1}{2}\begin{bmatrix} \mathbf{X} \\ \mathbf{Y} \end{bmatrix}^T \cdot \begin{bmatrix} \mathbf{C}_0 & \mathbf{C}_\gamma \\ \mathbf{C}_\gamma^T & \mathbf{C}_0 \end{bmatrix}^{-1} \cdot \begin{bmatrix} \mathbf{X} \\ \mathbf{Y} \end{bmatrix}\right), \quad (\text{A2})$$

subject to the "saddle" constraints (2D)

height $x_{0,2} + x_{2,0} = \nu, \quad x_{1,2} + x_{3,0} = 0, \quad x_{0,3} + x_{2,1} = 0,$ *zero gradient*

$$\kappa \cos(2\theta) = \frac{1}{2} (x_{4,0} - x_{0,4}), \quad \kappa \sin(2\theta) = -x_{1,3} - x_{3,1}.$$

parametrized curvature

TTT@ saddle?

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PDF(\mathbf{X})

$$\exp\left(-\frac{1}{2}\begin{bmatrix} \mathbf{X} \\ \mathbf{Y} \end{bmatrix}^T\right).$$

$$\begin{aligned}
 & x_{0,0,2} + x_{0,2,0} + x_{2,0,0} = \nu, \quad x_{1,0,2} + x_{1,2,0} + x_{3,0,0} = 0, & 3D \\
 & x_{0,1,2} + x_{0,3,0} + x_{2,1,0} = 0, \quad x_{0,0,3} + x_{0,2,1} + x_{2,0,1} = 0, \\
 & \kappa_{1,1} = \frac{1}{3}(x_{2,0,2} - x_{0,0,4} - 2x_{0,2,2} - x_{0,4,0} + x_{2,2,0} + 2x_{4,0,0}), \\
 & \kappa_{1,2} = x_{1,1,2} + x_{1,3,0} + x_{3,1,0}, \quad \kappa_{1,3} = x_{1,0,3} + x_{1,2,1} + x_{3,0,1}, \\
 & \kappa_{2,2} = \frac{1}{3}(x_{0,2,2} - x_{0,0,4} + 2x_{0,4,0} - 2x_{2,0,2} + x_{2,2,0} - x_{4,0,0}), \\
 & \kappa_{2,3} = x_{0,1,3} + x_{0,3,1} + x_{2,1,1}. & (B4)
 \end{aligned}$$

subject to the "saddle" constraints (2D)

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parametrized curvature

Define the spin at point \mathbf{r}_y along the z direction as the anti-symmetric contraction of the de-traced tidal field and hessian:

(2D)

$$L(\mathbf{r}_y) = \varepsilon_{ij} \bar{y}_{il} \bar{y}_{jmml} = (y_{2,0} - y_{0,2}) (y_{1,3} + y_{3,1}) + \frac{y_{1,1}}{2} (y_{0,4} - y_{4,0}) - \frac{y_{1,1}}{2} (y_{4,0} - y_{0,4}) . \quad (\text{A3})$$

It is then fairly straightforward to compute the corresponding constrained expectation, $\langle L | \text{pk} \rangle$, for L as

$$L_z(r, \theta, \kappa, \nu) = \int L(\mathbf{Y}) \text{PDF}(\mathbf{X}, \mathbf{Y} | \text{pk}) d\mathbf{X} d\mathbf{Y} . \quad (\text{A4})$$



e.g. for $n=-2$

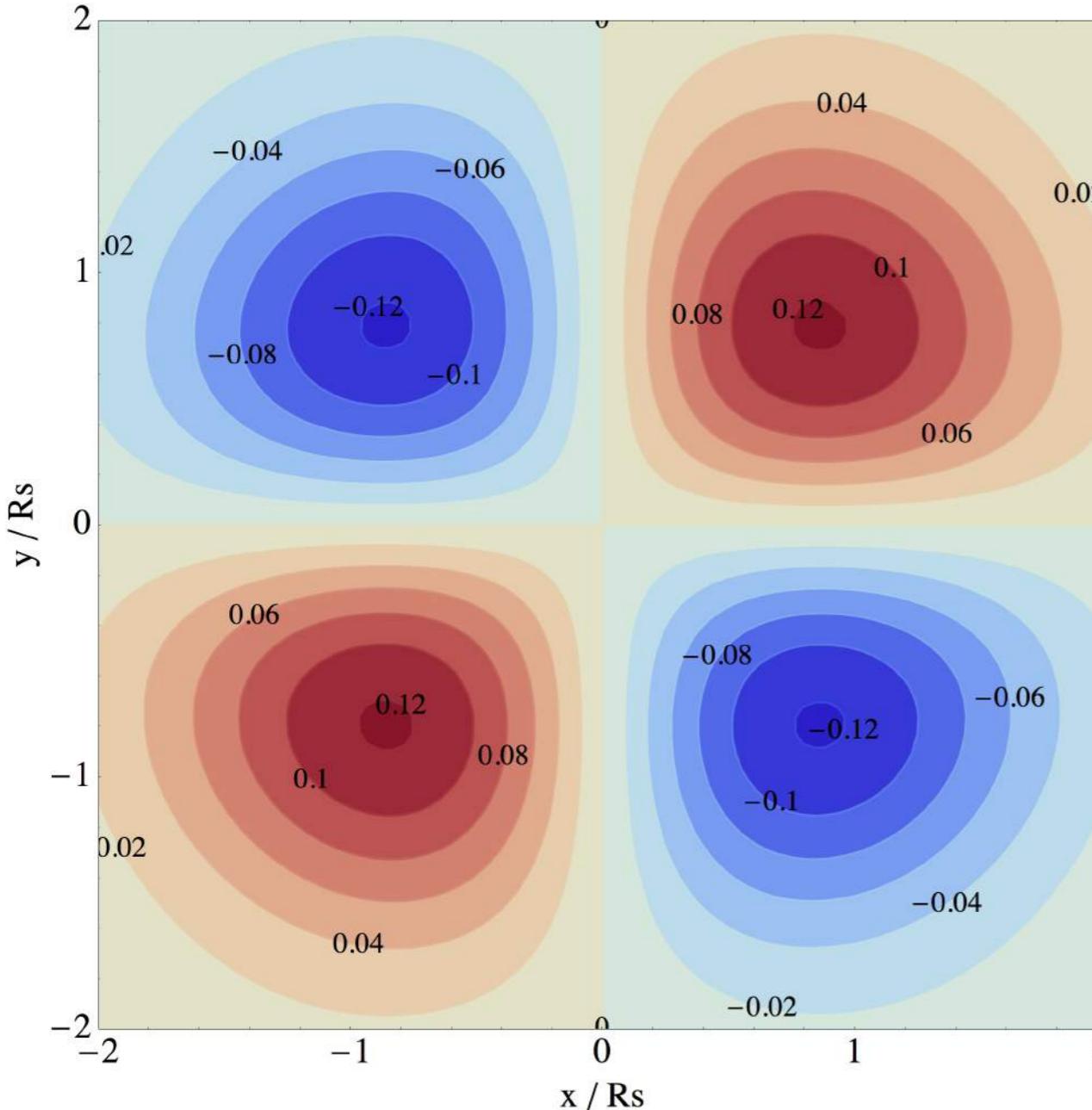
Incredibly simple prediction !

$$L_z = \kappa \frac{r^4 \sin(2\theta)}{144} e^{-\frac{r^2}{2}} \left(\sqrt{6} \kappa (r^2 - 4) \cos(2\theta) + 6 \right) .$$

asymmetry

finite extent

anti-symmetry



e.g. for $n=-2$

Incredibly simple prediction !

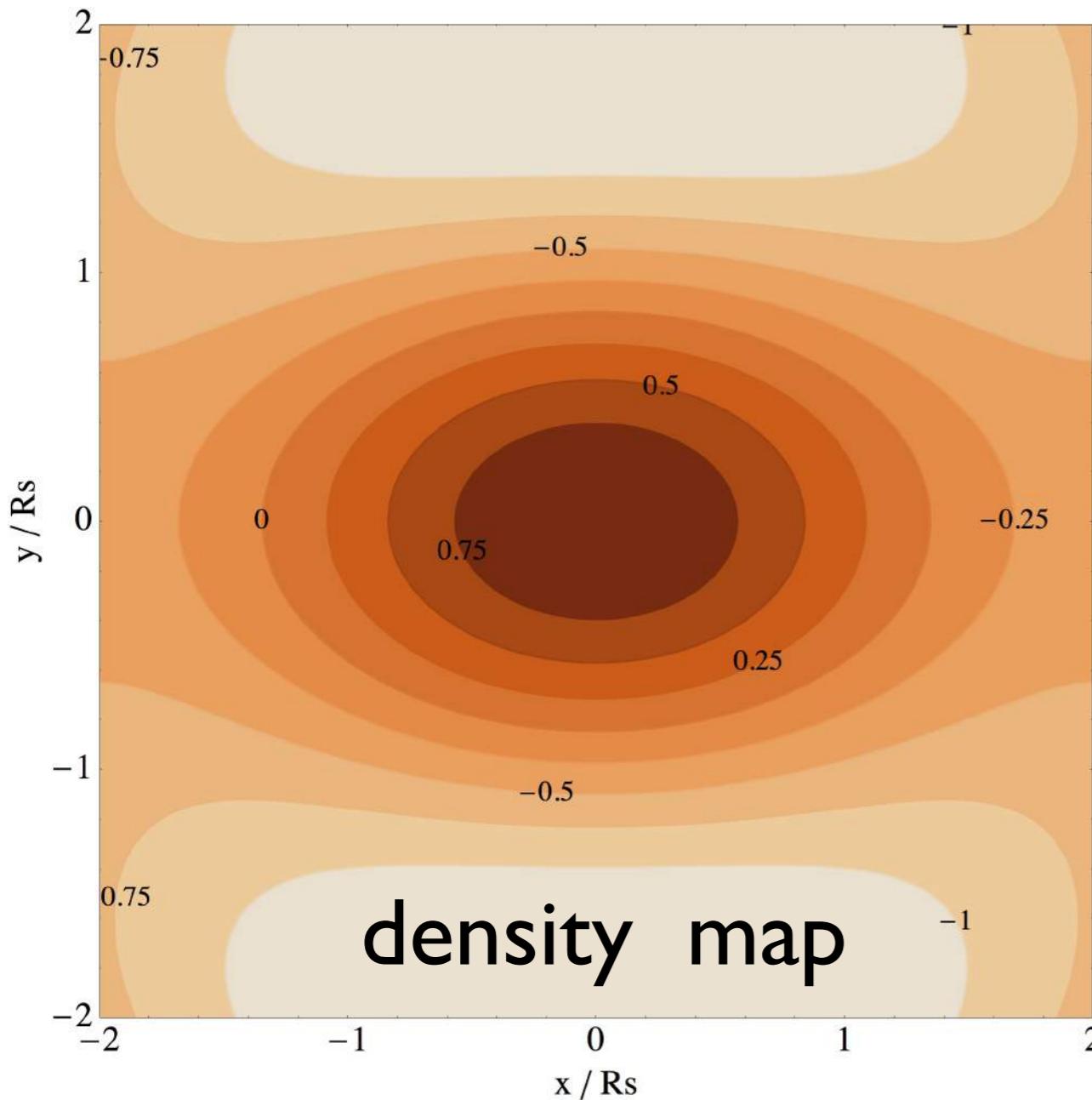
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asymmetry
finite extent
anti-symmetry

2D Theory of Tidal Torque @ saddle?

$$\delta(\mathbf{r}, \kappa, I_1, \nu | \text{ext}) = \frac{I_1(\xi_{\phi\delta}^{\Delta\Delta} + \gamma\xi_{\phi\phi}^{\Delta\Delta}) + \nu(\xi_{\phi\phi}^{\Delta\Delta} + \gamma\xi_{\phi\delta}^{\Delta\Delta})}{1 - \gamma^2} + 4 (\hat{\mathbf{r}}^T \cdot \bar{\mathbf{H}} \cdot \hat{\mathbf{r}}) \xi_{\phi\delta}^{\Delta+},$$

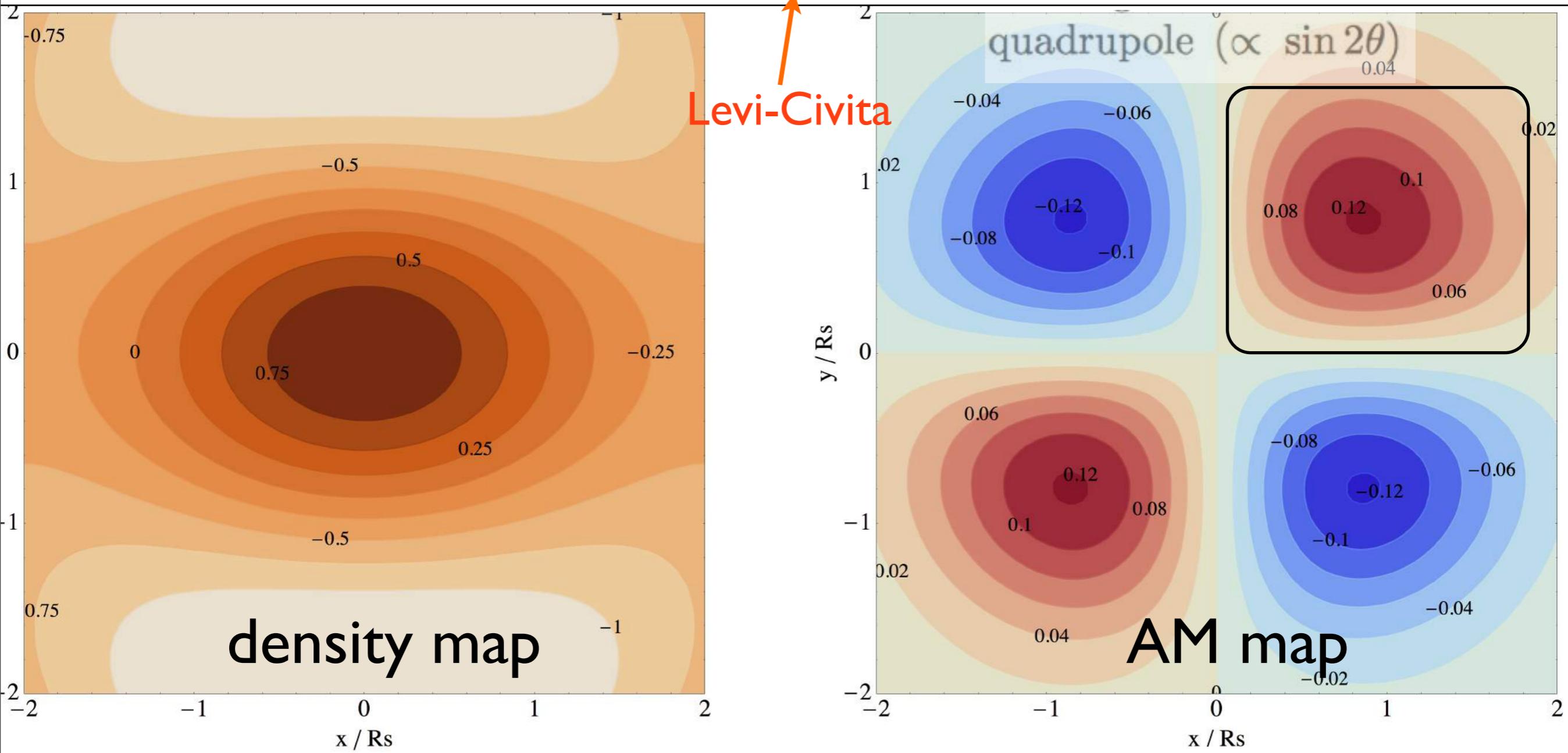
Hessian



$$f^+ = (f_{11} - f_{22})/2 \text{ and } f^\times = f_{12}.$$

2D Theory of Tidal Torque @ saddle?

$$\langle L_z | \text{ext} \rangle = L_z(\mathbf{r}, \kappa, I_1, \nu | \text{ext}) = -16(\hat{\mathbf{r}}^T \cdot \boldsymbol{\epsilon} \cdot \bar{\mathbf{H}} \cdot \hat{\mathbf{r}}) \left(L_z^{(1)}(r) + 2(\hat{\mathbf{r}}^T \cdot \bar{\mathbf{H}} \cdot \hat{\mathbf{r}}) L_z^{(2)}(r) \right)$$



$$L_z^{(1)}(r) = \frac{\nu}{1-\gamma^2} \left[(\xi_{\phi\phi}^{\Delta+} + \gamma \xi_{\phi\delta}^{\Delta+}) \xi_{\delta\delta}^{\times\times} - (\xi_{\phi\delta}^{\Delta+} + \gamma \xi_{\delta\delta}^{\Delta+}) \xi_{\phi\delta}^{\times\times} \right]$$

$$L_z^{(2)}(r) = (\xi_{\phi x}^{\Delta\Delta} \xi_{\delta\delta}^{\times\times} - \xi_{\phi\delta}^{\times\times} \xi_{\delta\delta}^{\Delta\Delta}) + \frac{I_1}{1-\gamma^2} \left[(\xi_{\phi\delta}^{\Delta+} + \gamma \xi_{\phi\phi}^{\Delta+}) \xi_{\delta\delta}^{\times\times} - (\xi_{\delta\delta}^{\Delta+} + \gamma \xi_{\phi\delta}^{\Delta+}) \xi_{\phi\delta}^{\times\times} \right]$$

In order to compute the spin distribution, the formalism developed in Section 2 is easily extended to 3D. A critical (including saddle condition) point constraint is imposed. The resulting mean density field subject to that constraint becomes (in units of σ_2):

$$\delta(\mathbf{r}, \kappa, I_1, \nu | \text{ext}) = \frac{I_1(\xi_{\phi\delta}^{\Delta\Delta} + \gamma\xi_{\phi\phi}^{\Delta\Delta})}{1 - \gamma^2} + \frac{\nu(\xi_{\phi\phi}^{\Delta\Delta} + \gamma\xi_{\phi\delta}^{\Delta\Delta})}{1 - \gamma^2} + \frac{15}{2} (\hat{\mathbf{r}}^T \cdot \bar{\mathbf{H}} \cdot \hat{\mathbf{r}}) \xi_{\phi\delta}^{\Delta+}, \quad (3.1)$$

where again $\bar{\mathbf{H}}$ is the *detraced* Hessian of the density and $\hat{\mathbf{r}} = \mathbf{r}/r$ and we define in 3D $\xi_{\phi x}^{\Delta+}$ as $\xi_{\phi\delta}^{\Delta+} = \langle \Delta\delta, \phi^+ \rangle$, with $\phi^+ = \phi_{11} - (\phi_{22} + \phi_{33})/2$. Note that $\hat{\mathbf{r}}^T \cdot \bar{\mathbf{H}} \cdot \hat{\mathbf{r}}$ is a scalar quantity defined explicitly as $\hat{r}_i \bar{H}_{ij} \hat{r}_j$. As in 2D, the expected spin can also be computed. In 3D, the spin is a vector, which components are given by $L_i = \epsilon_{ijk}\delta_{kl}\phi_{lj}$, with ϵ the rank 3 Levi Civita tensor. It is found to be orthogonal to the separation and can be written as the sum of two terms

$$\mathbf{L}(\mathbf{r}, \kappa, I_1, \nu | \text{ext}) = -15 \left(\mathbf{L}^{(1)}(r) + \mathbf{L}^{(2)}(\mathbf{r}) \right) \cdot (\hat{\mathbf{r}}^T \cdot \epsilon \cdot \bar{\mathbf{H}} \cdot \hat{\mathbf{r}}), \quad (3.2)$$

where $\mathbf{L}^{(1)}$ depends on height, ν , and on the trace of the Hessian I_1 but not on orientation

$$\begin{aligned} \mathbf{L}^{(1)}(r) = & \left(\frac{\nu}{1 - \gamma^2} \left[(\xi_{\phi\phi}^{\Delta+} + \gamma\xi_{\phi\delta}^{\Delta+})\xi_{\delta\delta}^{\times\times} - (\xi_{\phi\delta}^{\Delta+} + \gamma\xi_{\delta\delta}^{\Delta+})\xi_{\phi\delta}^{\times\times} \right] \right. \\ & \left. + \frac{I_1}{1 - \gamma^2} \left[(\xi_{\phi\delta}^{\Delta+} + \gamma\xi_{\phi\phi}^{\Delta+})\xi_{\delta\delta}^{\times\times} - (\xi_{\delta\delta}^{\Delta+} + \gamma\xi_{\phi\delta}^{\Delta+})\xi_{\phi\delta}^{\times\times} \right] \right) \mathbb{I}_3, \end{aligned}$$

and $\mathbf{L}^{(2)}(\mathbf{r})$ now depends on $\bar{\mathbf{H}}$ and on orientation:

$$\begin{aligned} \mathbf{L}^{(2)}(\mathbf{r}) = & -\frac{5}{8} \left[2((\xi_{\phi\delta}^{\Delta+} - \xi_{\phi\delta}^{\Delta\Delta})\xi_{\delta\delta}^{\times\times} - (\xi_{\delta\delta}^{\Delta+} - \xi_{\delta\delta}^{\Delta\Delta})\xi_{\phi\delta}^{\times\times}) \bar{\mathbf{H}} \right. \\ & \left. + ((7\xi_{\delta\delta}^{\Delta\Delta} + 5\xi_{\delta\delta}^{\Delta+})\xi_{\phi\delta}^{\times\times} - (7\xi_{\phi\delta}^{\Delta\Delta} + 5\xi_{\phi\delta}^{\Delta+})\xi_{\delta\delta}^{\times\times})(\hat{\mathbf{r}}^T \cdot \bar{\mathbf{H}} \cdot \hat{\mathbf{r}}) \mathbb{I}_3 \right], \end{aligned}$$

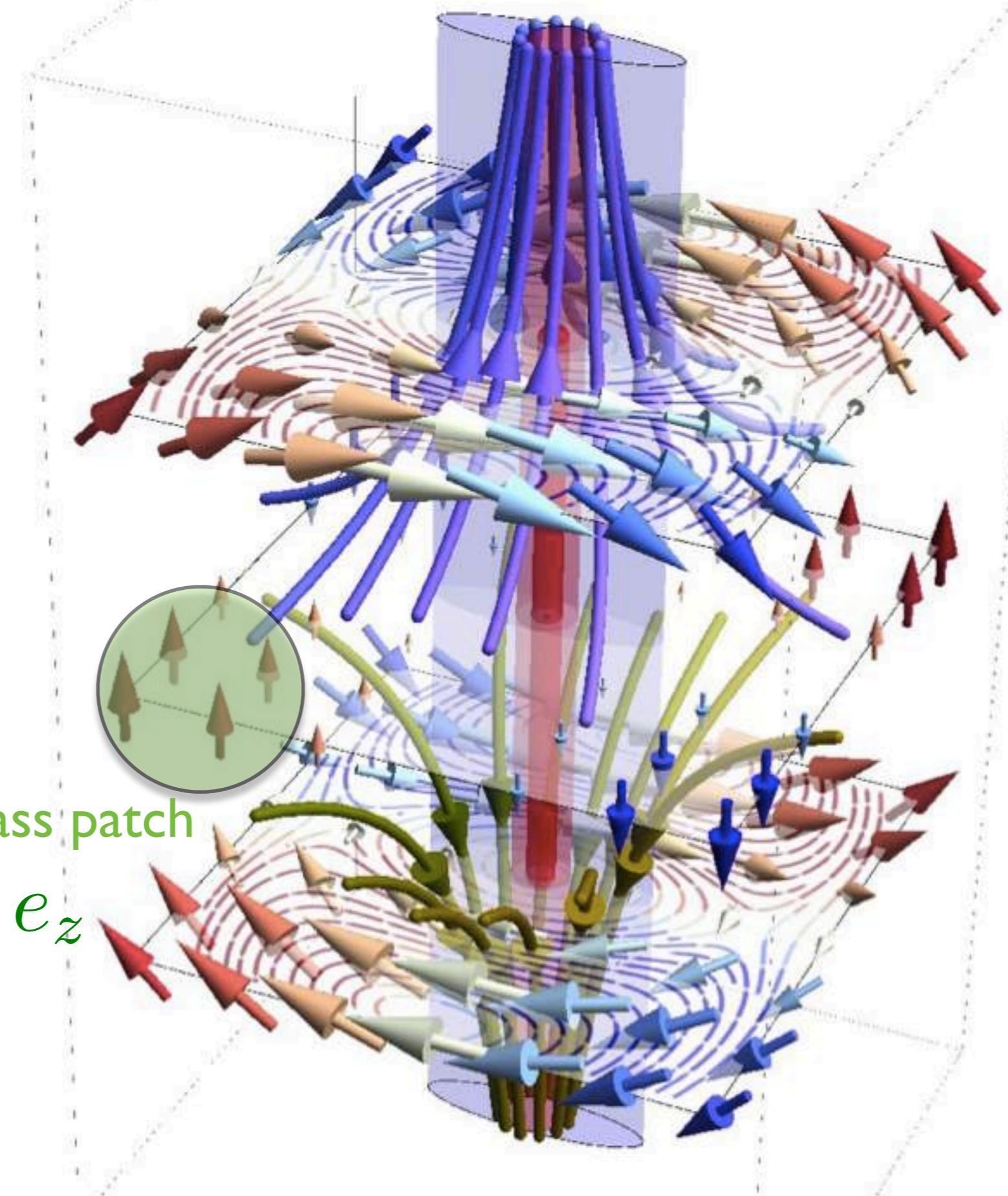
3D Transition mass ?

Lagrangian theory
capture spin flip !

Transition mass
associated
with **size**
of quadrant

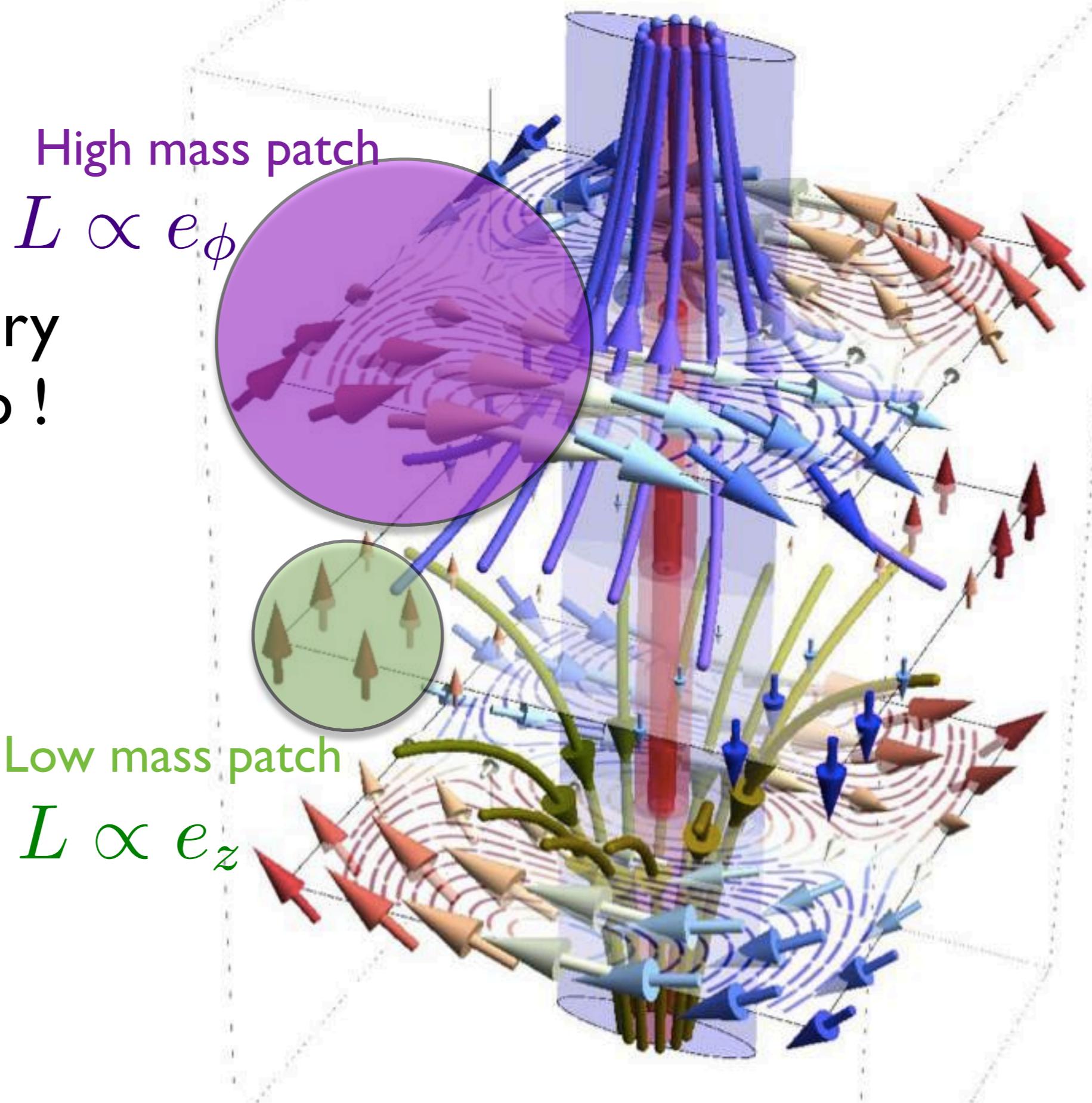
Low mass patch

$$L \propto e_z$$



3D Transition mass ?

Transition mass
associated
with **size**
of quadrant



Geometry of the saddle provides a natural ‘metric’ (local frame as defined by Hessian @ saddle) relative to which **dynamical evolution** of DH is predicted.

Cloud in
cloud effect

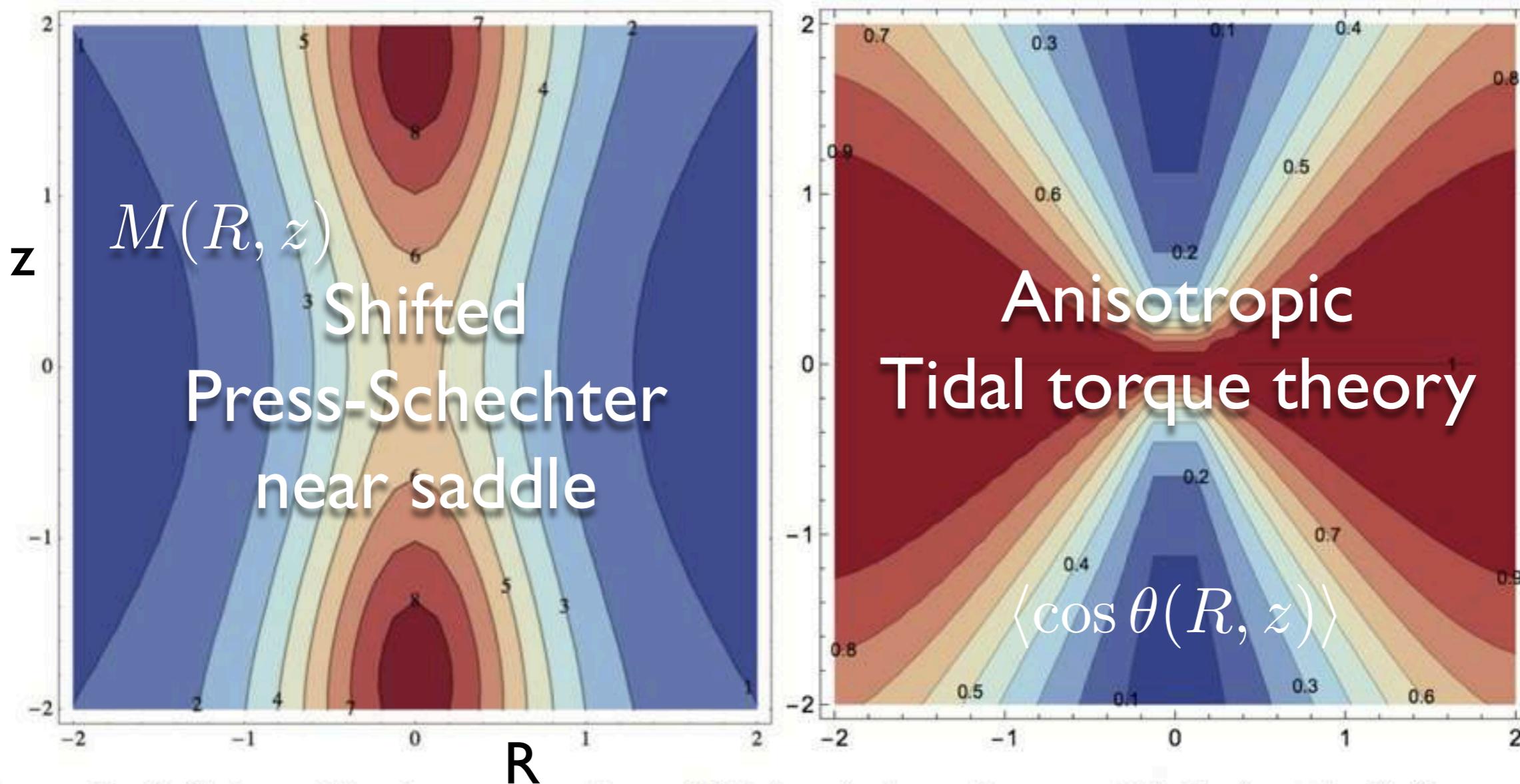


Figure 5. Left: logarithmic cross section of $M_p(r, z)$ along the most likely (vertical) filament (in units of $10^{12} M_\odot$). Right: corresponding cross section of $\langle \cos \hat{\theta} \rangle(r, z)$. The mass of halos increases towards the nodes, while the spin flips.

geometric split \longrightarrow mass split

Geometry of the saddle provides a natural ‘metric’ (local frame as defined by Hessian @ saddle) relative to which **dynamical evolution** of DH is predicted.

Cloud in
cloud effect

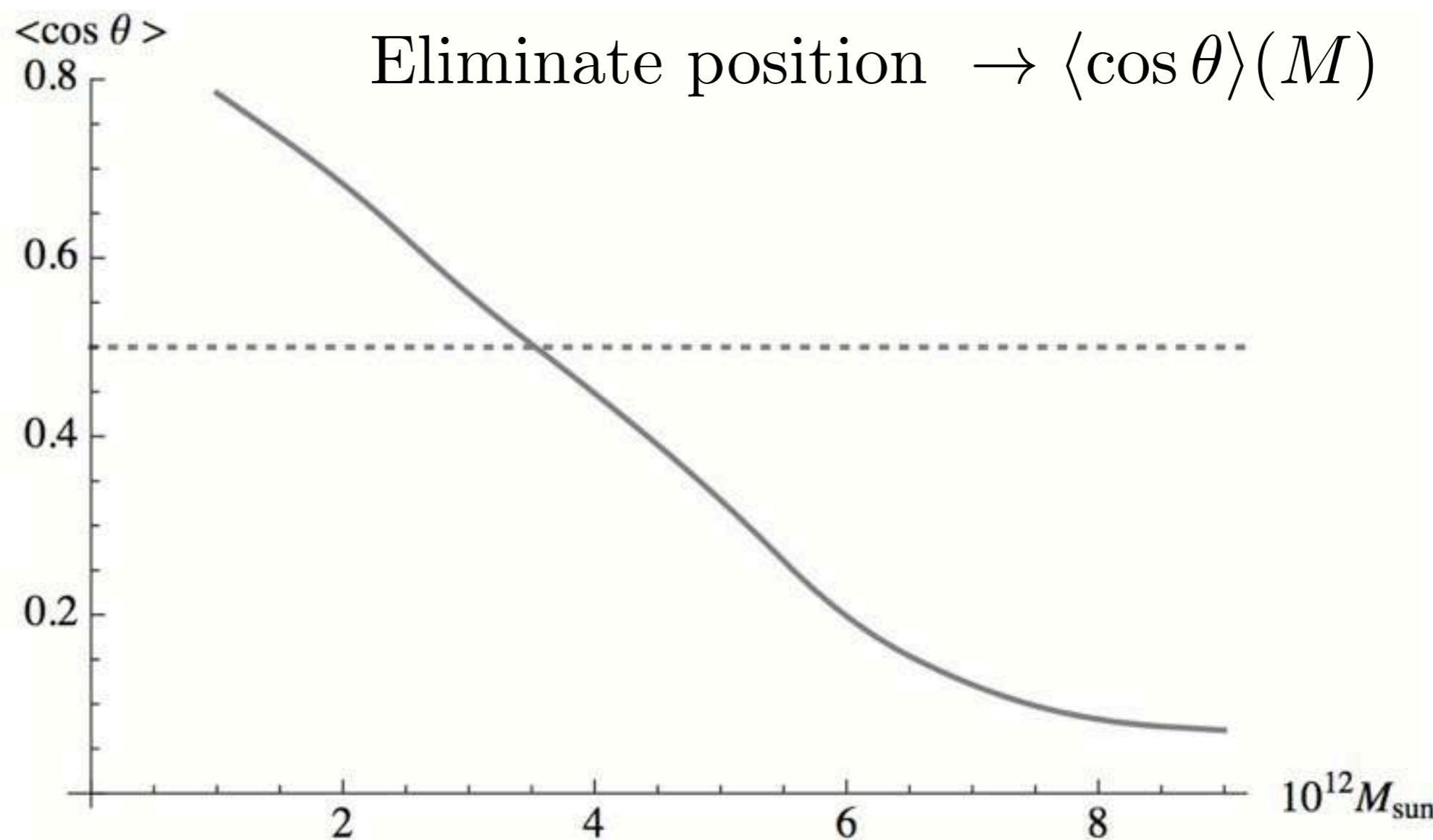


Figure 6. Mean alignment between spin and filament as a function of mass for a filament smoothing scale of 5 Mpc/ h . The spin flip transition mass is around $4 \cdot 10^{12} M_{\odot}$.

geometric split \longrightarrow mass split

Link with Eulerian vorticity?

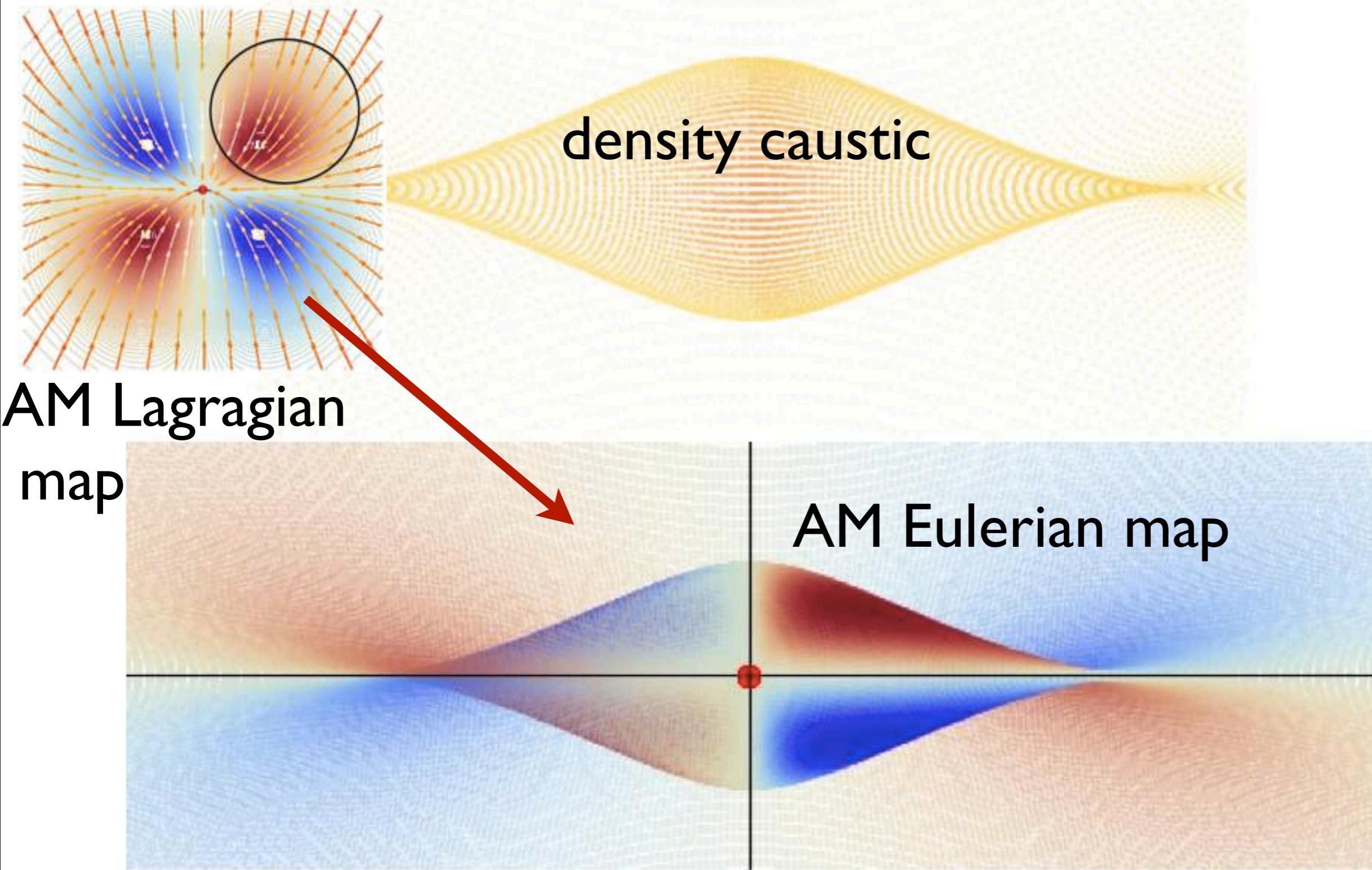
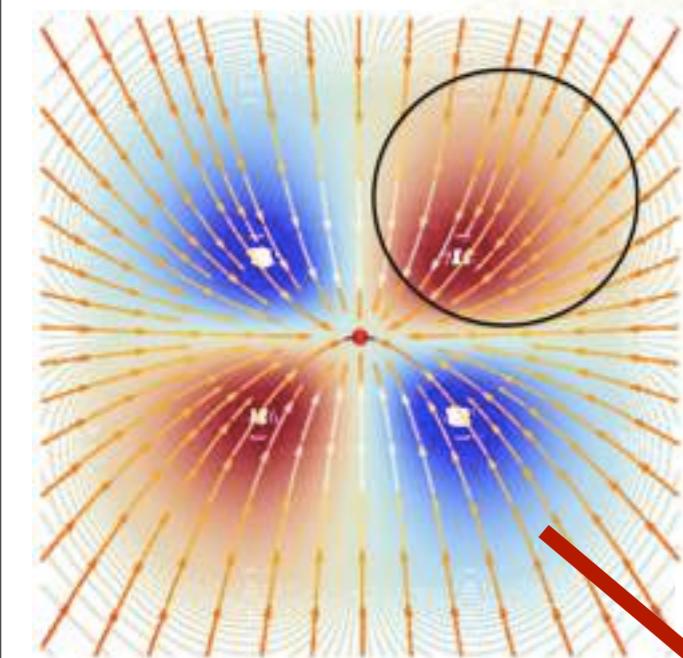


Figure 5. top: Density caustic; Bottom: Zeldovitch mapping of the spin distribution

Link with Eulerian vorticity?



AM Lagrangian
map

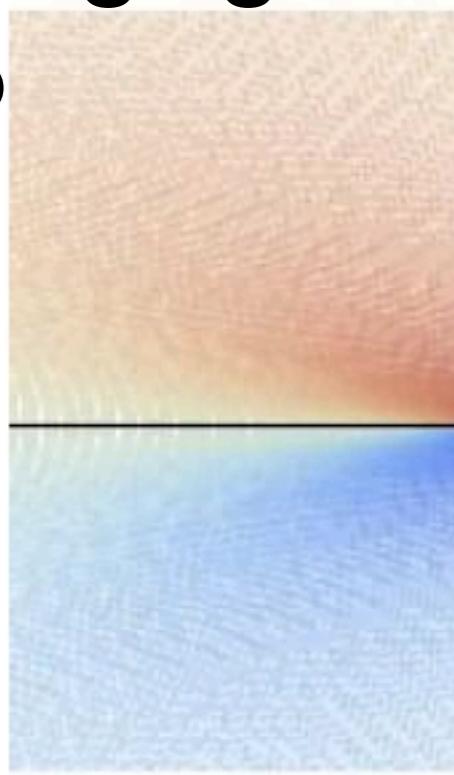
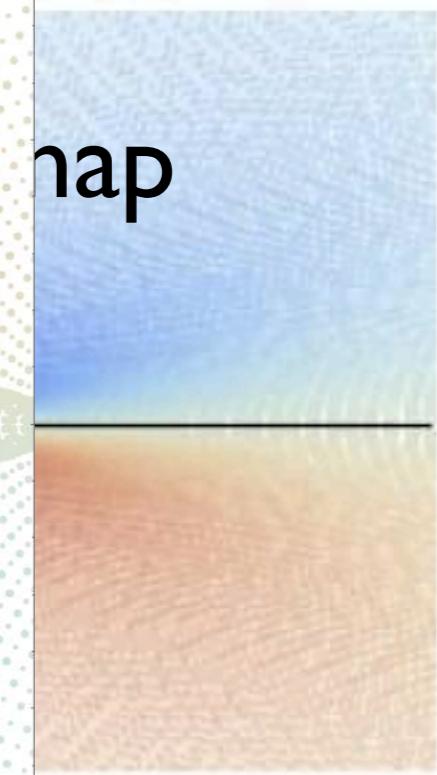
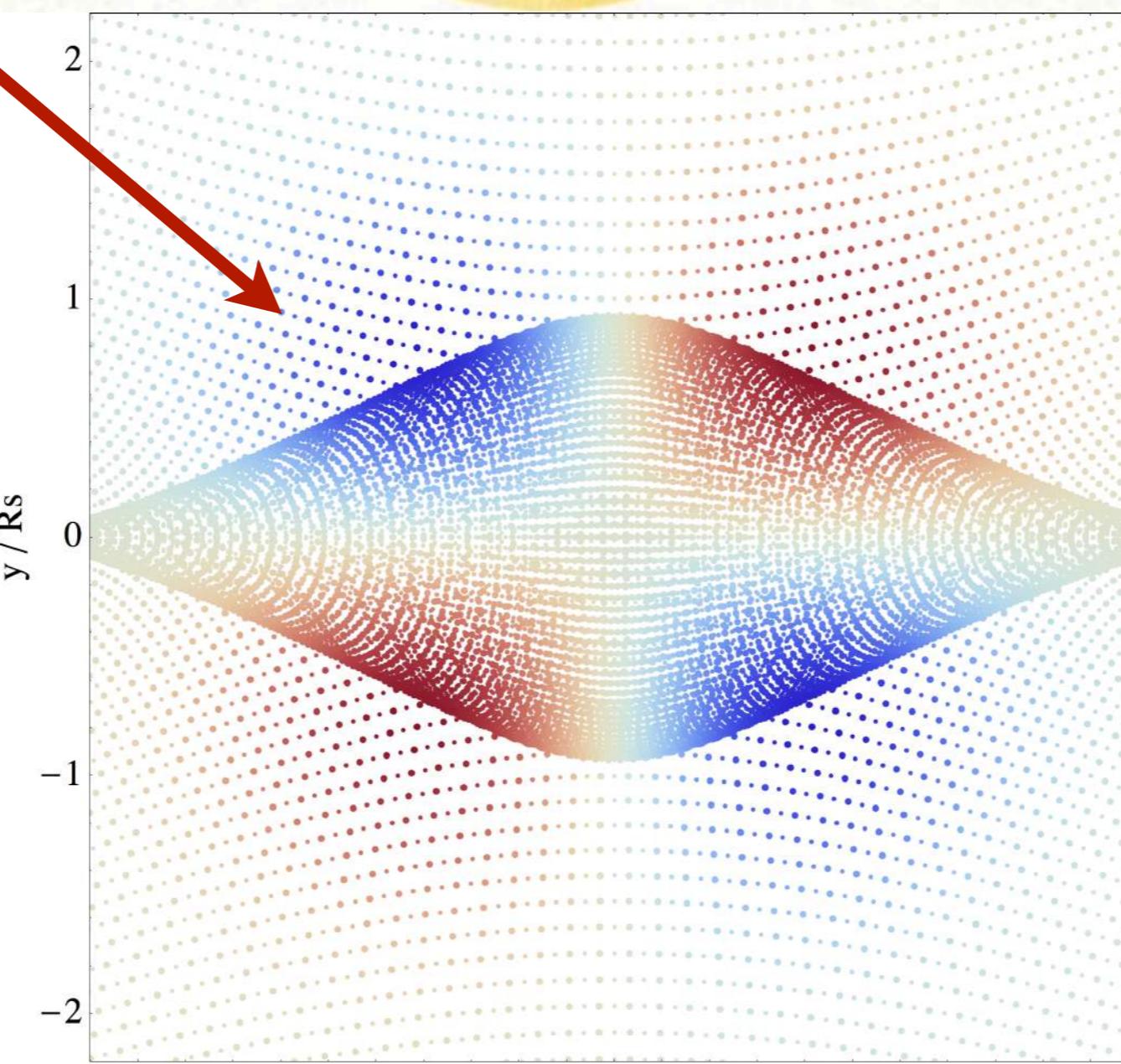
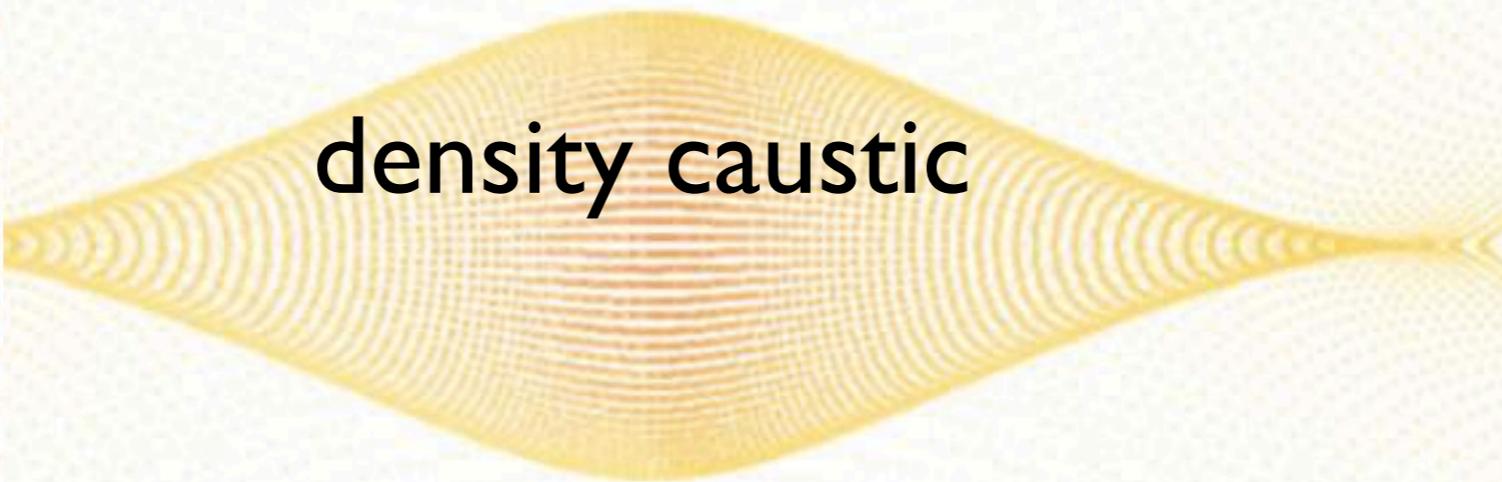
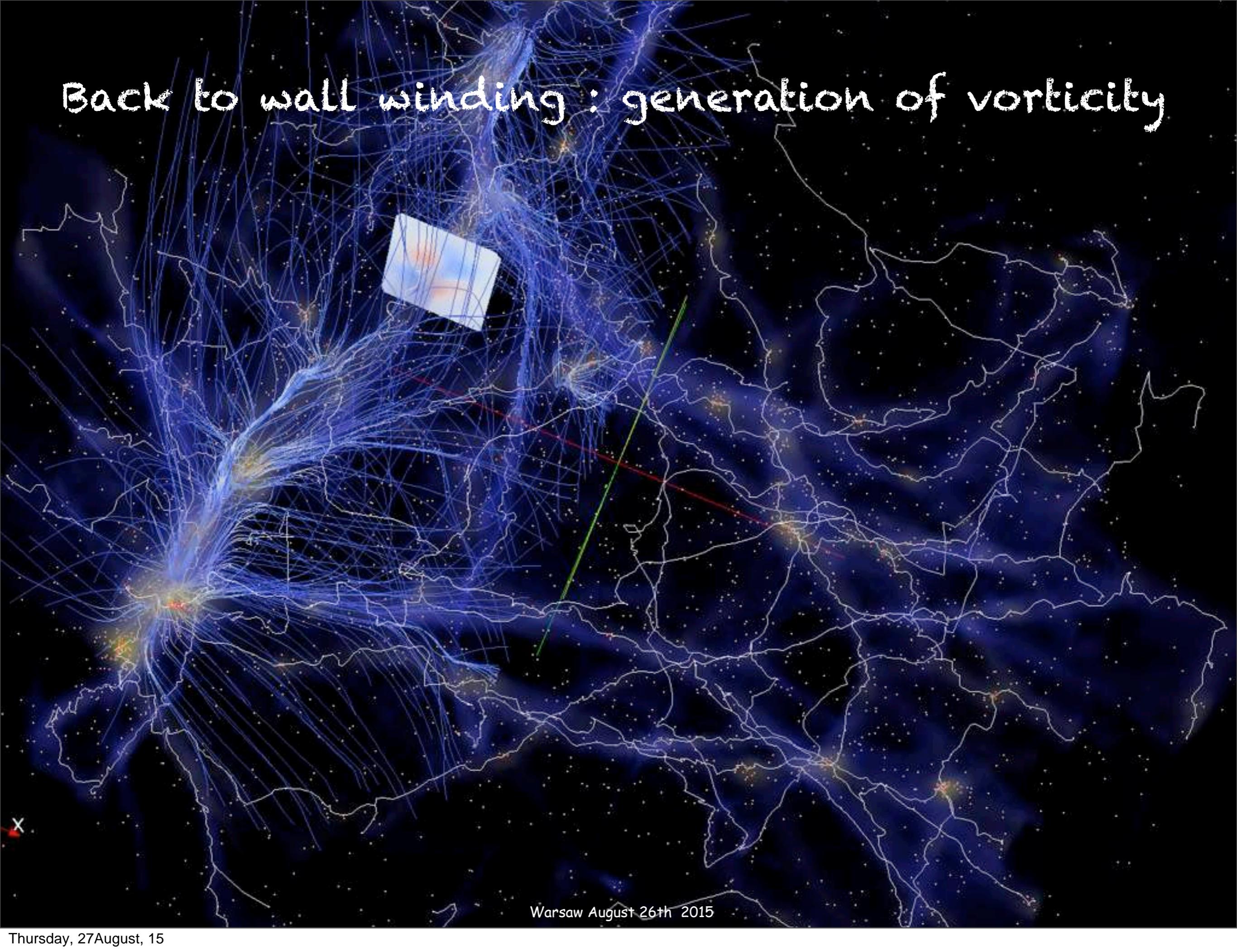


Figure 5. top
the spin distri



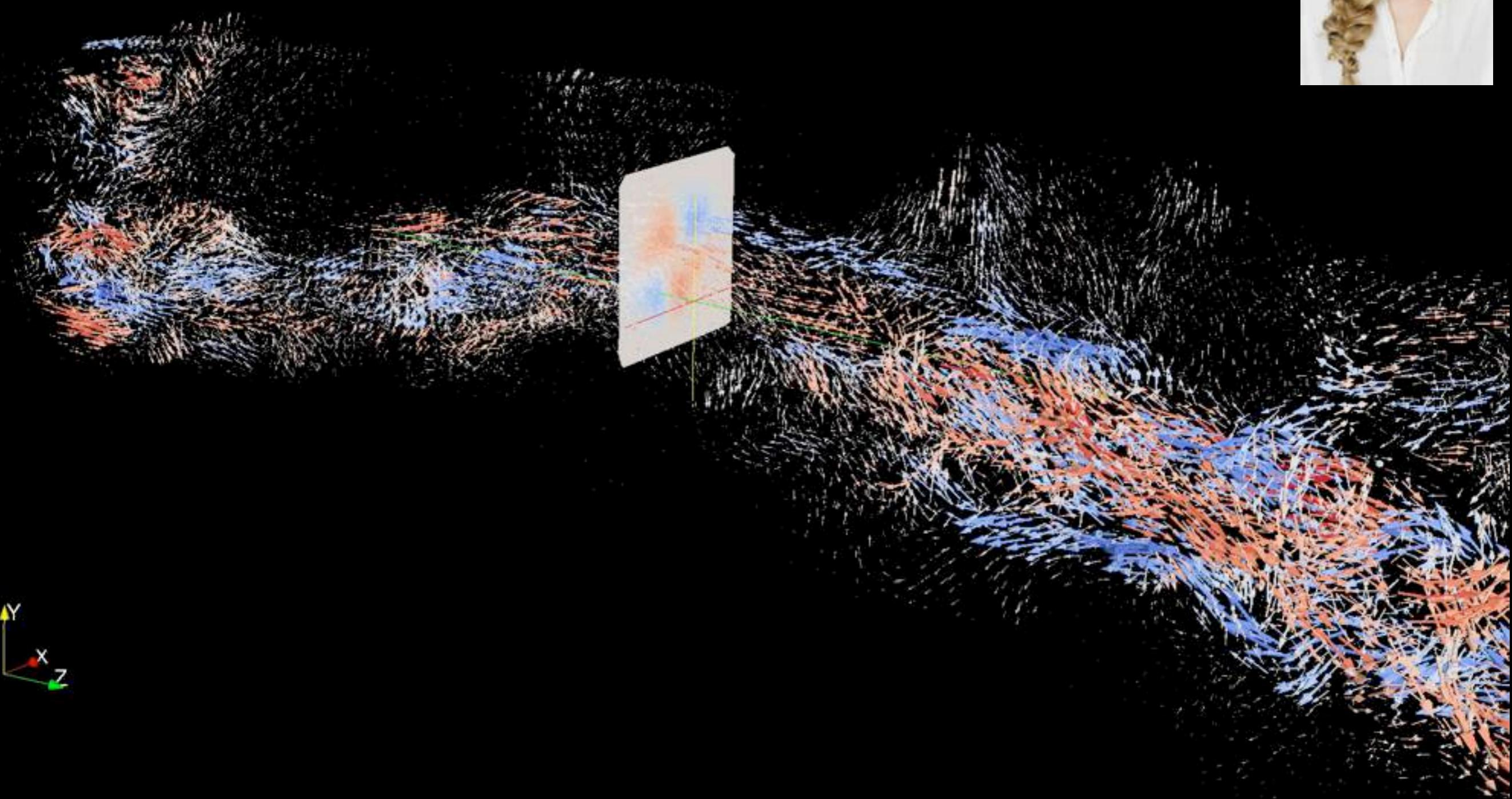
mapping of

Back to wall winding : generation of vorticity



Warsaw August 26th 2015

Alignment of vorticity with cosmic web



braids structure of vorticity.

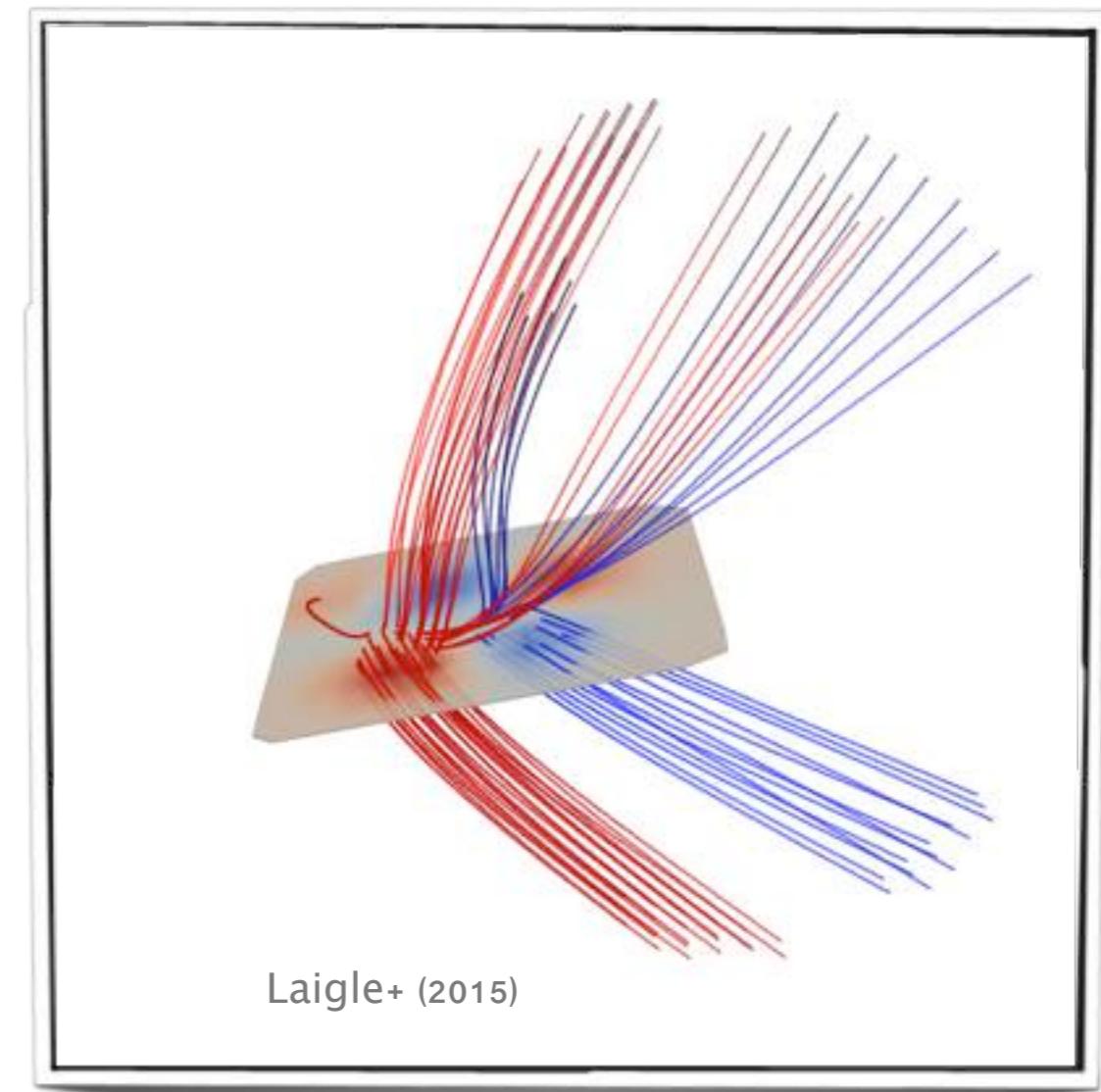
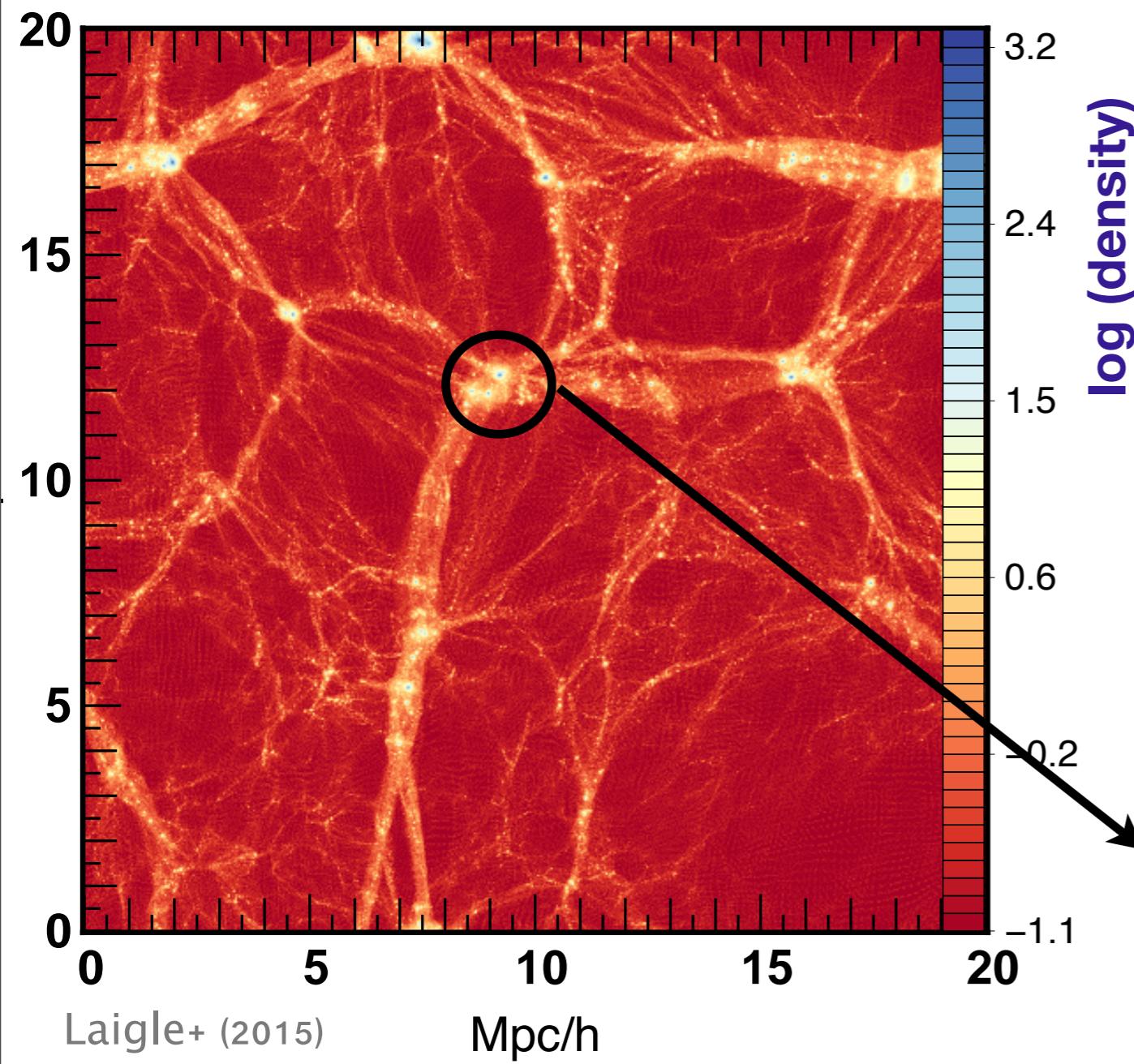
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Growth of large-scale structure

In the initial phase of structure formation, **flows are laminar and curl-free**.

This is no longer valid **at the shell-crossing**.

Thin slice of a DM simulation at $z=0$.



What happens when cosmic flows cross?

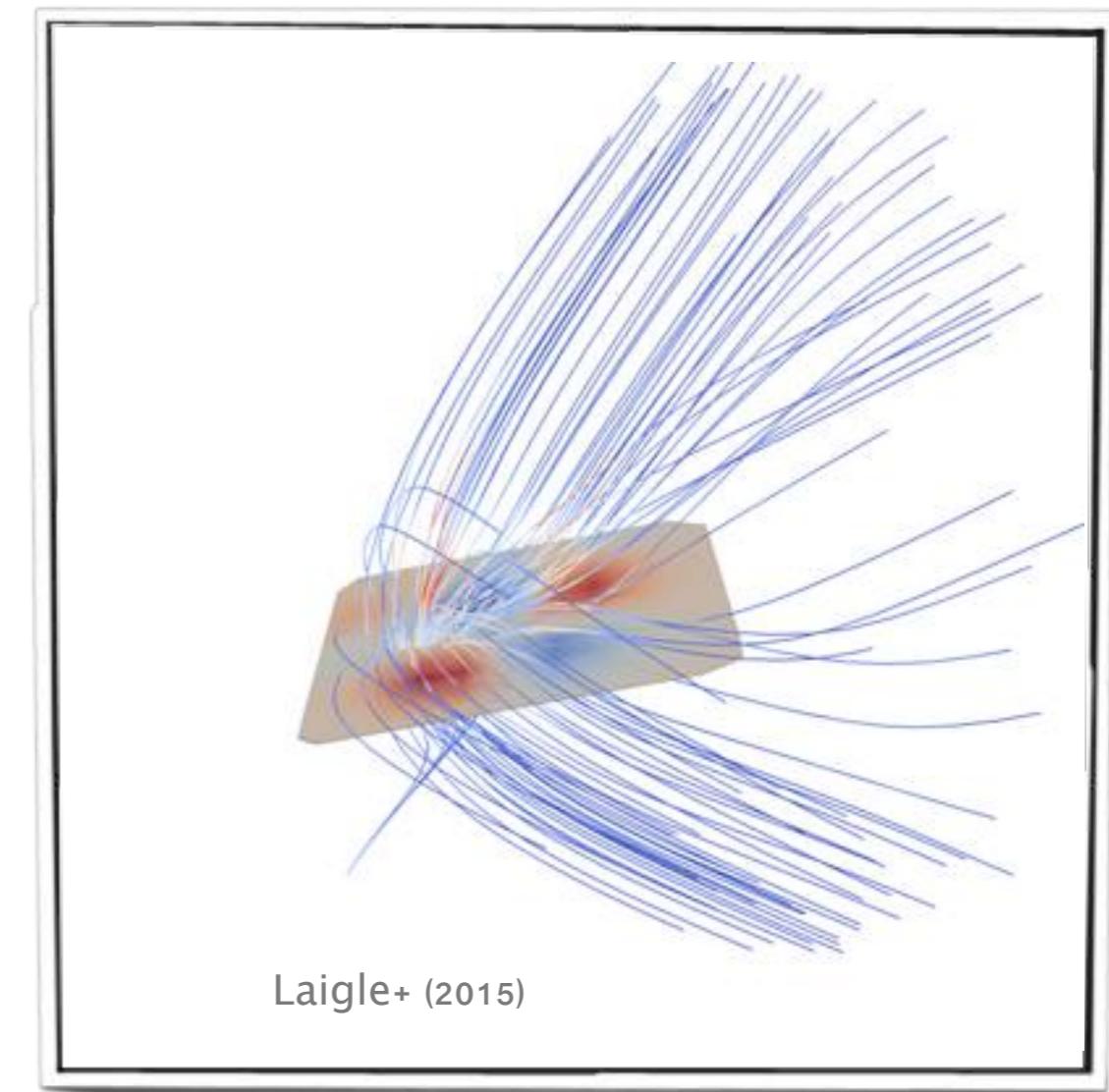
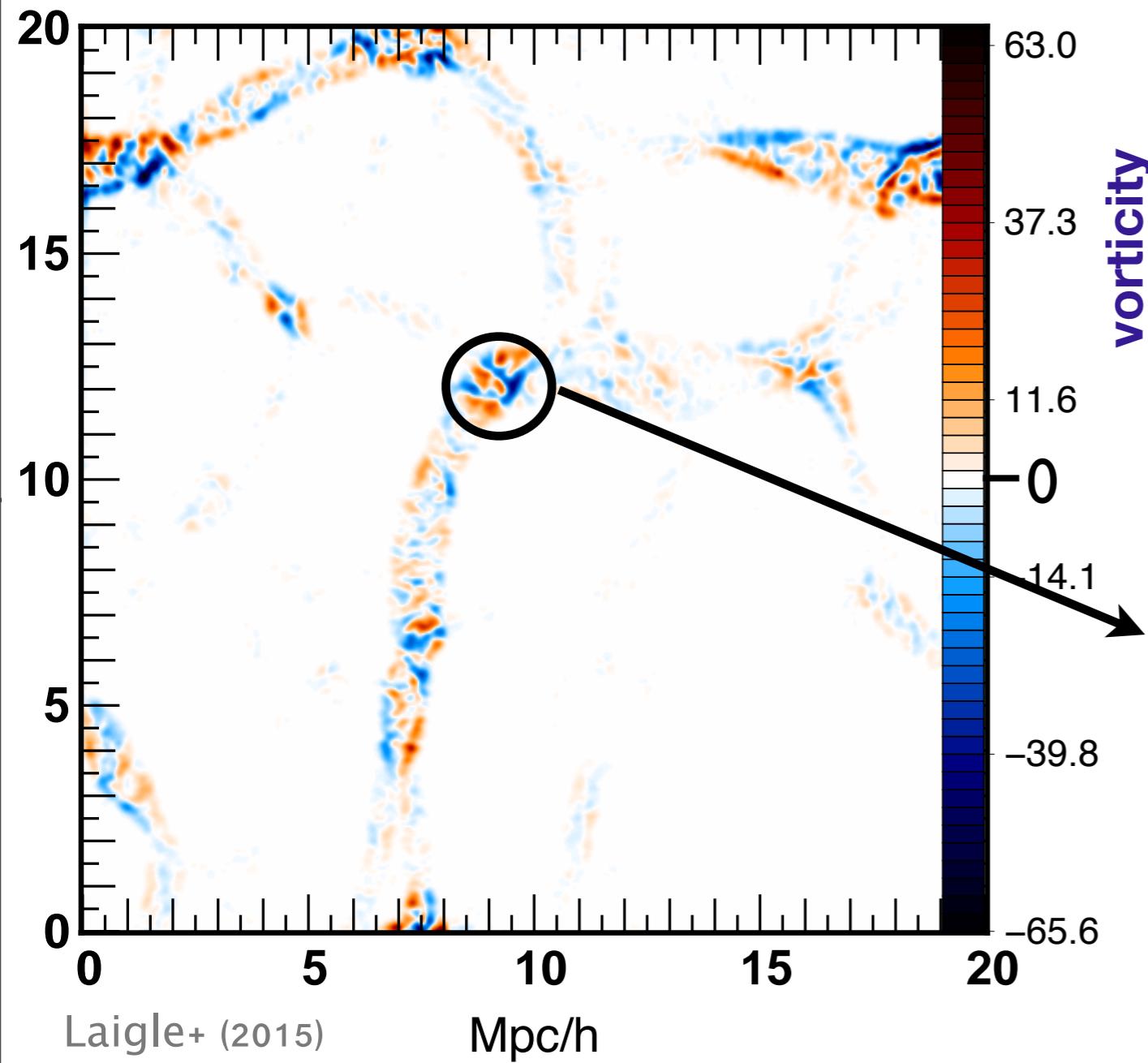
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Vorticity generation

In the initial phase of structure formation, **flows are laminar and curl-free.**

This is no longer valid **at the shell-crossing**.

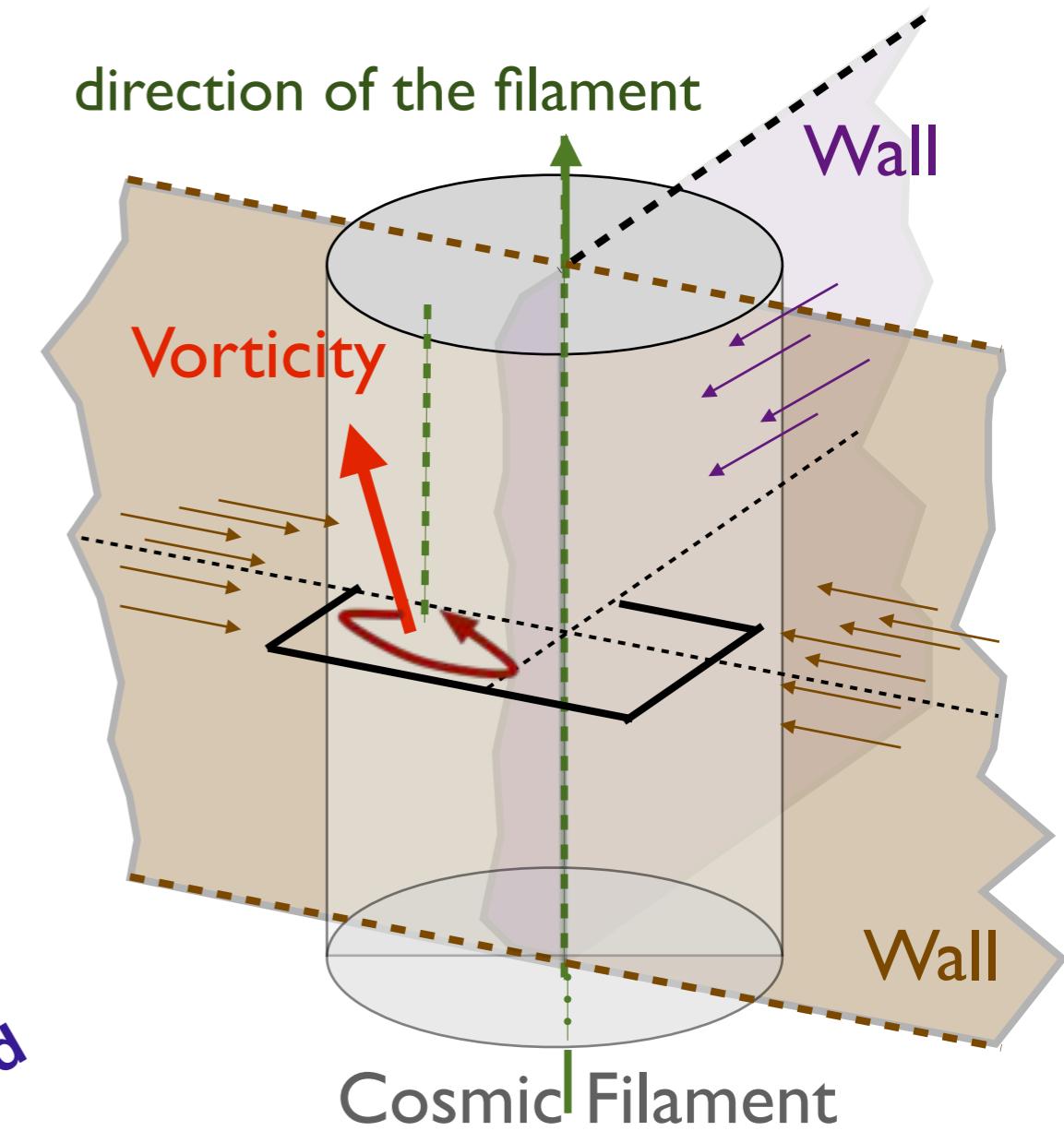
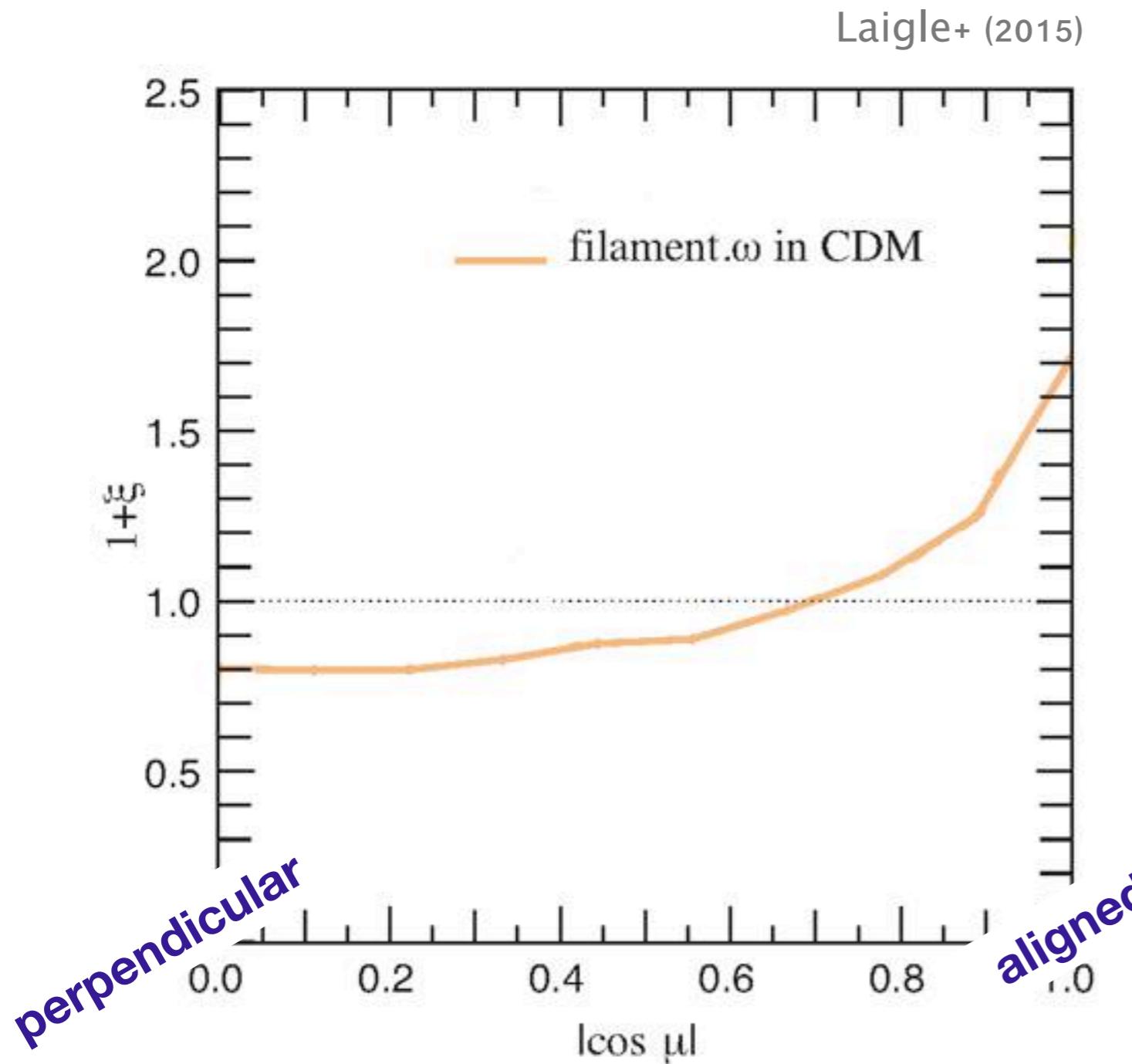
Thin slice of a DM simulation at $z=0$.



Vorticity is generated and is confined in the filaments.

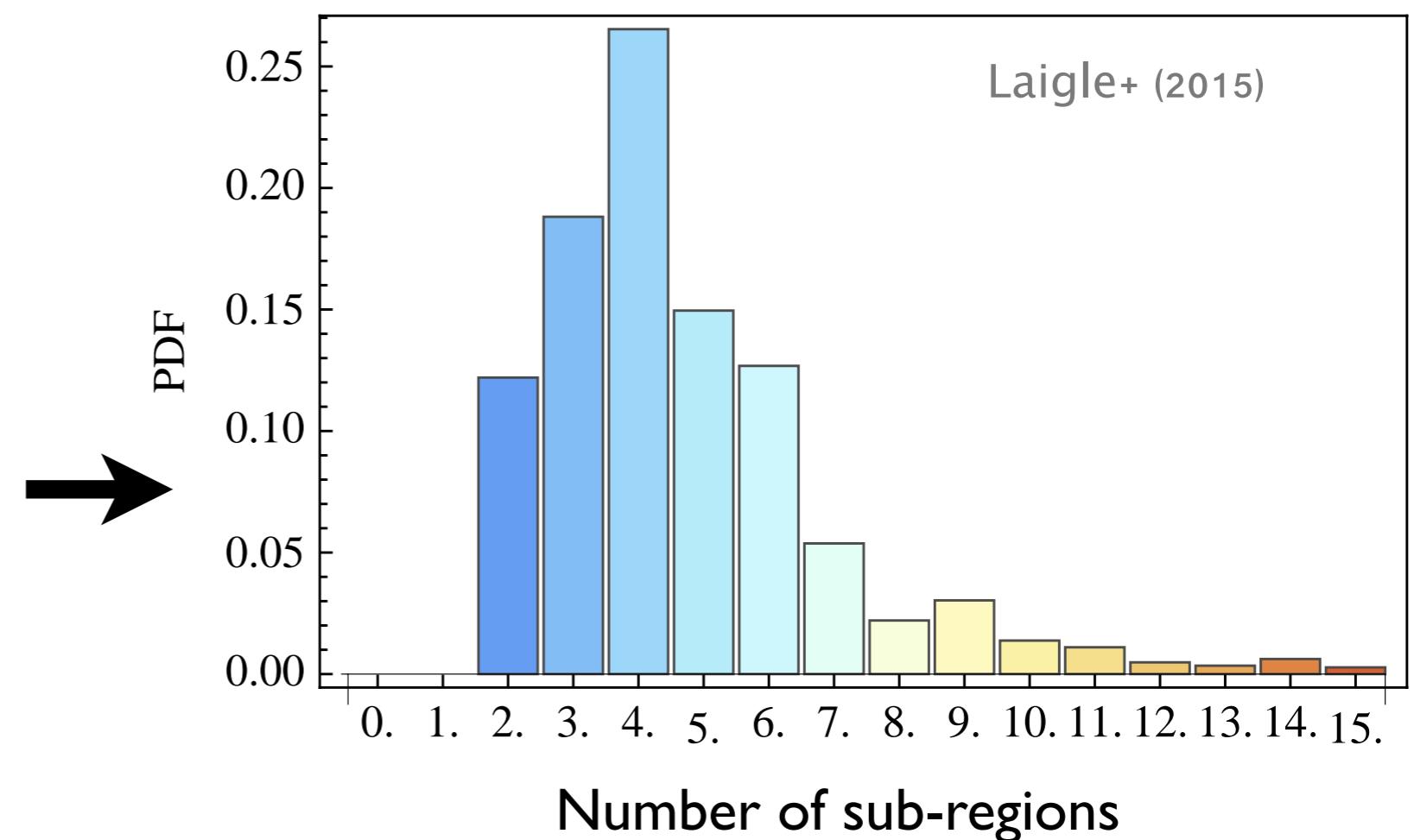
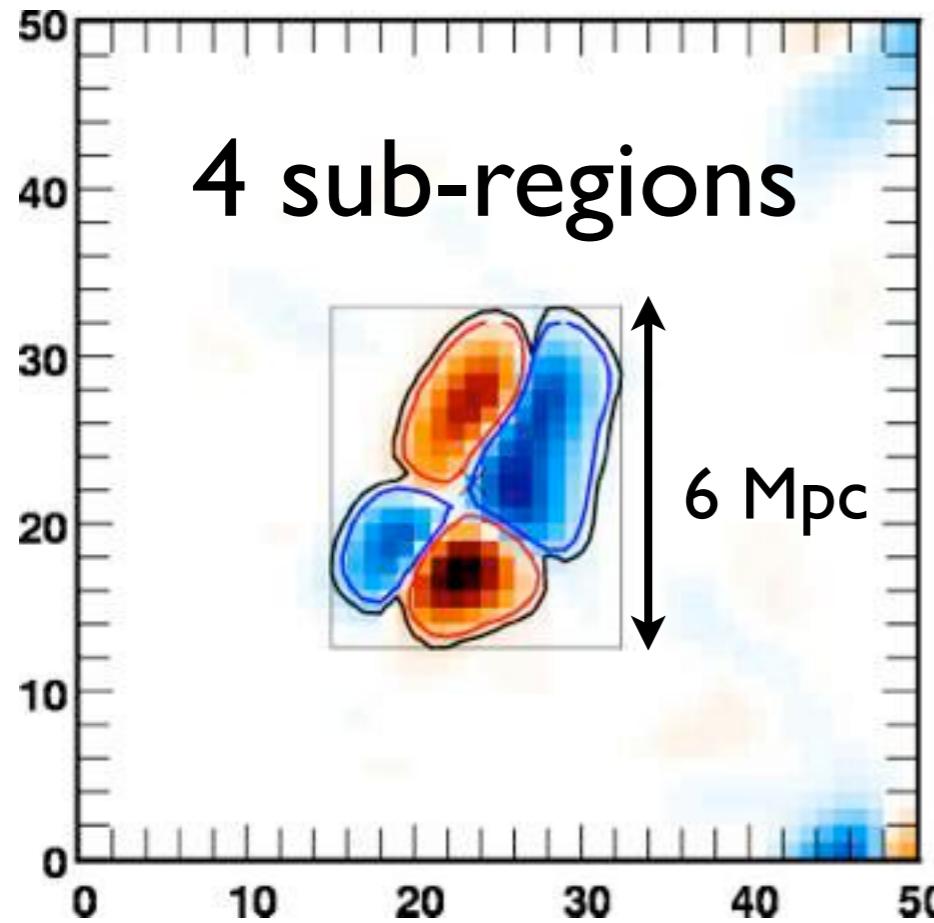
Alignment of vorticity with filaments

Vorticity is aligned with the filaments.

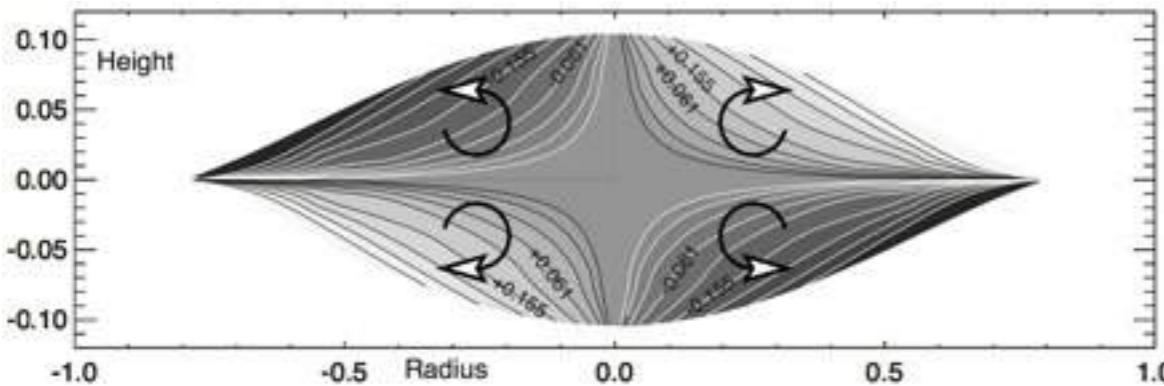


Filaments and walls are identified with DISPERSE: Sousbie+ (2011).

Geometry of the vorticity cross-section



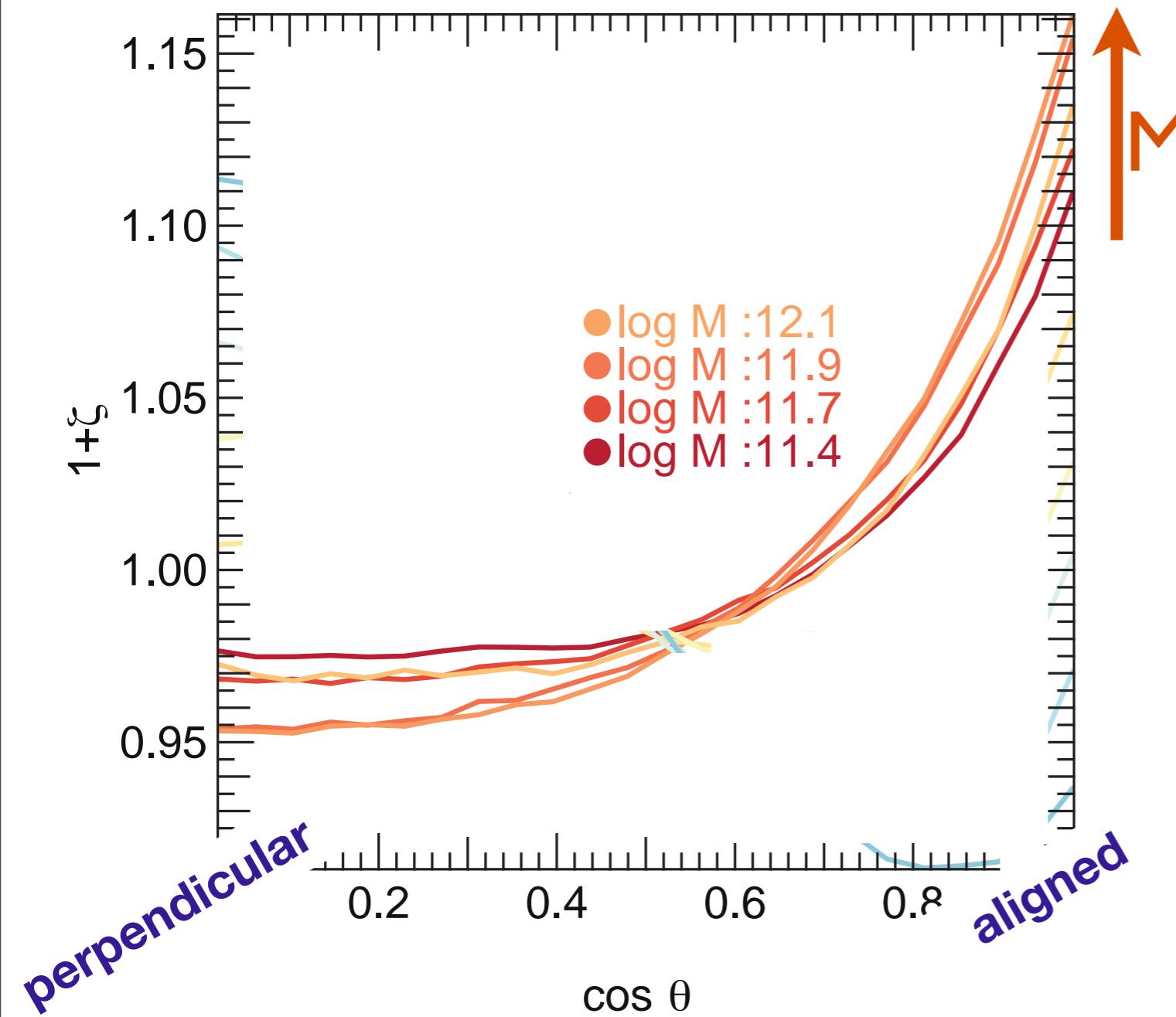
Pichon & Bernardeau (1999)



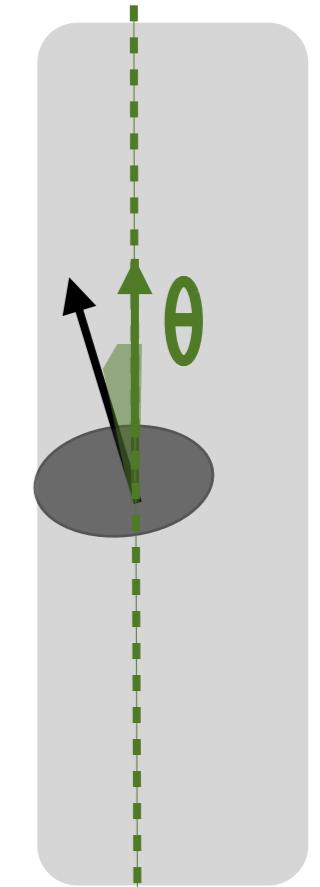
Theoretical prediction from Pichon & Bernardeau 1999

Cross-sections are typically divided in 4 quadrants.

Low-mass halos are aligned with the filament



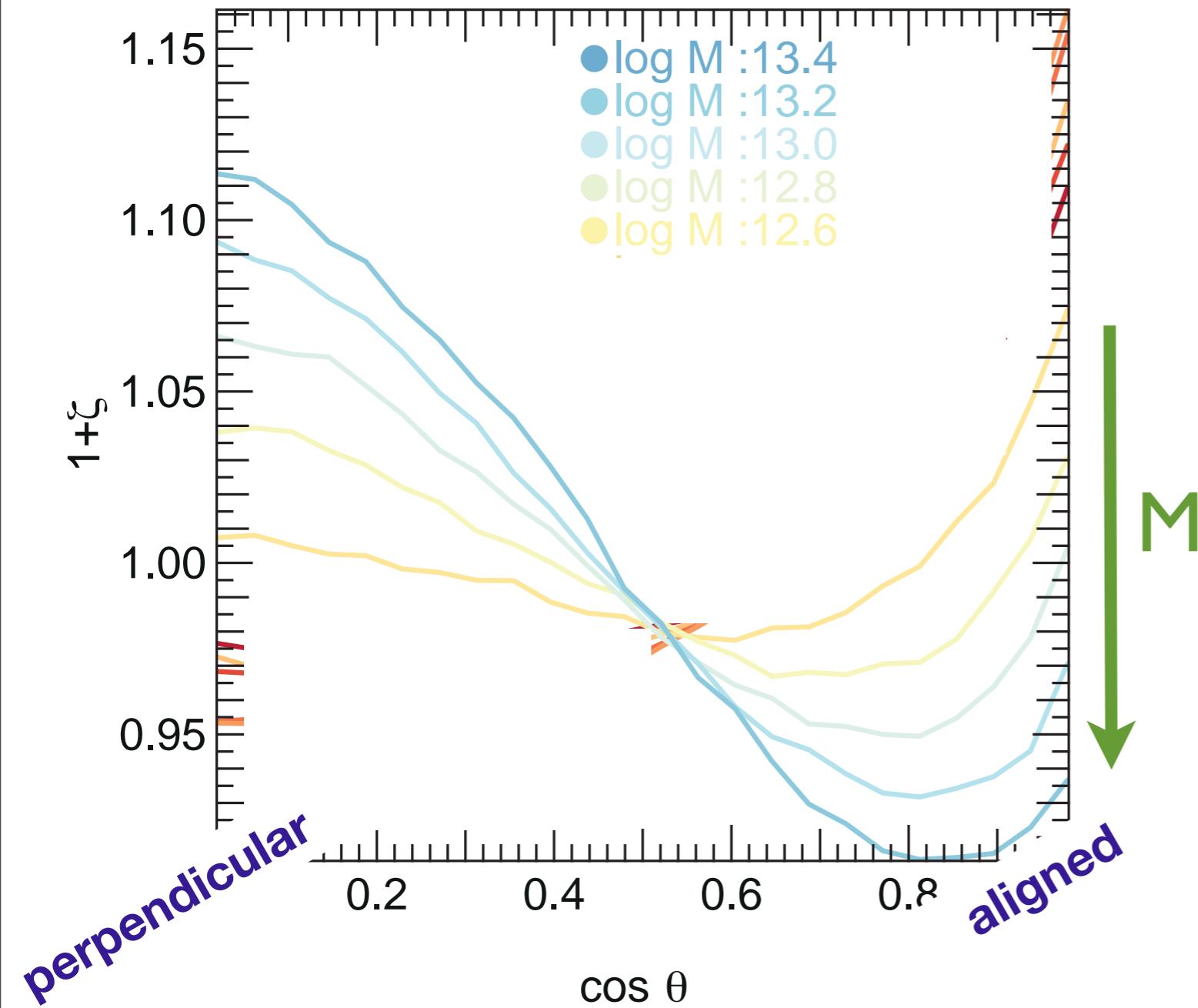
- $M_{\text{halo}} < M_{\text{crit}}$
alignment of halo spin with filament increases with mass.
- M_{crit}



Laigle+ (2015)
in DM simulation (10^5 haloes)

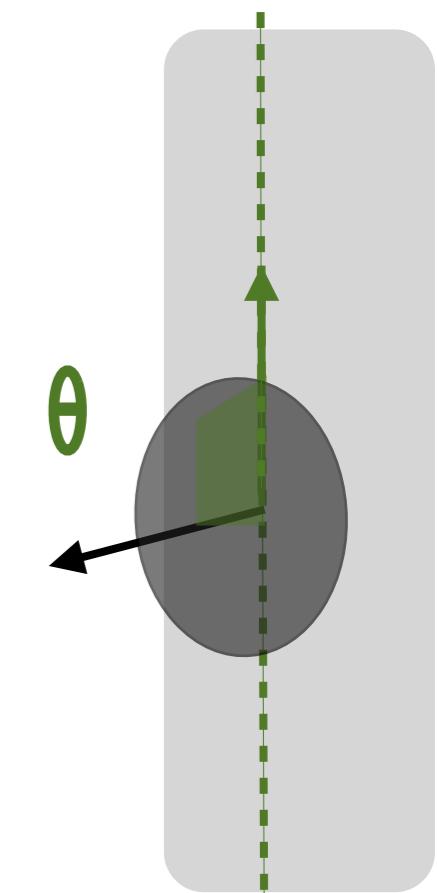
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High-mass halos are perpendicular



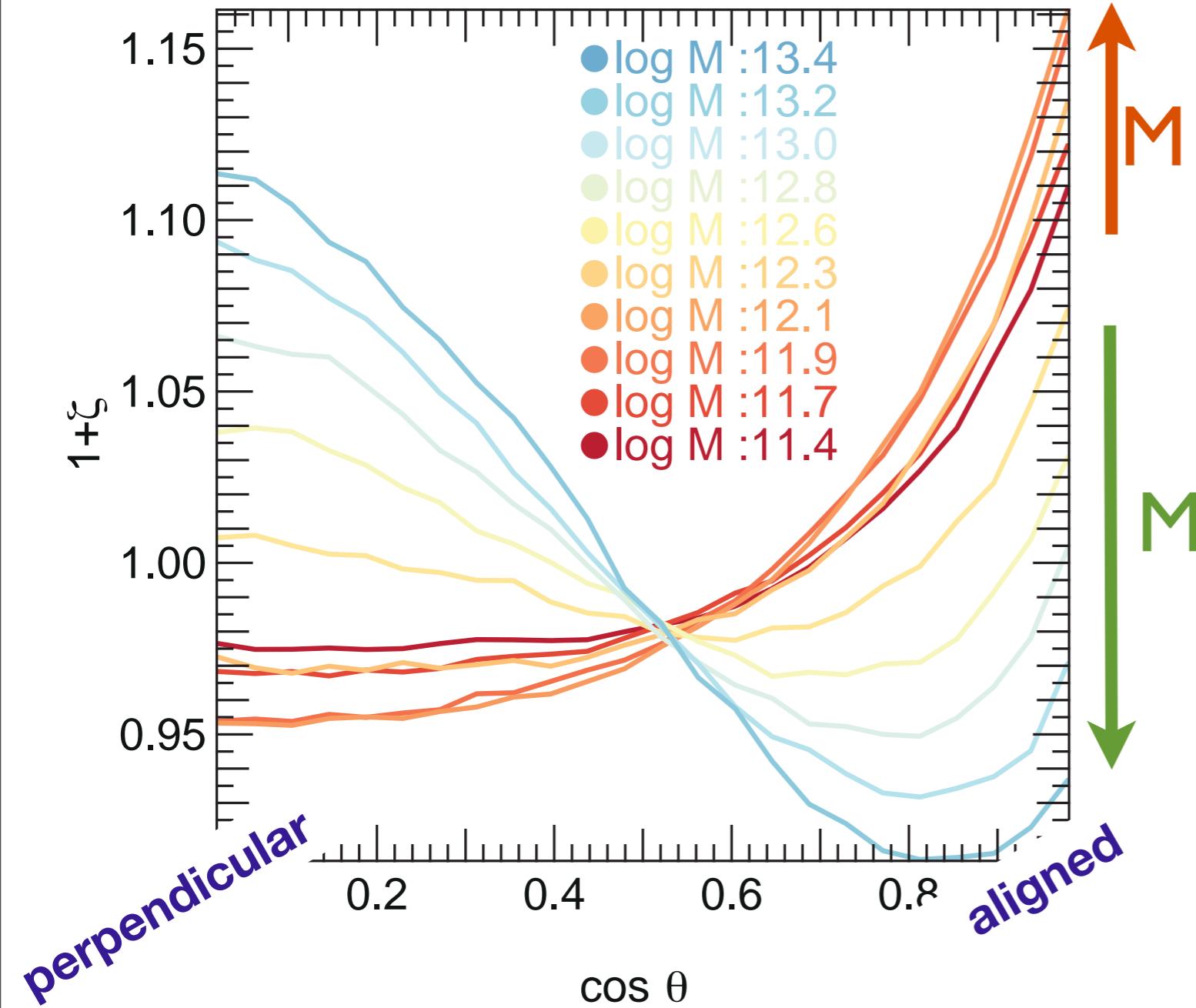
Laigle+ (2015)
in DM simulation (10^5 haloes)

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- M_{crit}
- $M_{\text{halo}} > M_{\text{crit}}$
- halo spin tends to be **perpendicular** to the filament.

Mass dependent Halo spin - filament alignment



- $M_{\text{halo}} < M_{\text{crit}}$
alignment of halo spin
with filament increases
with mass.
- M_{crit}
- $M_{\text{halo}} > M_{\text{crit}}$
halo spin tends to be
perpendicular to the
filament.

Why?

When?

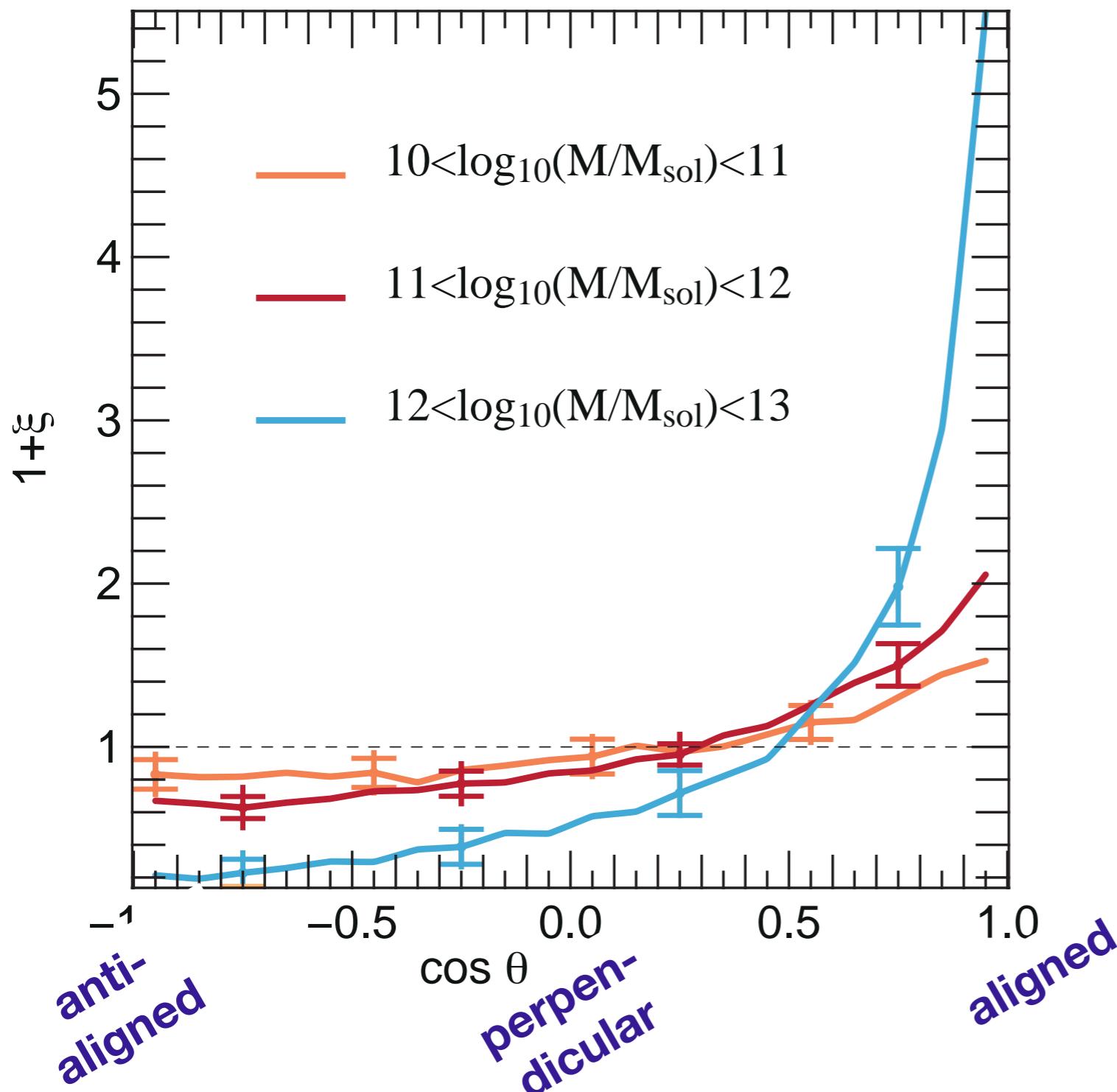
Why?

Laigle+ (2015)
in DM simulation (10^5 haloes)

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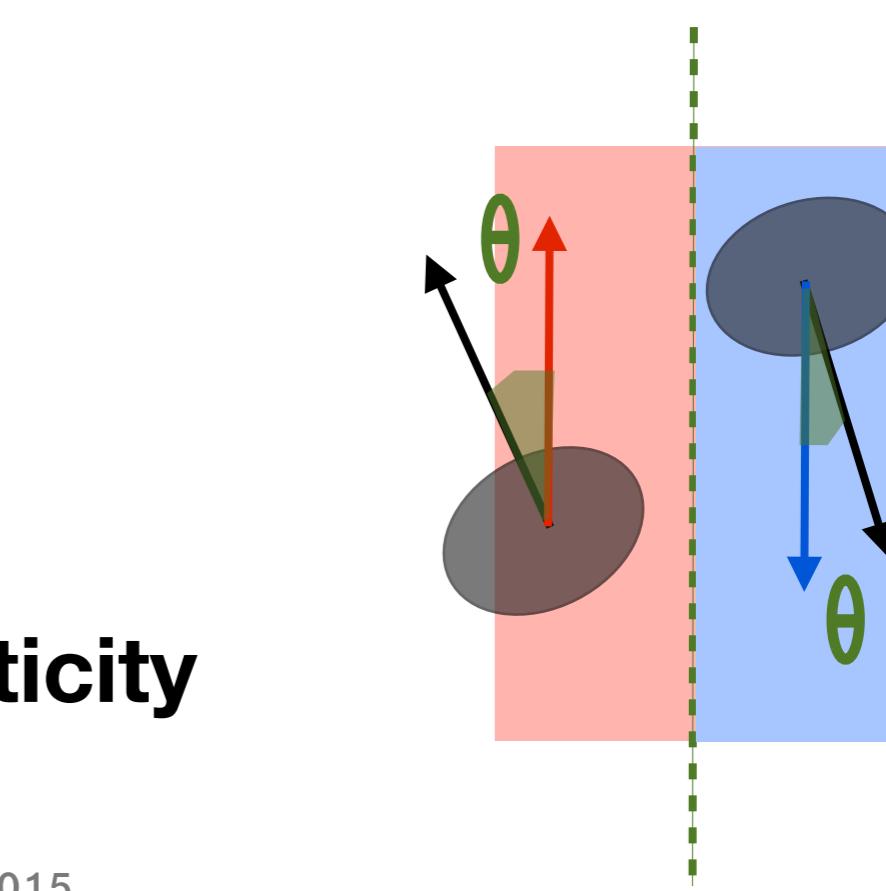
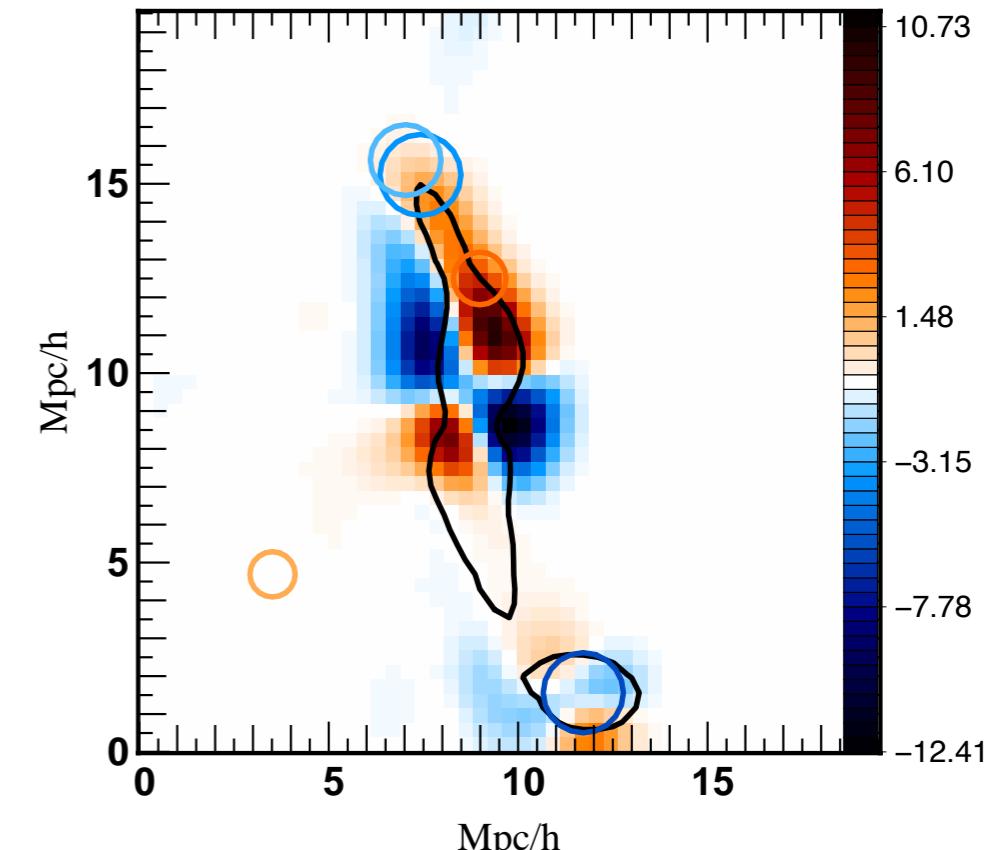
Halo-spin vorticity alignment

vorticity along the filament

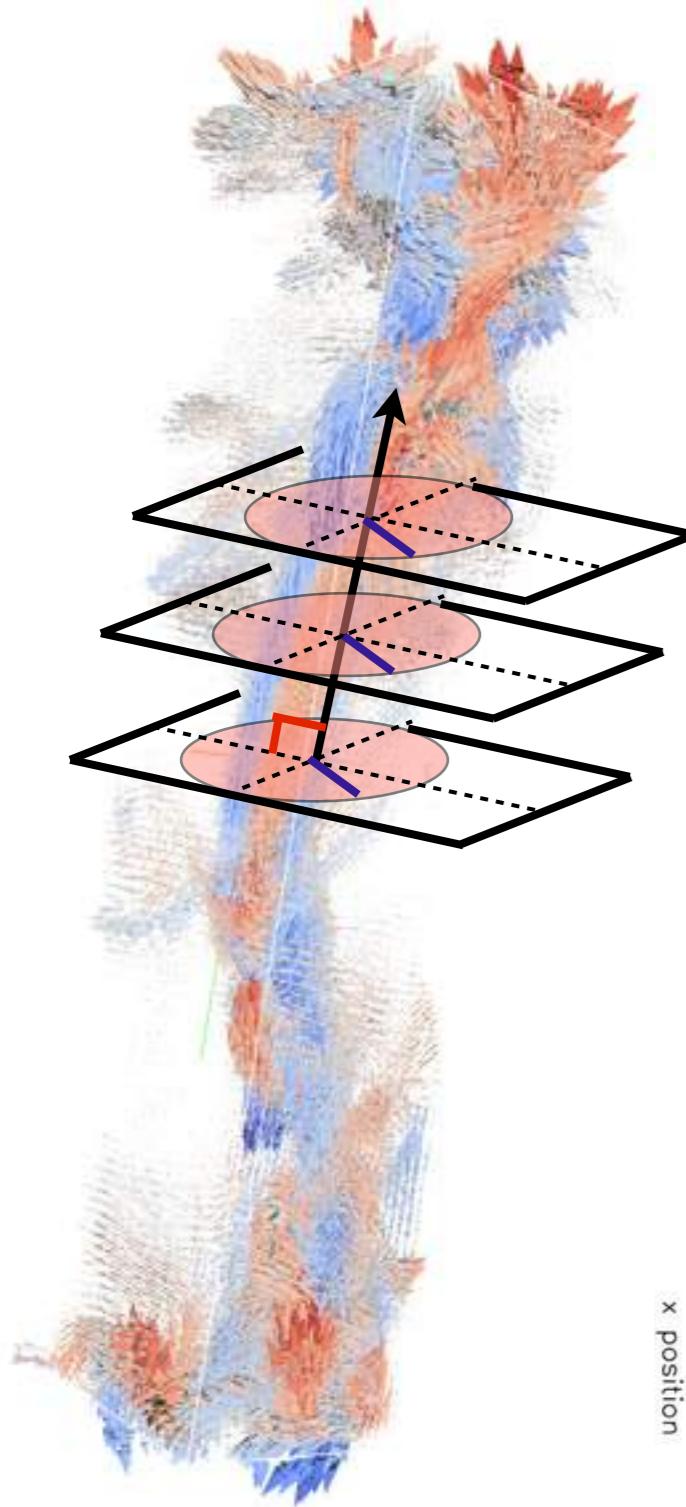


Halo spins are aligned with the vorticity in quadrants.

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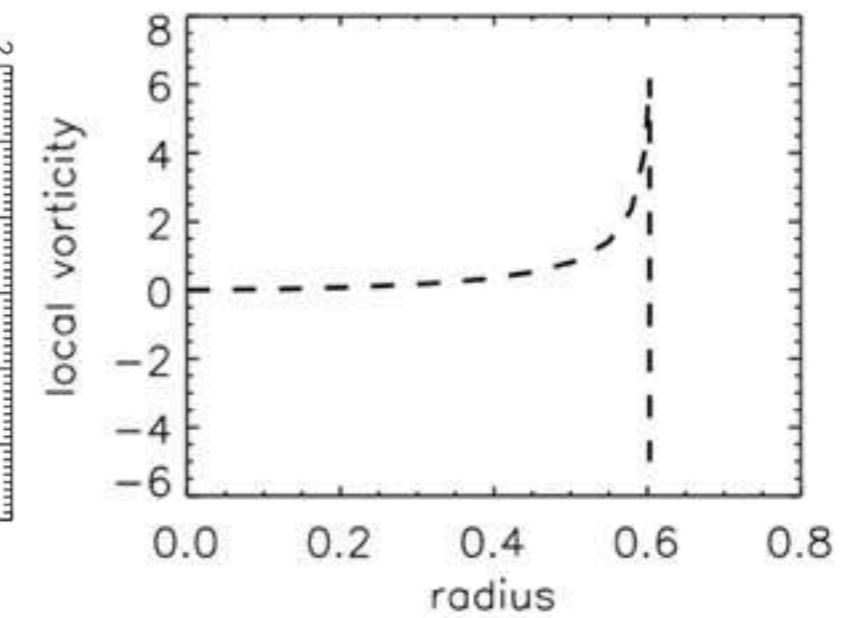
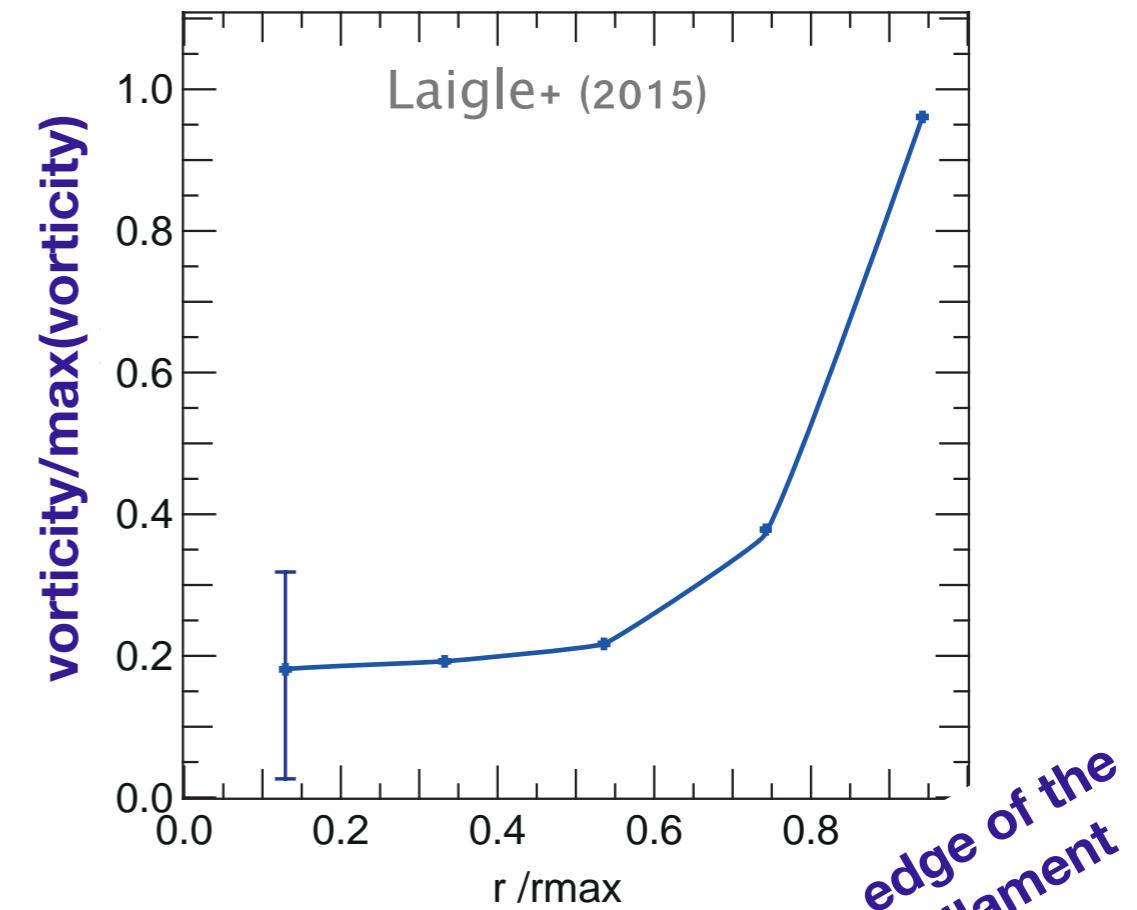
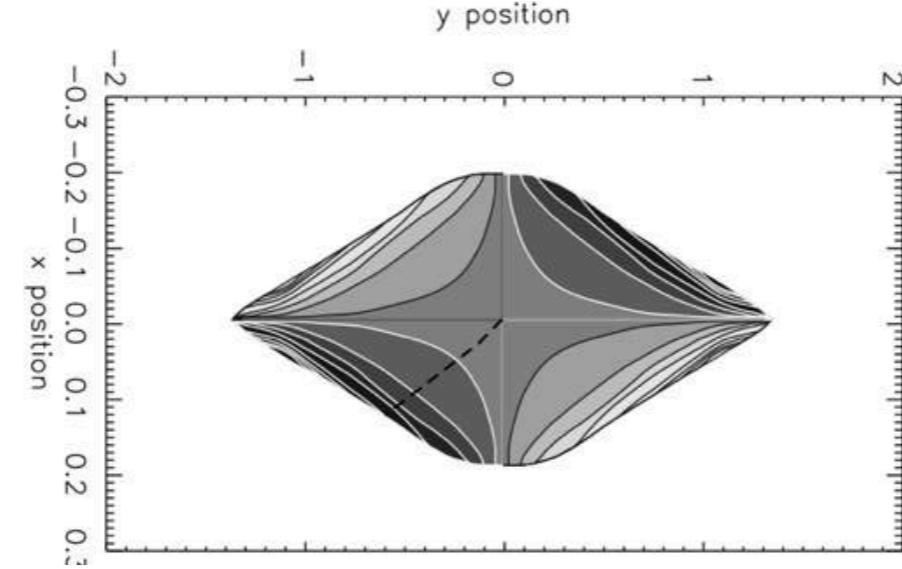


Geometry of the vorticity cross-section



Stacked profile

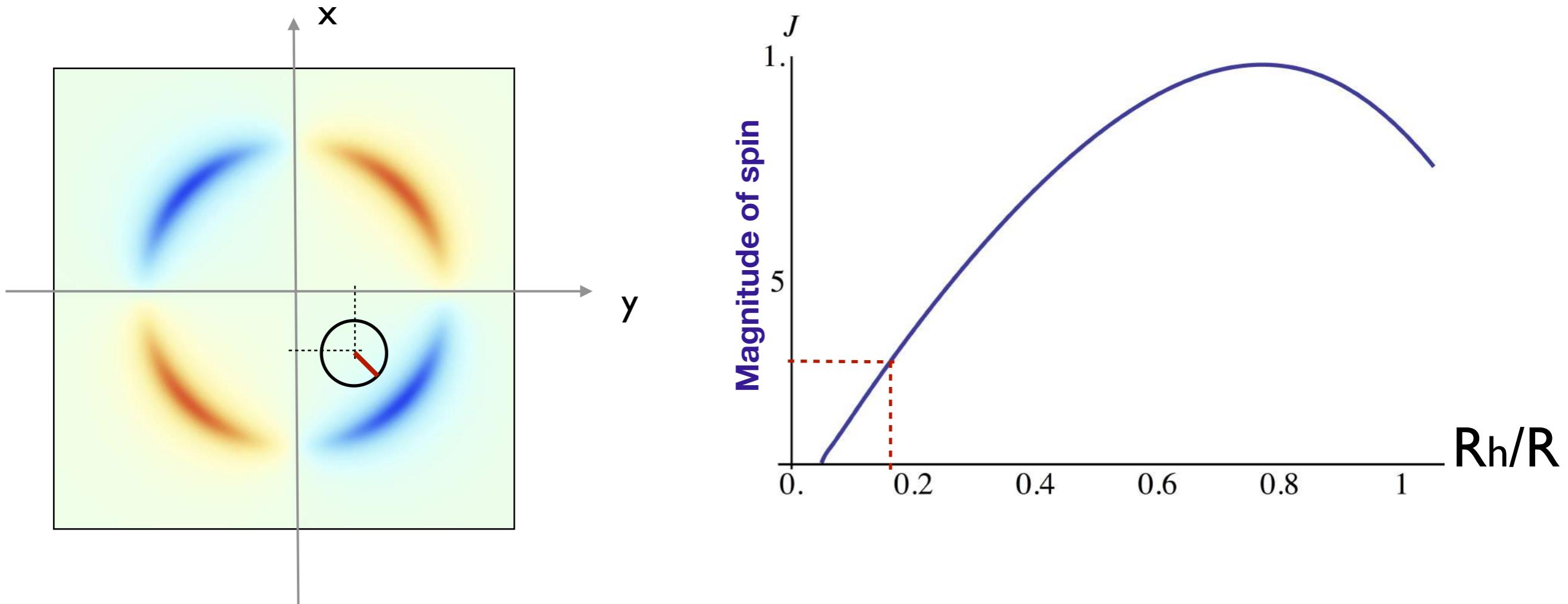
Pichon & Bernardeau (1999)



High vorticity regions are located at the edges of the filament.

Mass transition for spin alignment

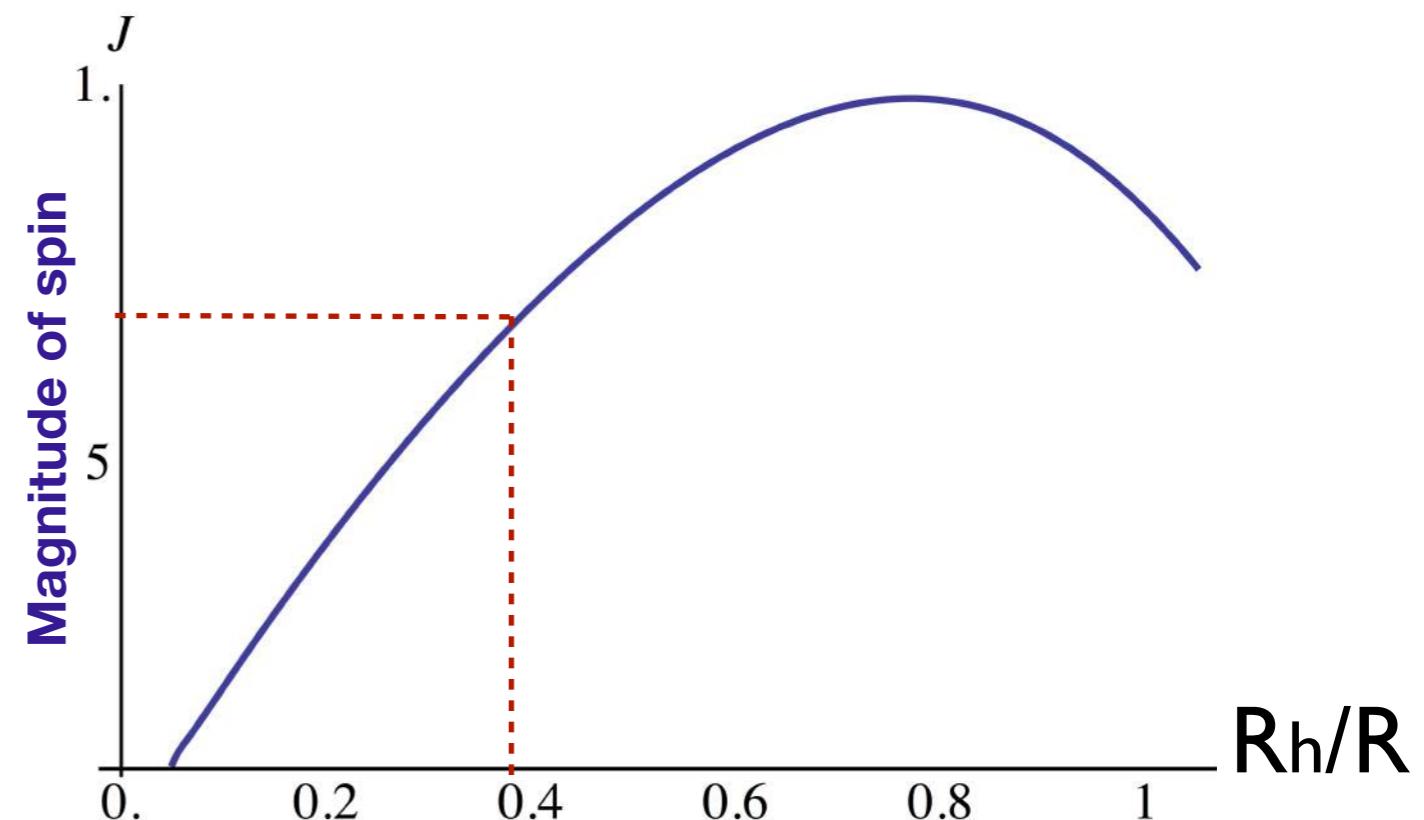
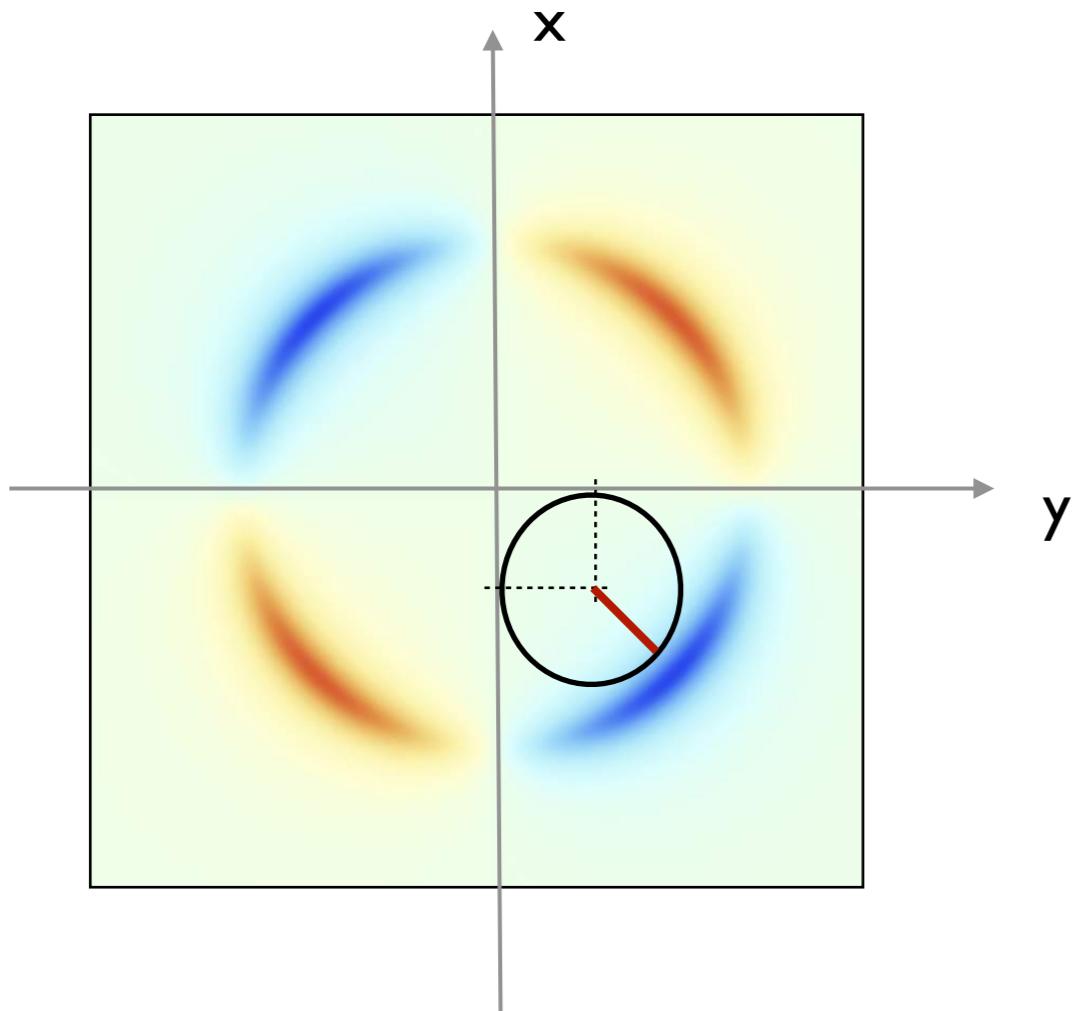
Idealized toy model: The position is fixed and the radius of the halo increases:



Transition mass is correlated with the size of the quadrants.

Mass transition for spin alignment

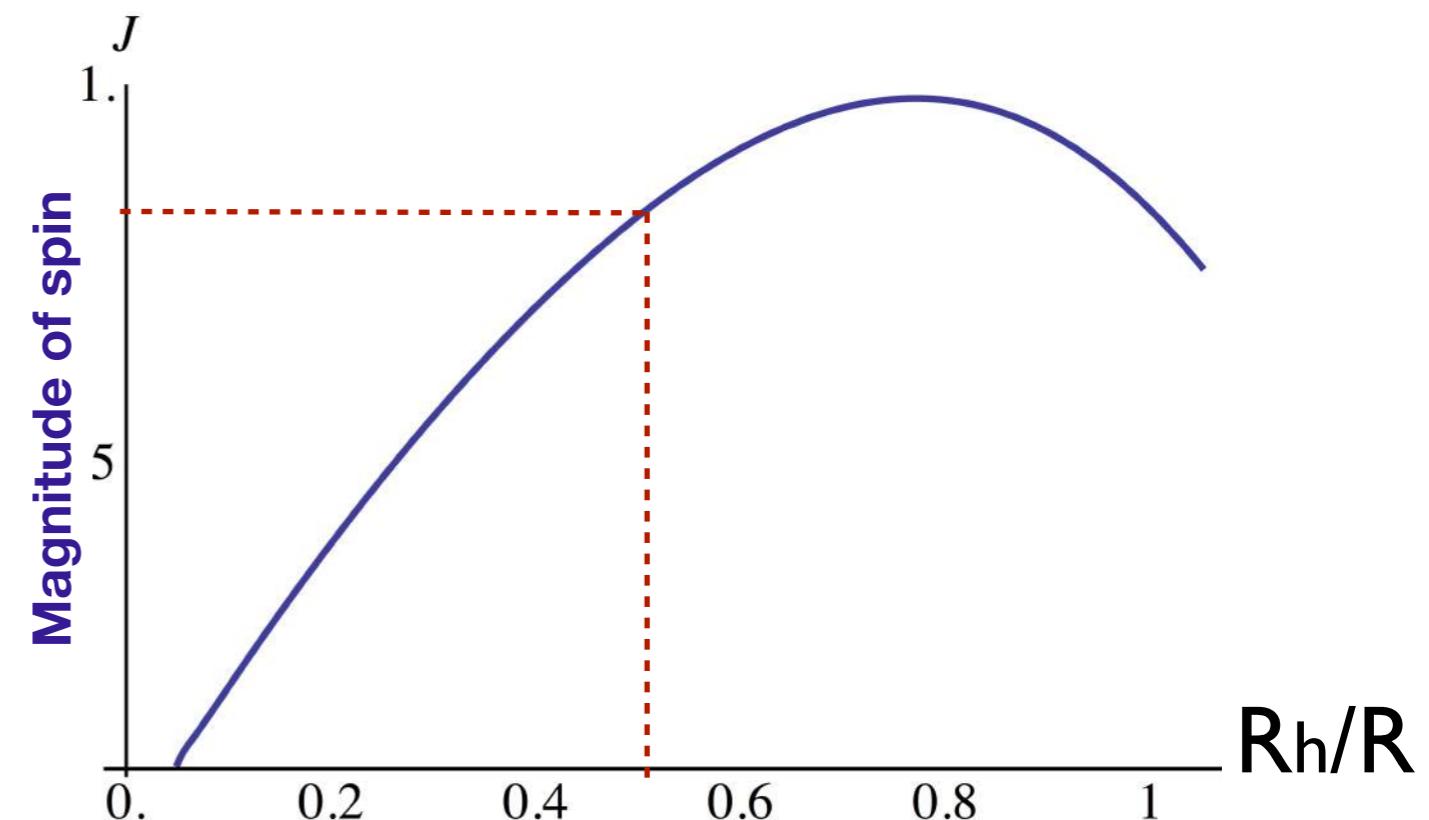
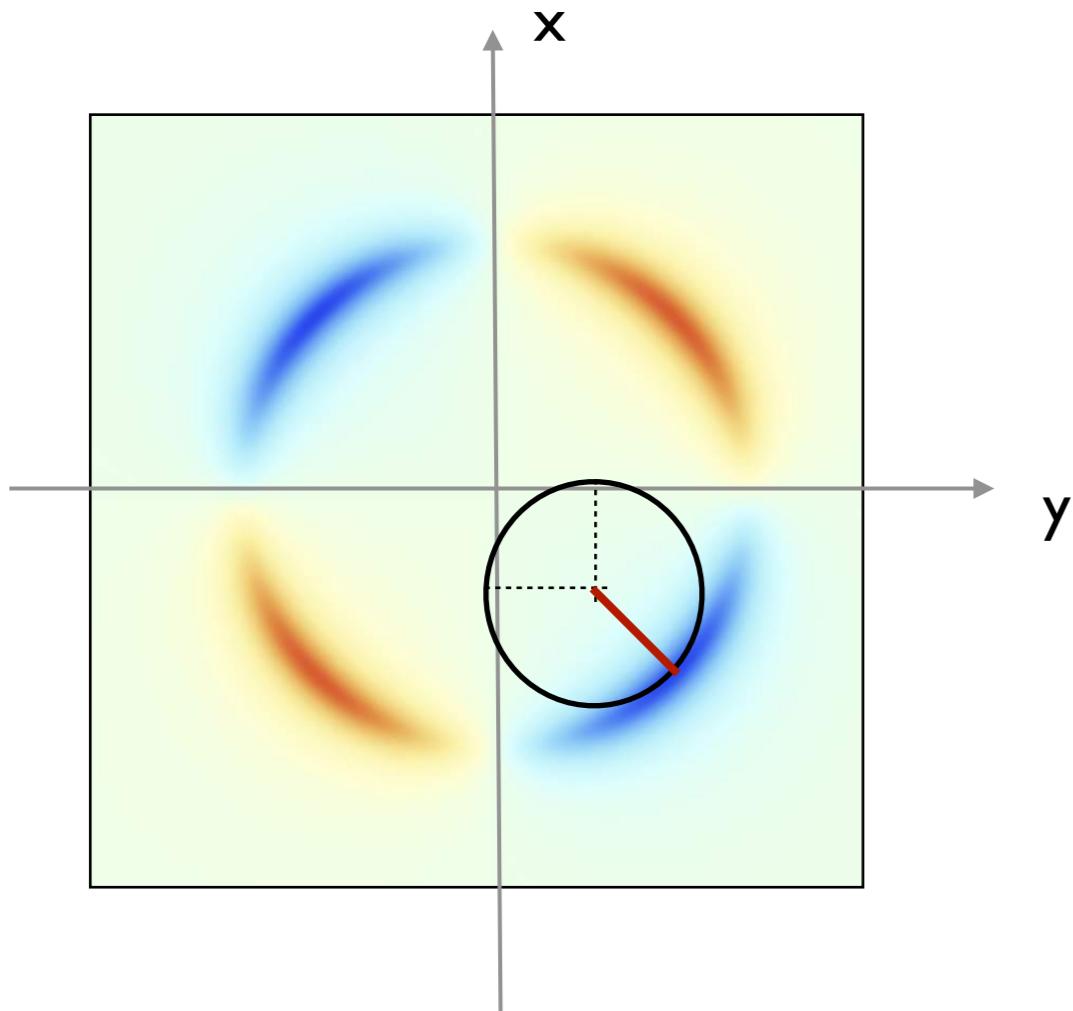
Idealized toy model: The position is fixed and the radius of the halo increases:



Transition mass is correlated with the size of the quadrants.

Mass transition for spin alignment

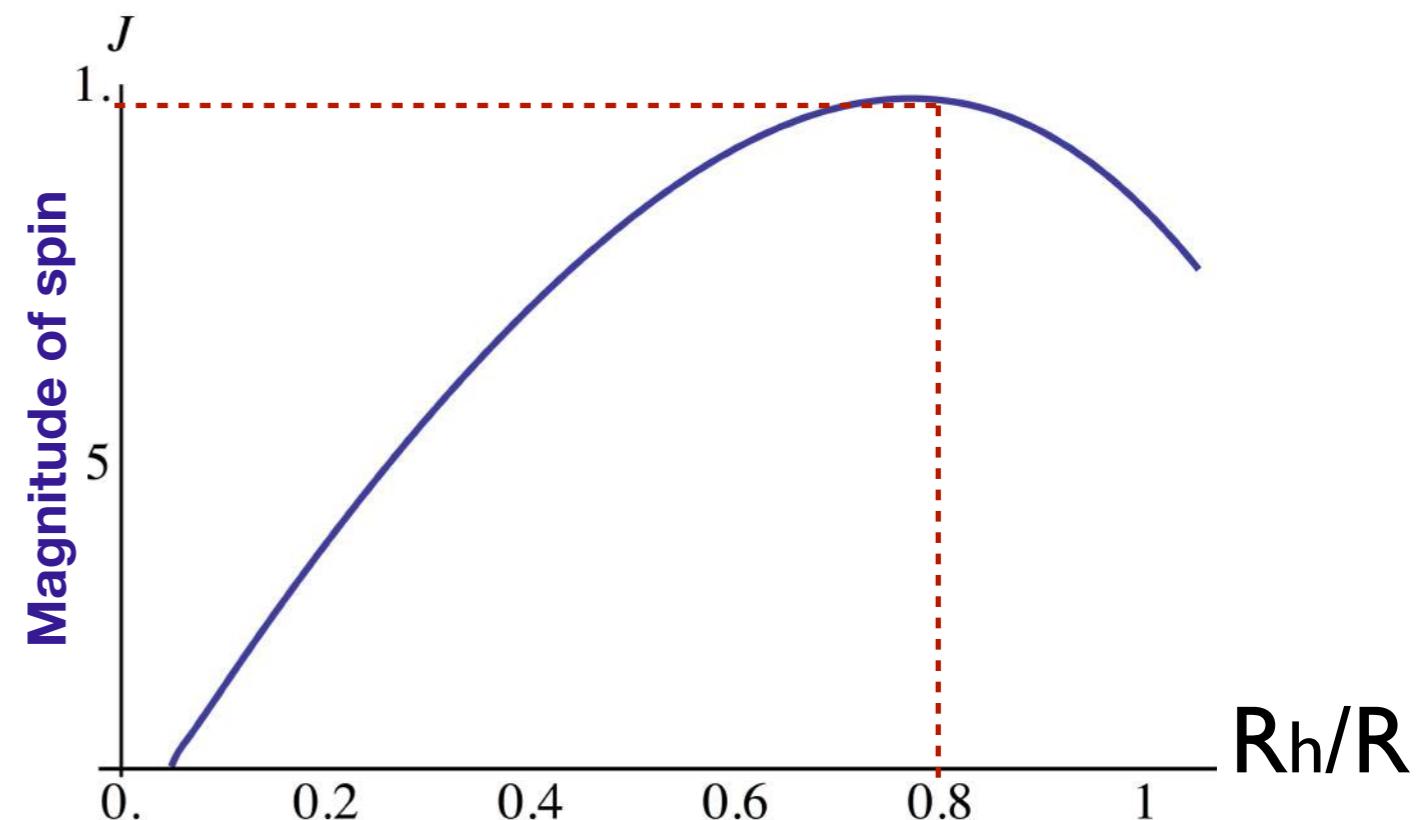
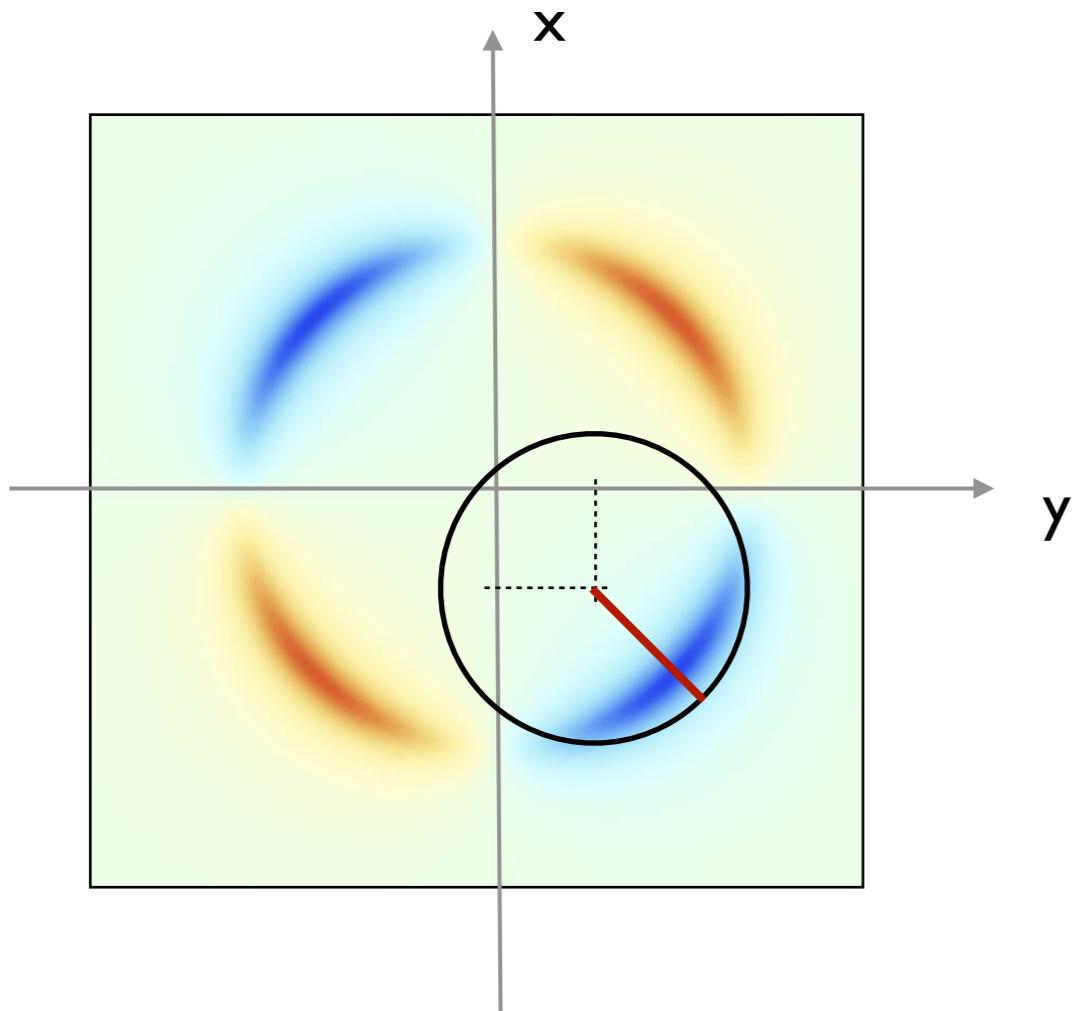
Idealized toy model: The position is fixed and the radius of the halo increases:



Transition mass is correlated with the size of the quadrants.

Mass transition for spin alignment

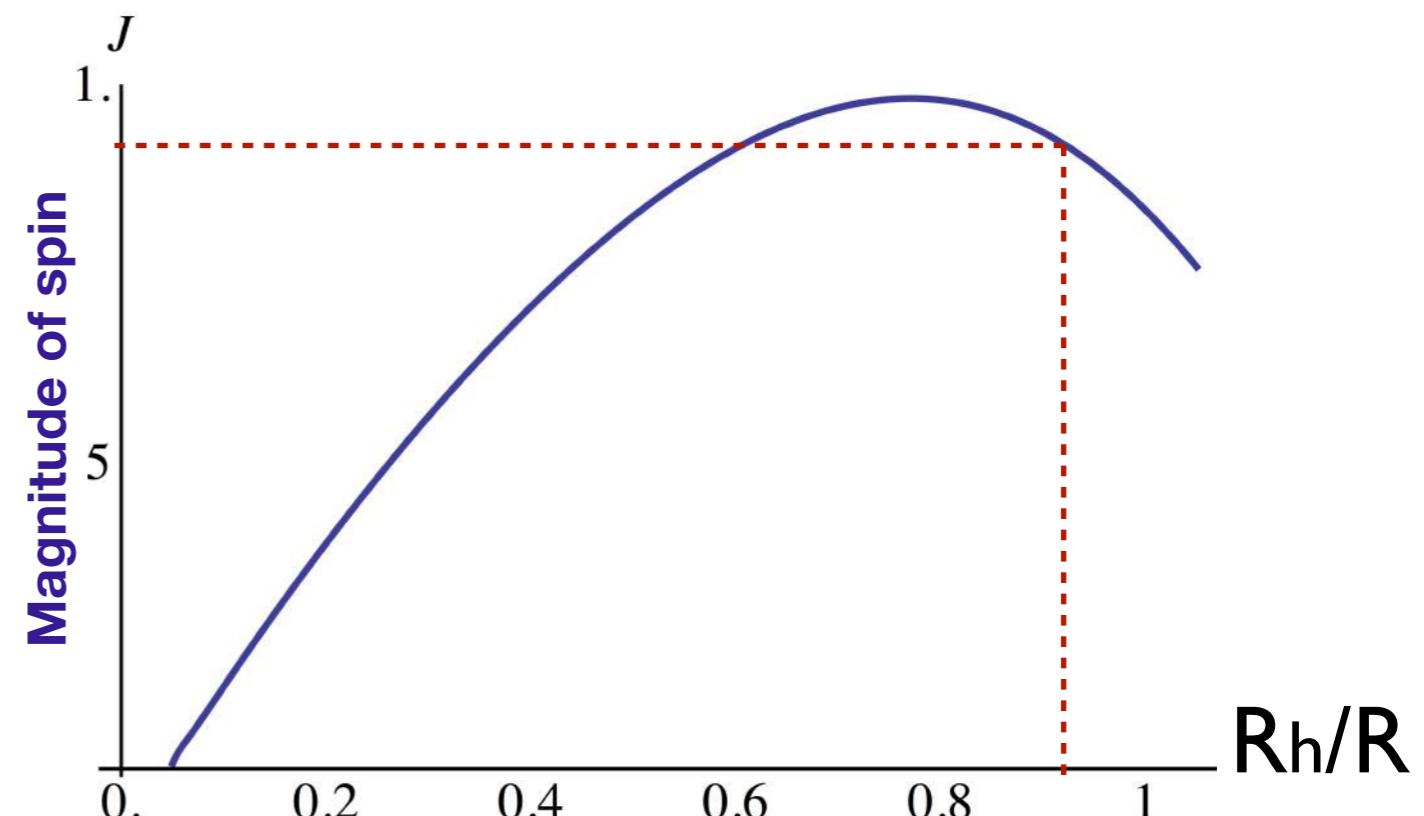
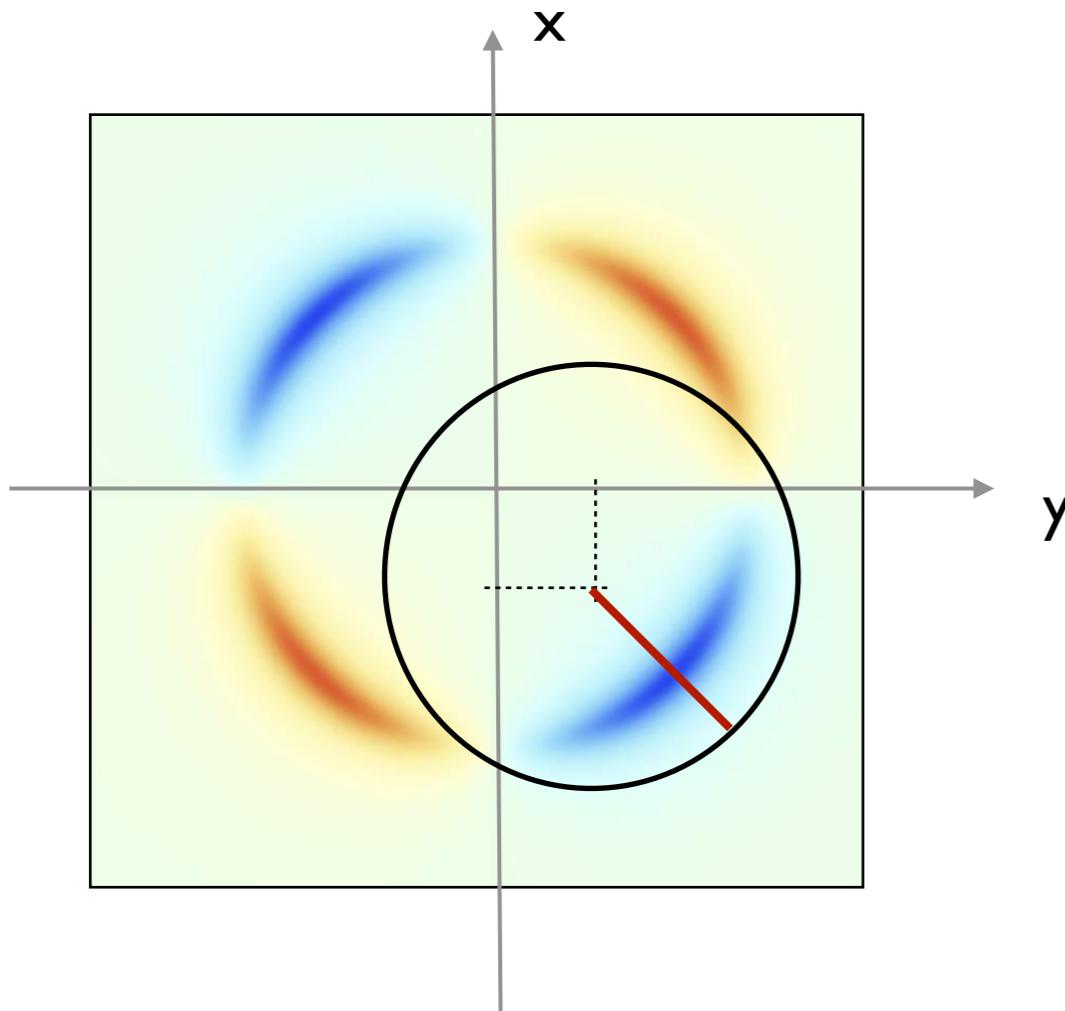
Idealized toy model: The position is fixed and the radius of the halo increases:



Transition mass is correlated with the size of the quadrants.

Mass transition for spin alignment

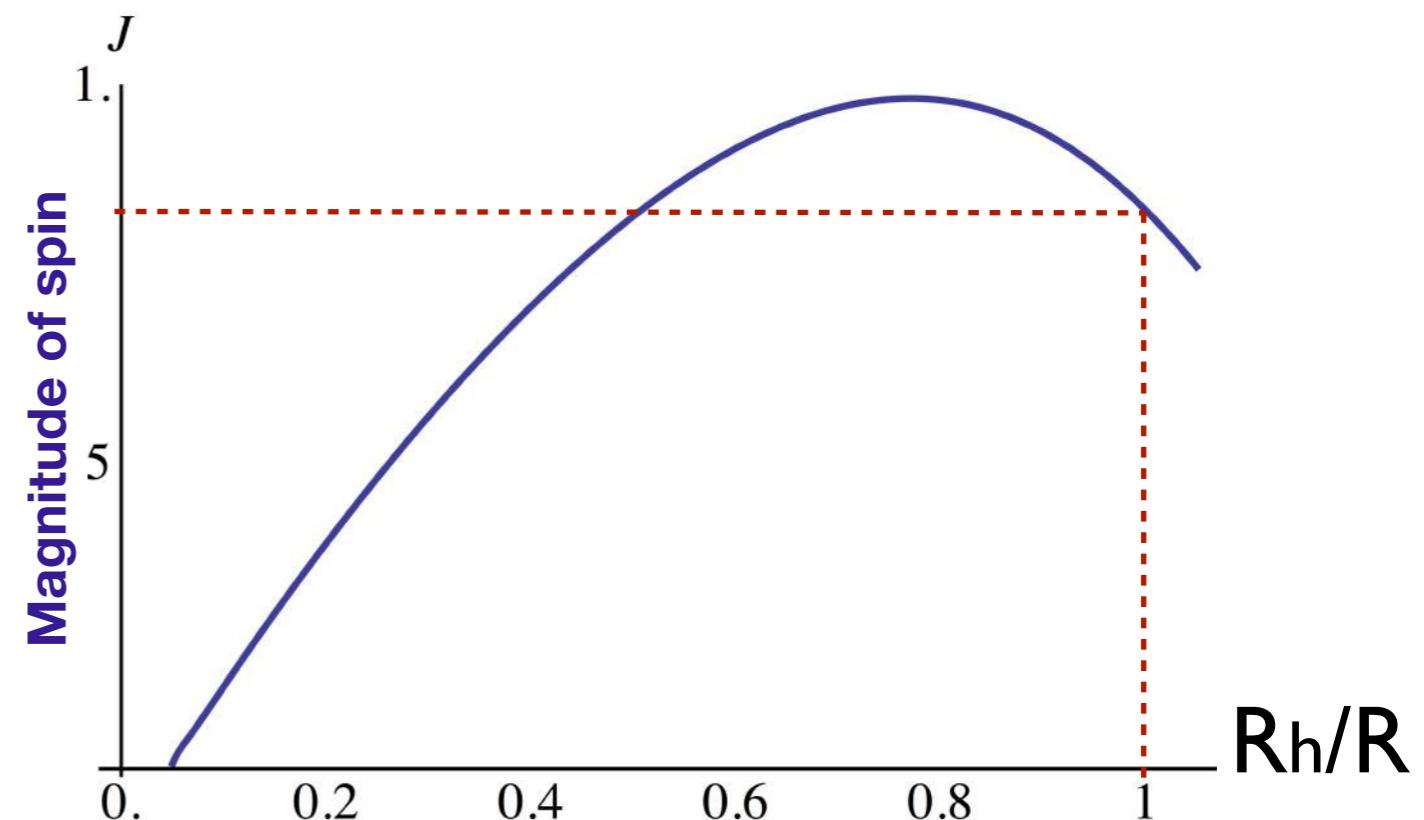
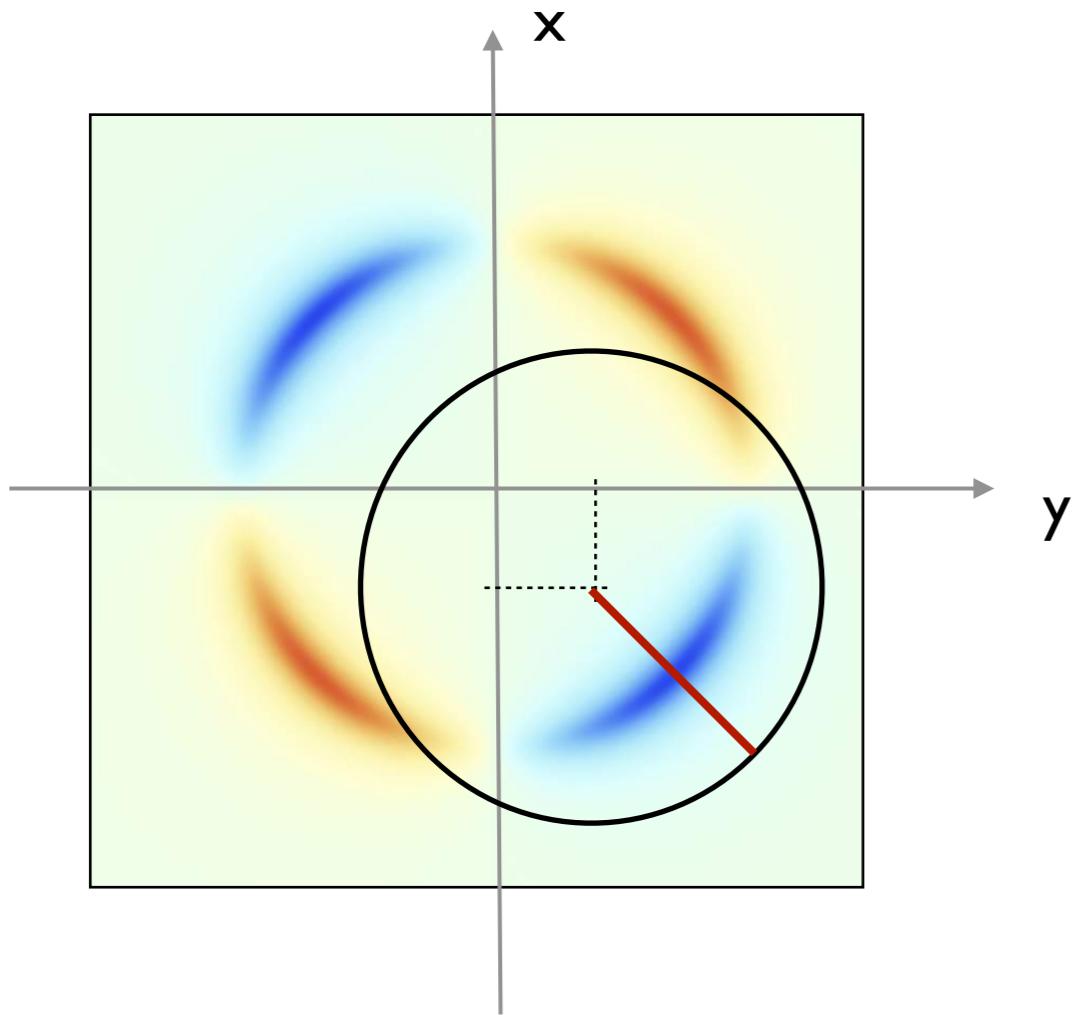
Idealized toy model: The position is fixed and the radius of the halo increases:



Transition mass is correlated with the size of the quadrants.

Mass transition for spin alignment

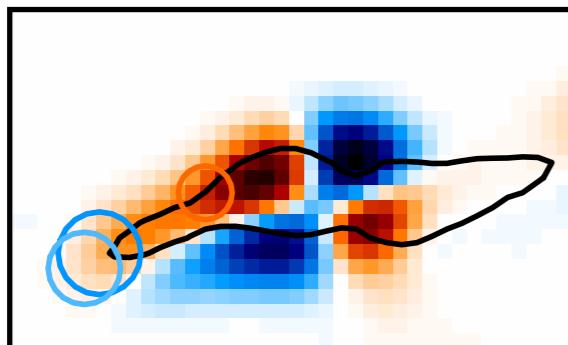
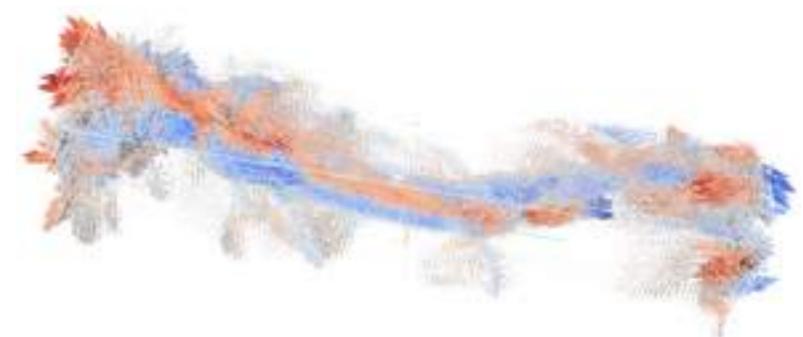
Idealized toy model: The position is fixed and the radius of the halo increases:



Transition mass is correlated with the size of the quadrants.

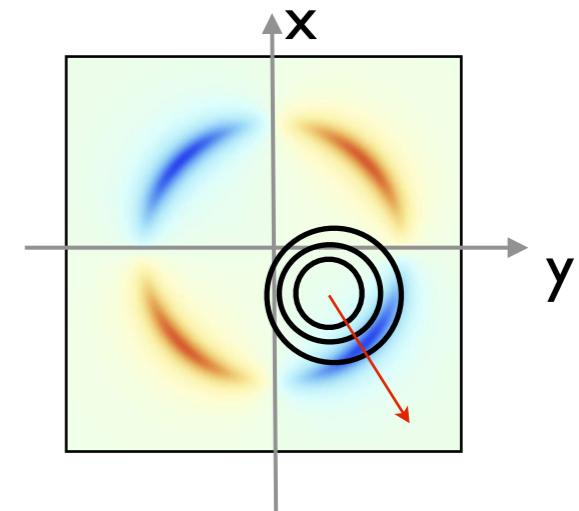
in short...

- ❖ Vorticity is confined in the filaments, and aligned with them. The cross-section with a plane perpendicular to the filament is typically quadripolar.



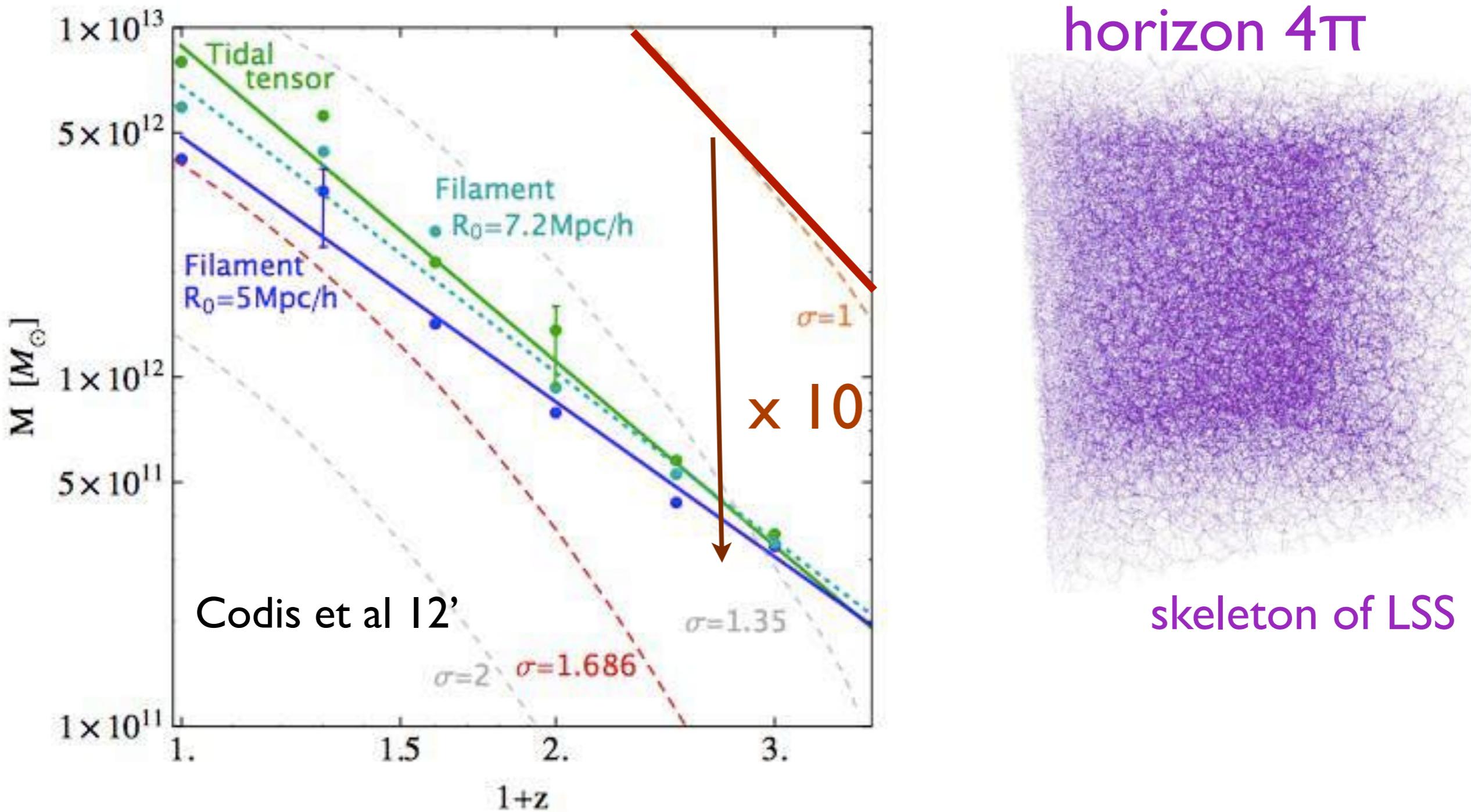
- ❖ Halo spins are aligned with the same polarity than vorticity in quadrants.

- ❖ Qualitatively, the transition mass in the alignment could be correlated with the size of the quadrant.



Explain transition mass? YES!

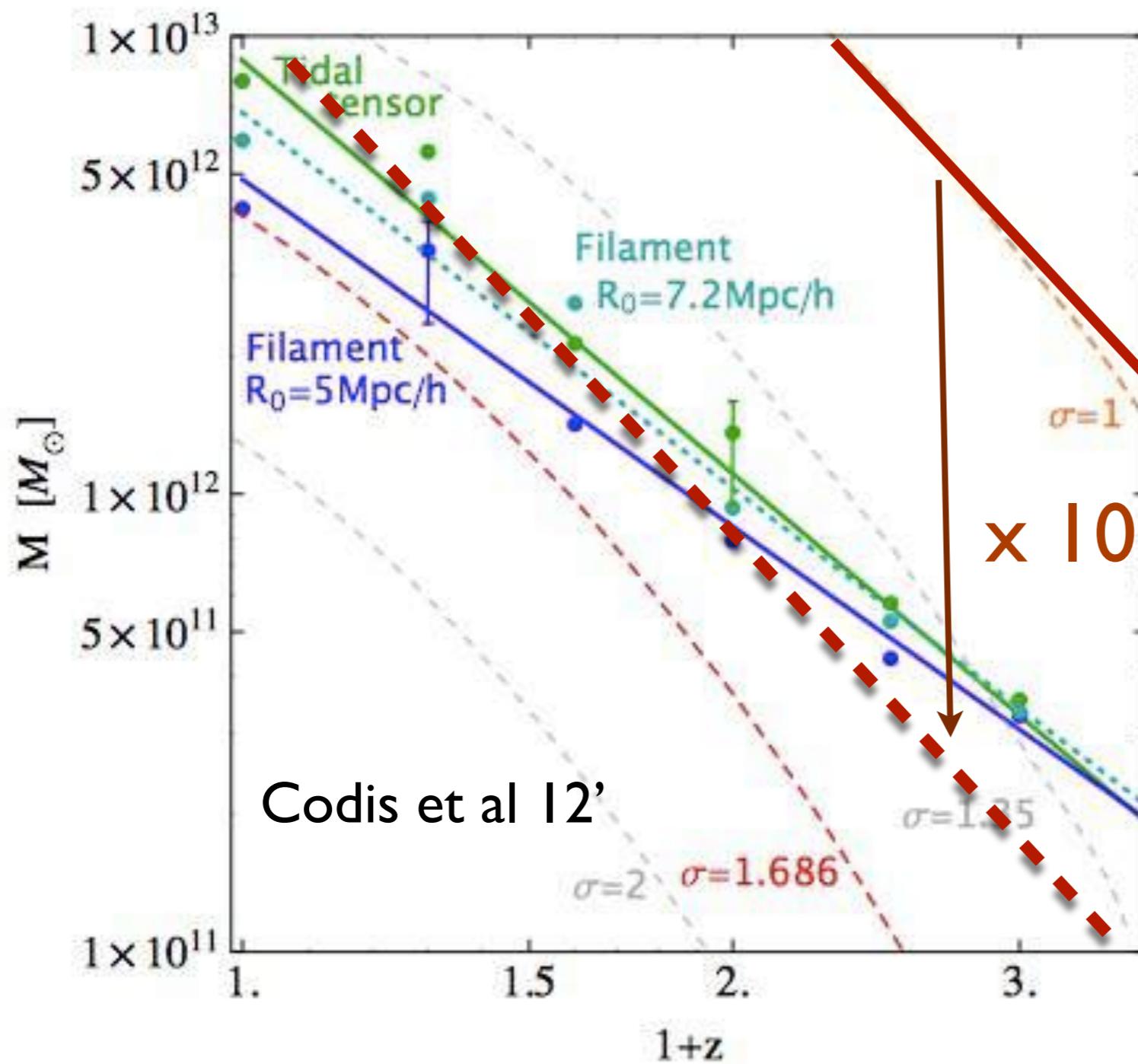
Transition mass versus redshift



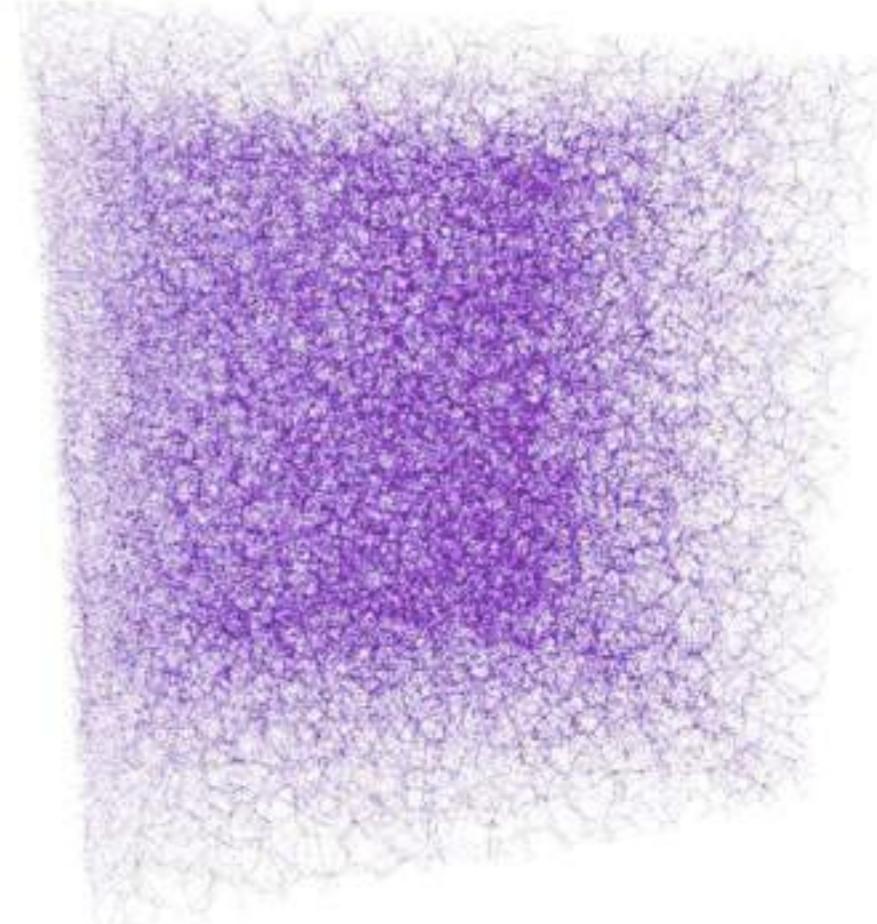
Only 2 **ingredients**: a) spin is spin one b) filaments flattened

Explain transition mass? YES!

Transition mass versus redshift

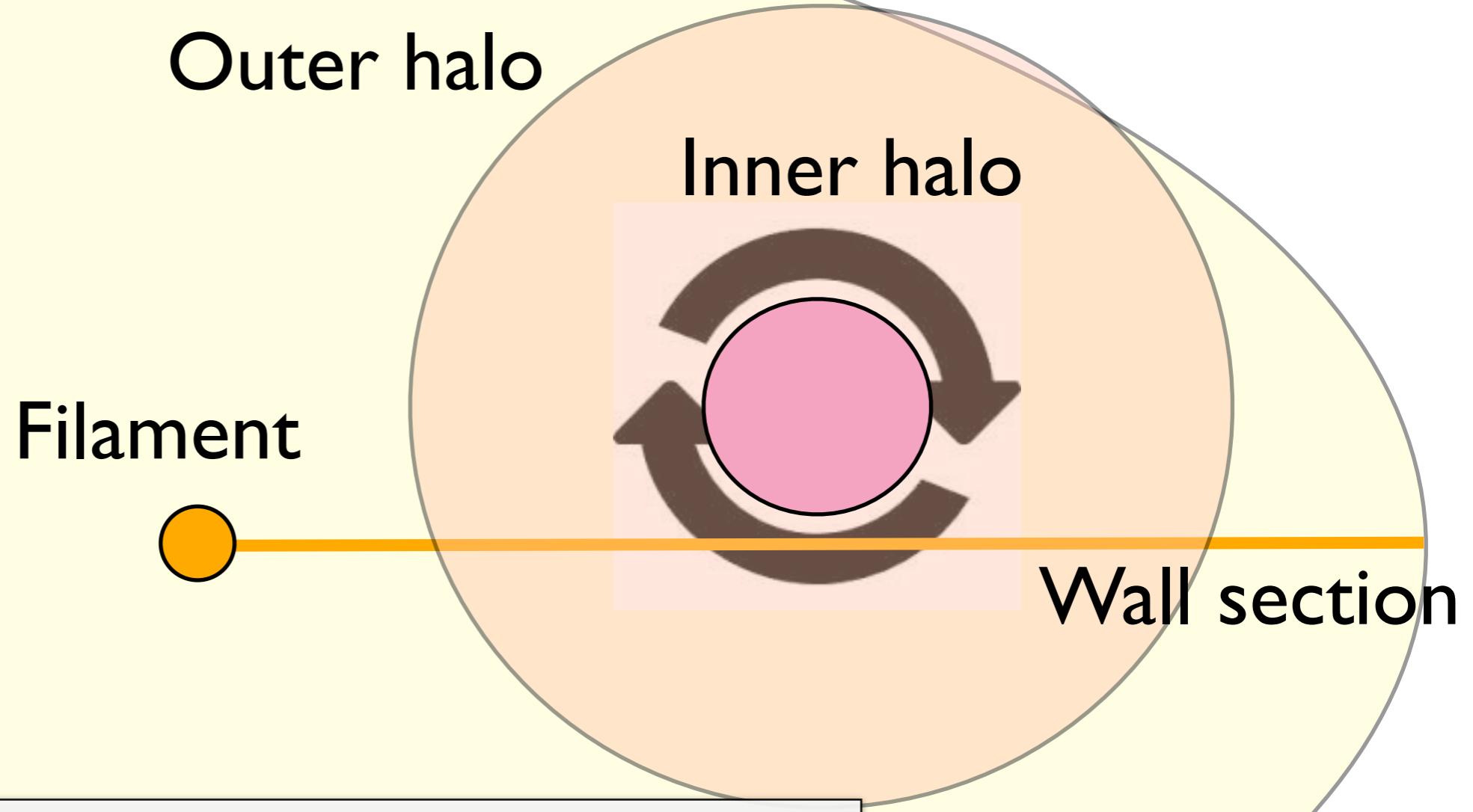


horizon 4π

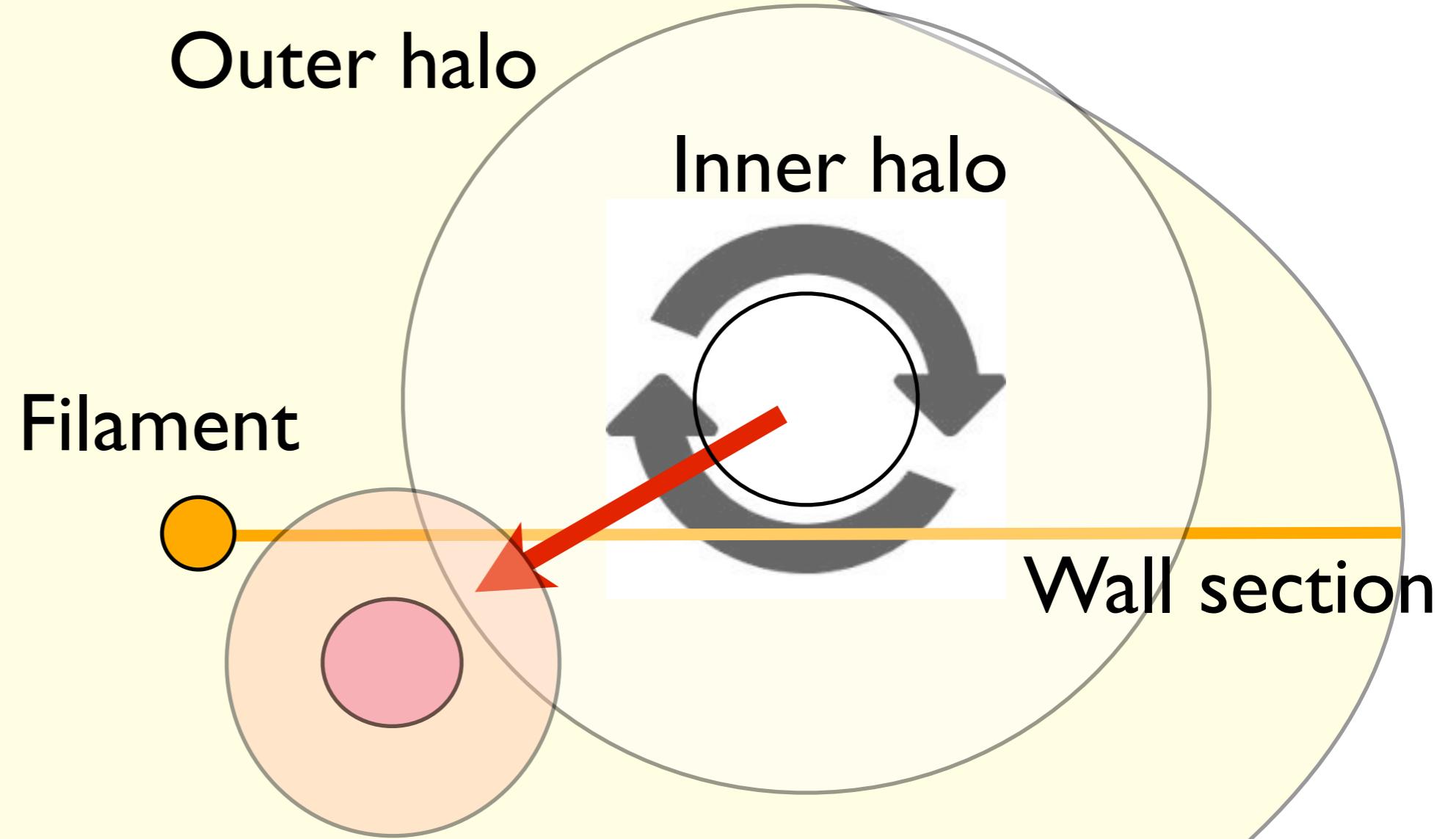


skeleton of LSS

Only 2 ingredients: a) spin is spin one b) filaments flattened

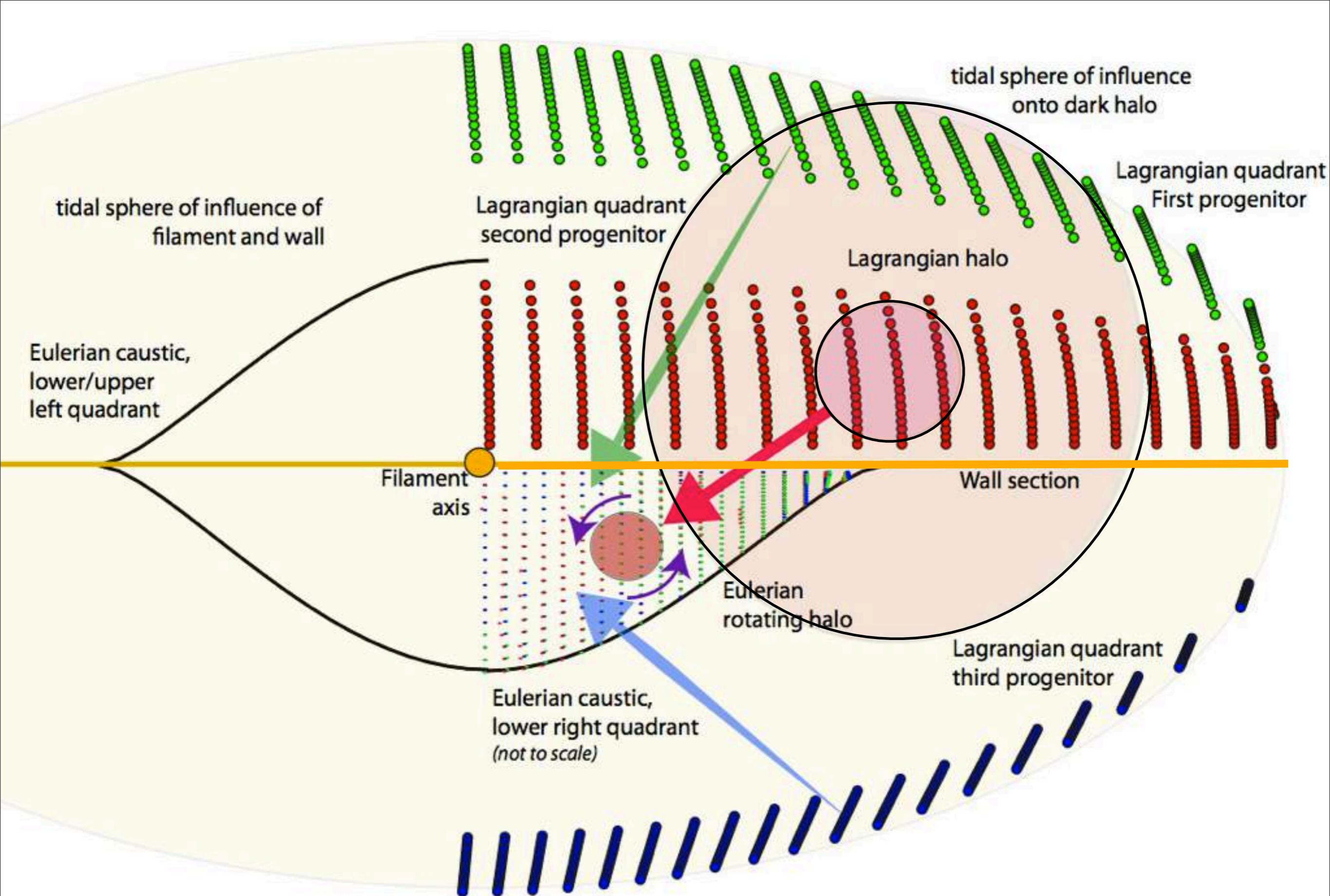


Connecting Eulerian &
Lagrangian theories



+Zeldovitch boost

BUT in // from the pt of view of LSS



Complementary vorticity advection view

Take home message...

- Morphology (= AM stratification) driven by LSS in cosmic web: it explains Es & Sps where, how & why from **ICs**
- Signature in correlation between *spin* and *internal kinematic* structure of cosmic web on larger scales.
- Process driven by simple PBS/biassed clustering dynamics:
 - requires updating TTT to **saddles**: simple theory :-)
 - can be expressed into an Eulerian theory via vorticity

Where galaxies form does matter, and can be traced back to ICs

*Flattened filaments generate point-reflection-symmetric AM/vorticity distribution:
they induce the observed spin transition mass*

- which is why the δ -Web is the best :-)

What about galaxies ??

- Horizon-AGN simulation Jade (CINES)
(PI Y. Dubois, Co-I J. Devriendt & C. Pichon)
 - $L_{\text{box}}=100 \text{ Mpc}/h$
 - 1024^3 DM particles $M_{\text{DM, res}}=8 \times 10^7 M_{\text{sun}}$
 - Finest cell resolution $dx=1 \text{ kpc}$
 - Gas cooling & UV background heating
 - Low efficiency star formation
 - Stellar winds + SNII + SNIa
 - O, Fe, C, N, Si, Mg, H
 - AGN feedback radio/quasar
- Outputs
(backed up and analyzed on BEYOND)
 - Simulation outputs
 - Lightcones ($1^\circ \times 1^\circ$) performed on-the-fly
 - Dark Matter (position, velocity)
 - Gas (position, density, velocity, pressure, chemistry)
 - Stars (position, mass, velocity, age, chemistry)
 - Black holes (position, mass, velocity, accretion rate)
- $z=1.5$ using 3 Mhours on 4096 cores

PART IV

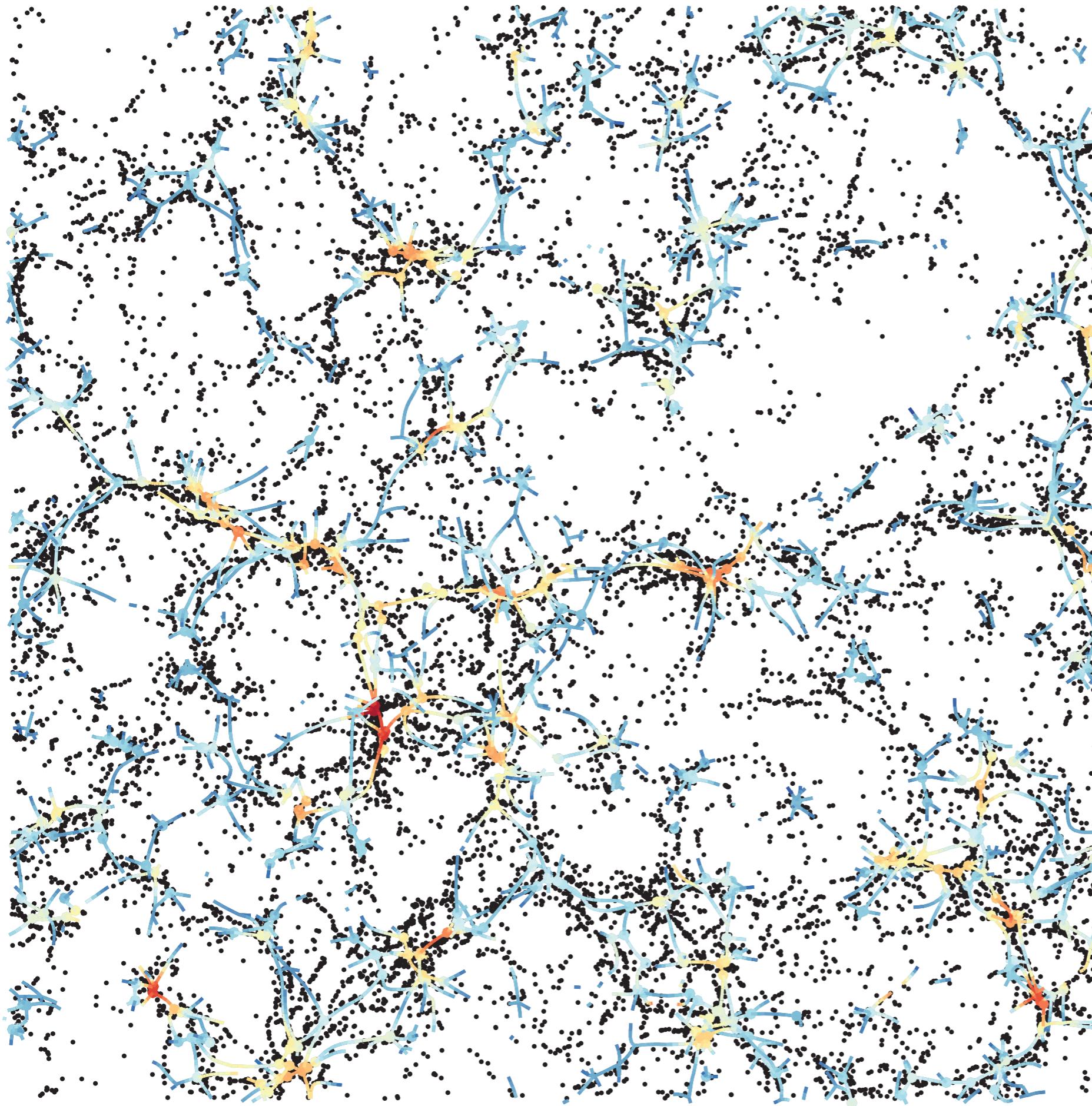
horizon-AGN.projet-horizon.fr

$z=1.2$

Part V Outline

- ⦿ Can morphology/physics trace spin flip?
- ⦿ Are transition masses consistent?
- ⦿ The fate of forming galaxies
- ⦿ The fate of merging galaxies

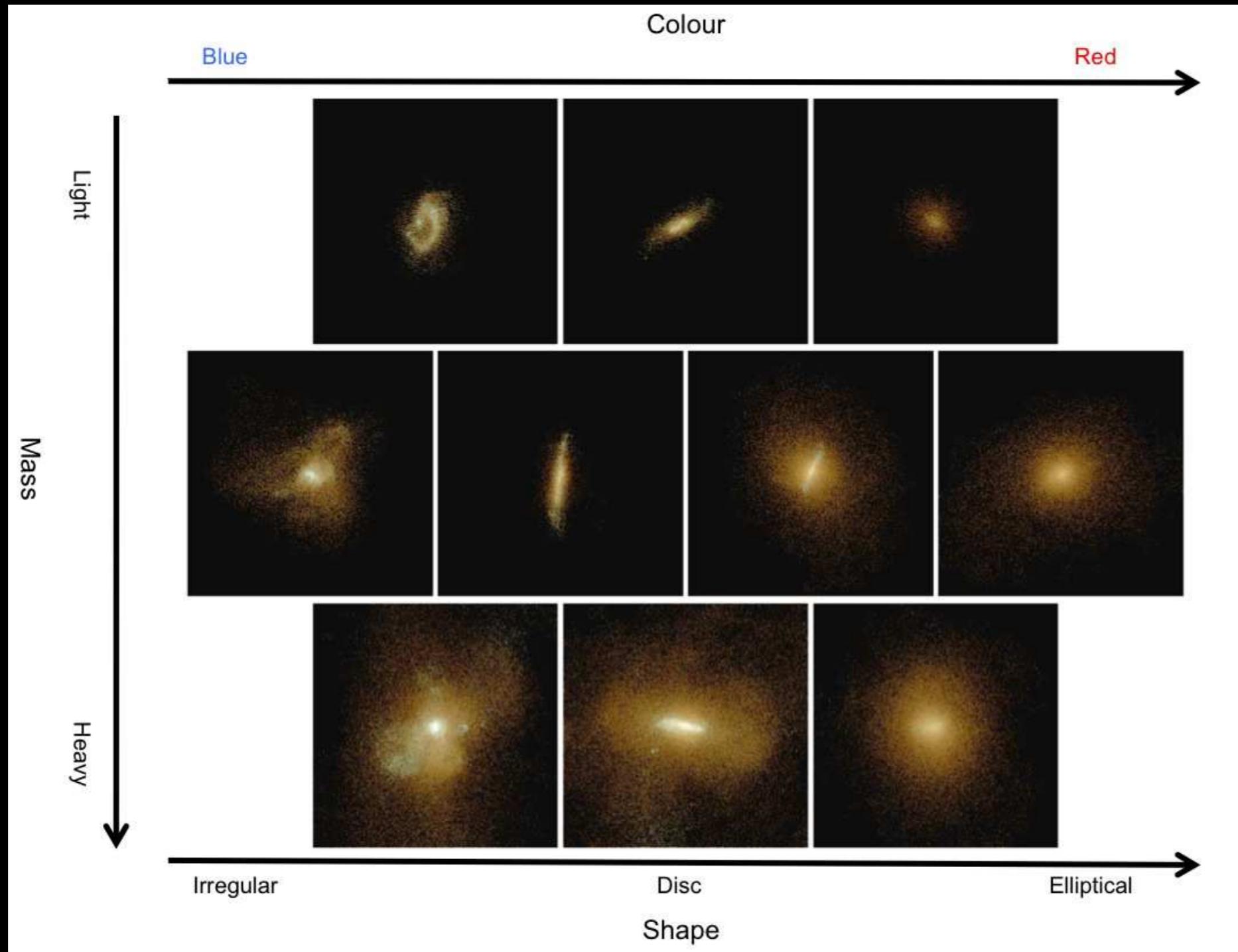
Galaxies versus dense filaments



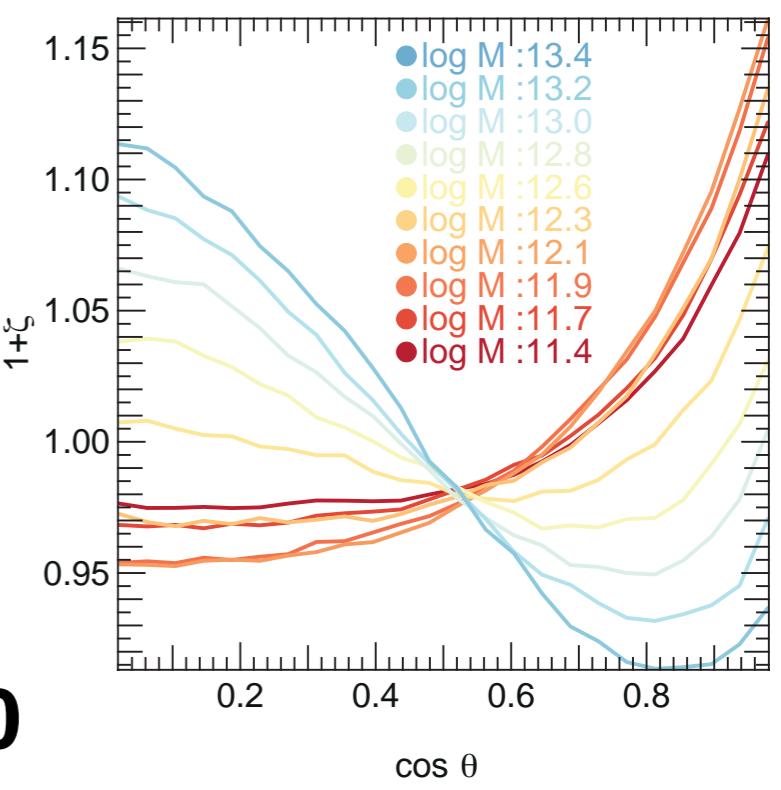
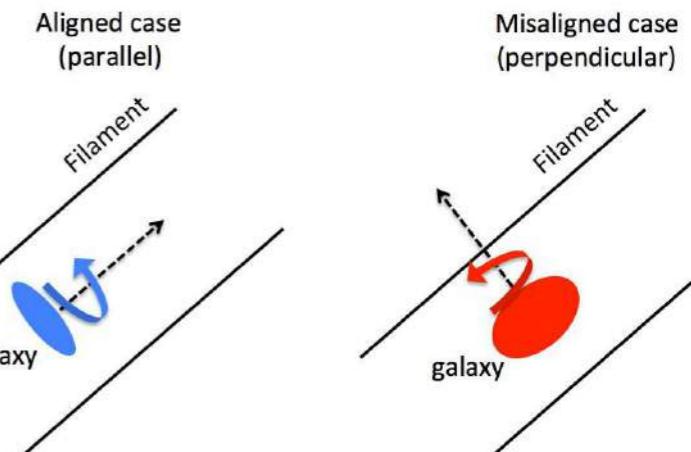
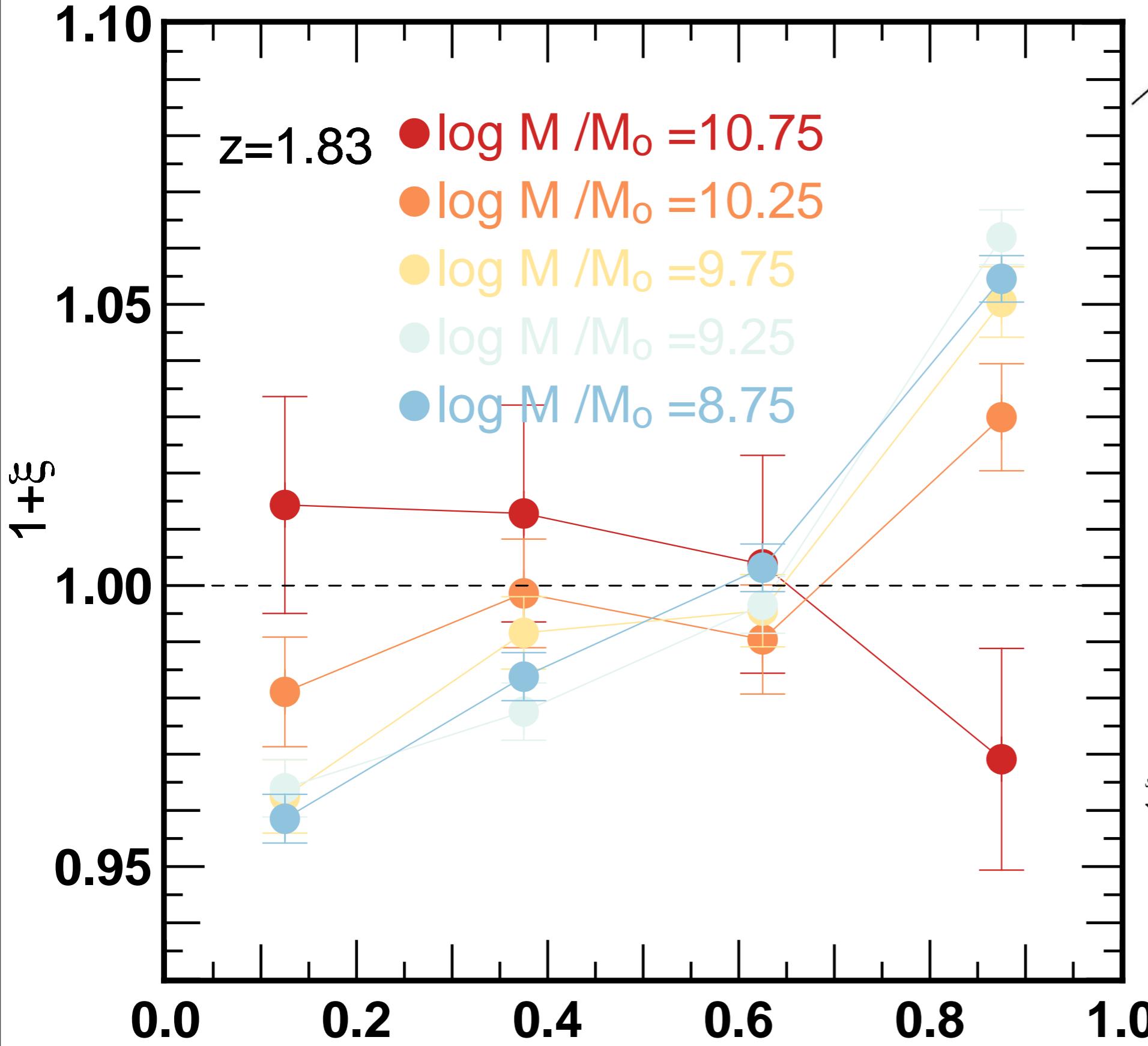
Galaxies are strongly
clustered
near filaments

can morphology trace spin flip ?

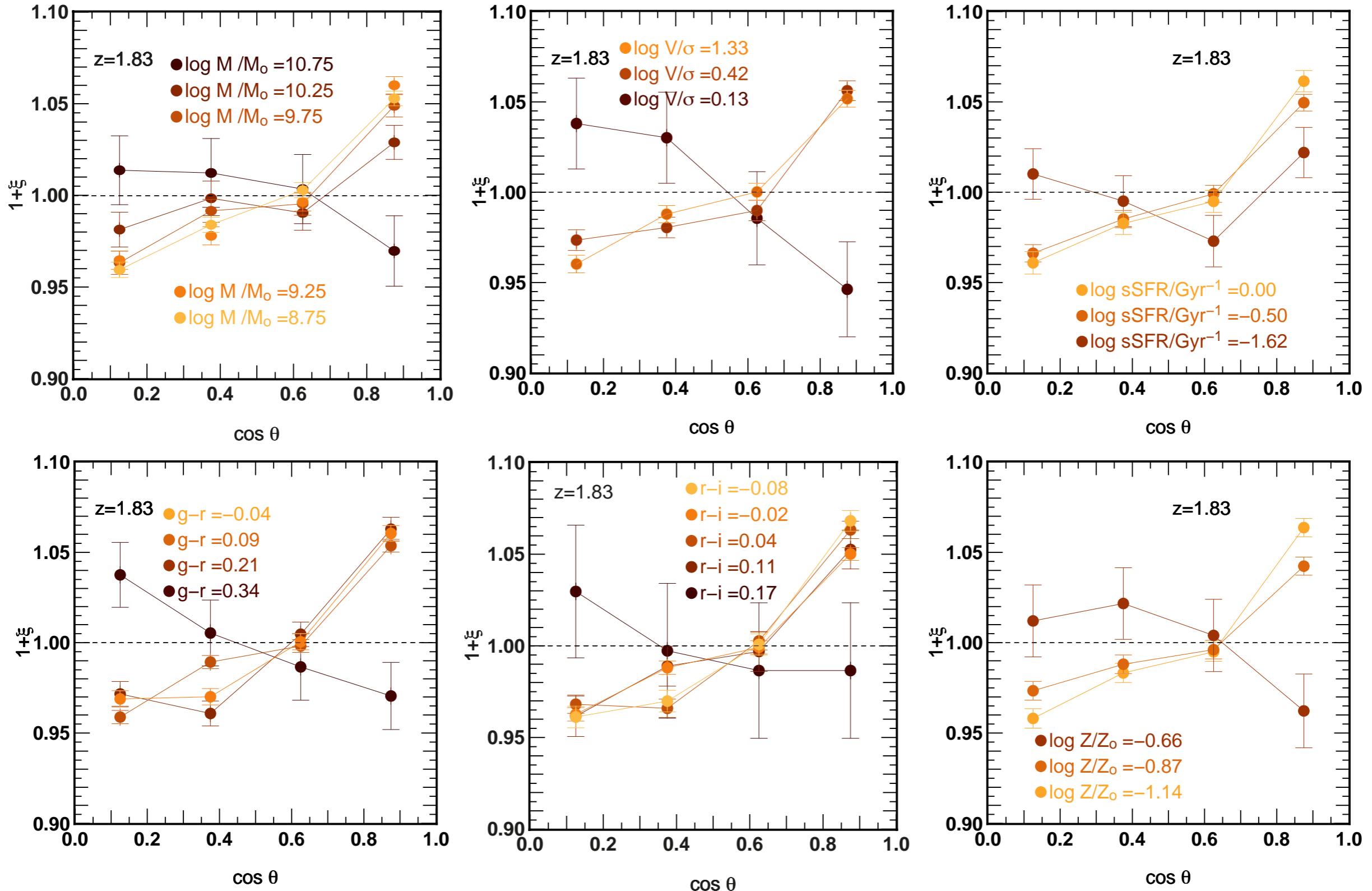
- thanks to AGN feedback we have morphological diversity



Filament-galactic spin & mass



Can morphological/physical properties of galaxies trace spin flip?

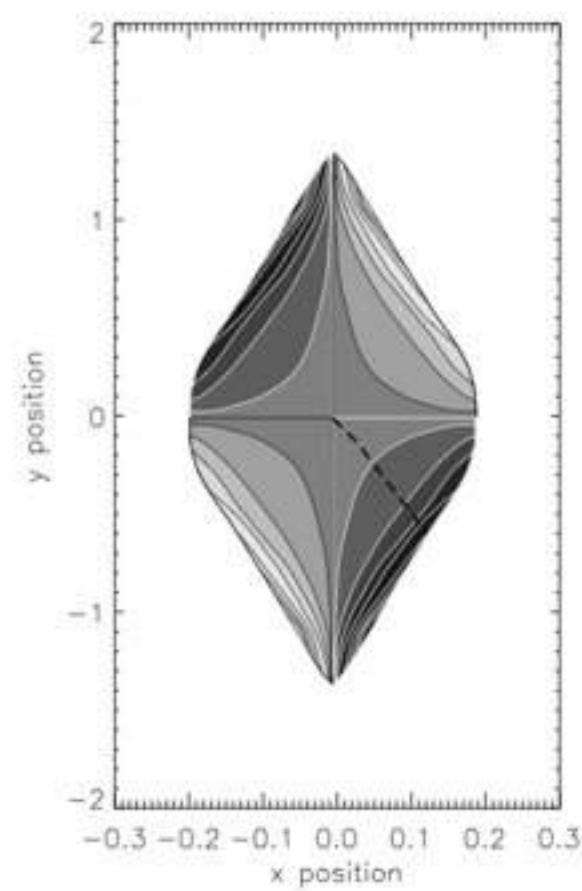
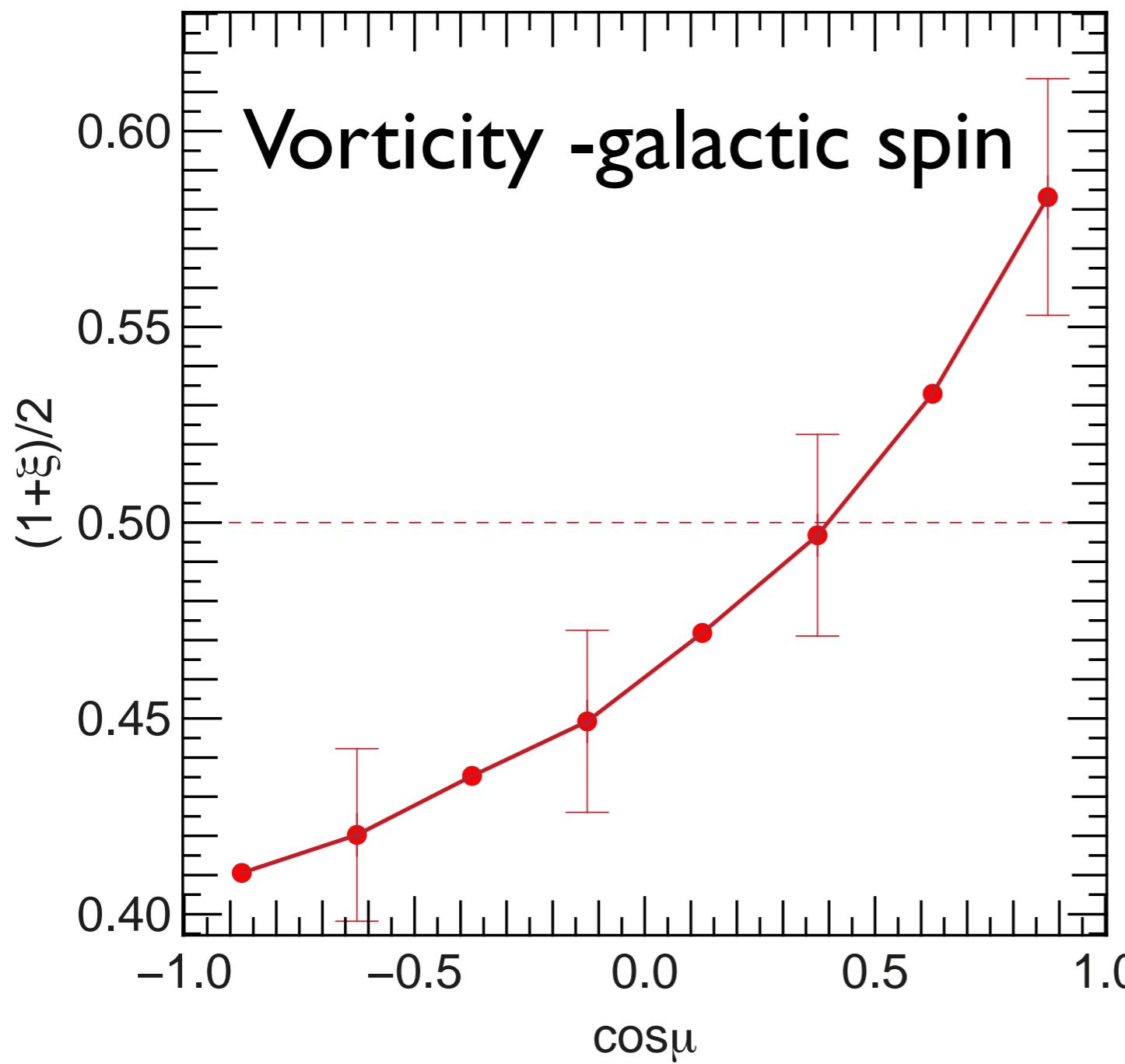


is morphometric
transition mass
consistent
with DM ?

Final point 1/2:
low mass galaxies

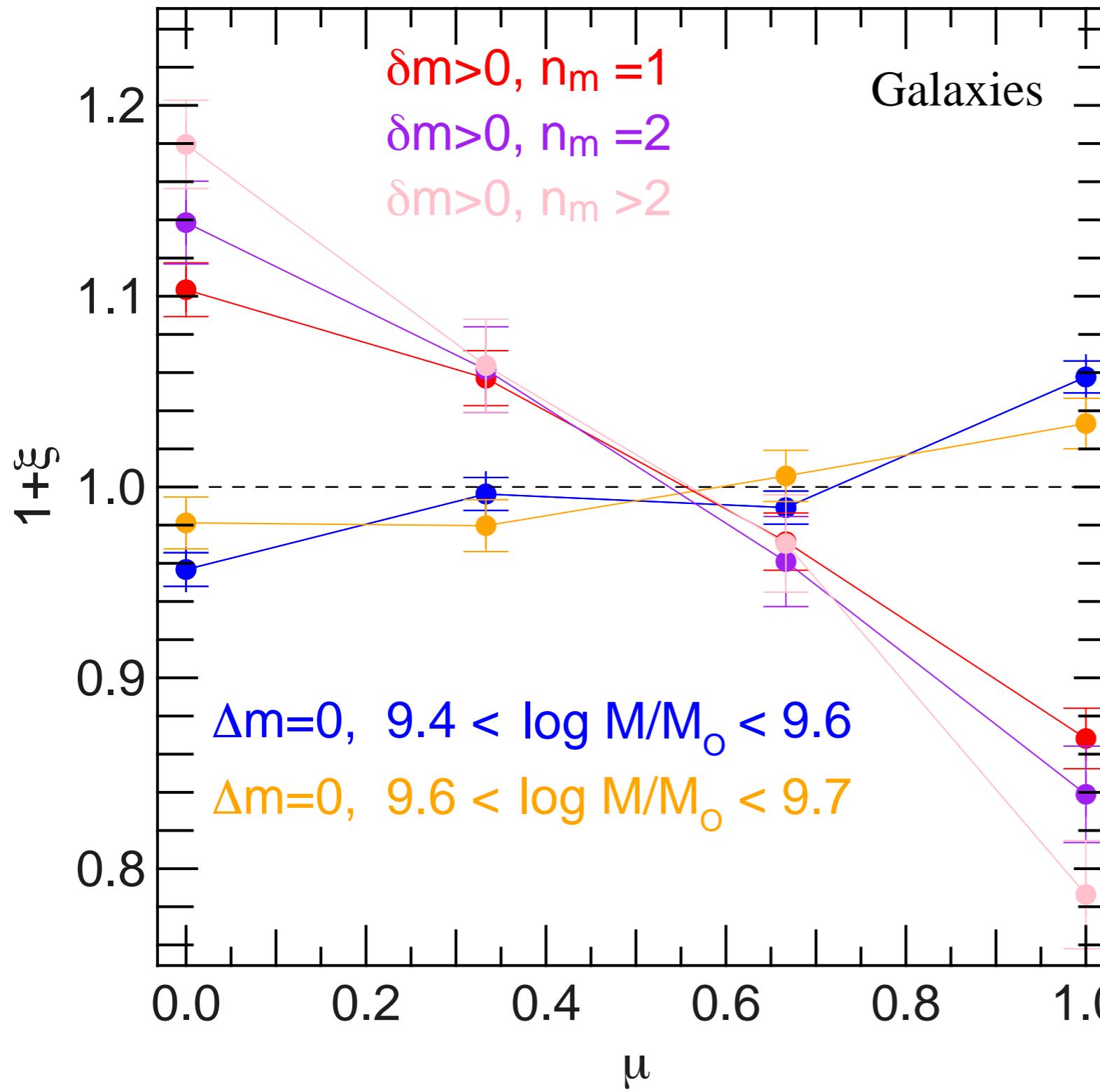
What is the *physical* origin of low mass galaxies spin-filament alignment ?

Vorticity arising from kin. structure of filament!



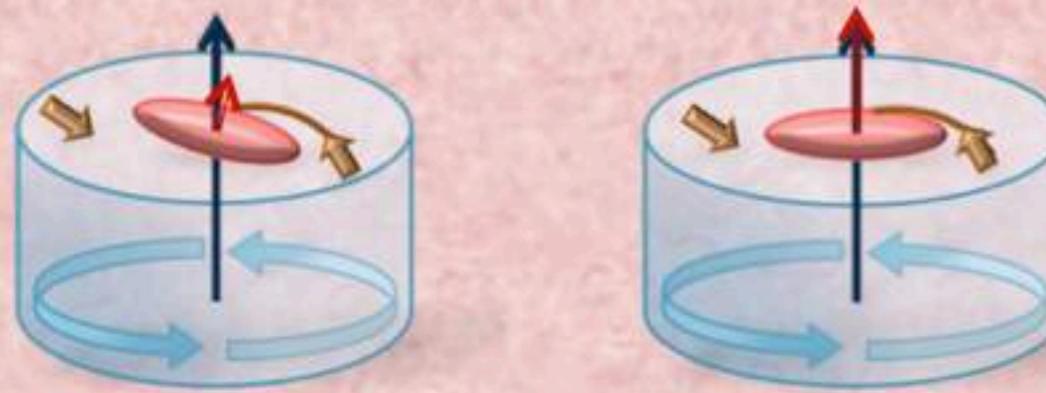
Final point 2/2:
high mass galaxies

What is the *physical* origin of spin flip? high mass galaxies merge!

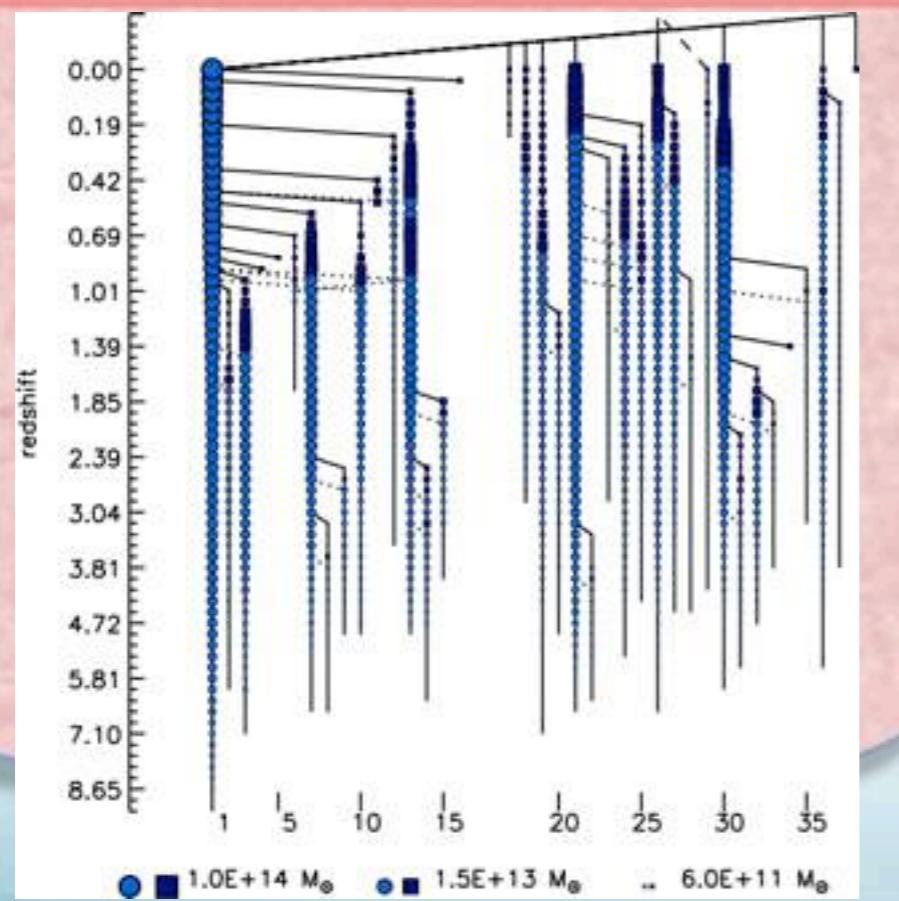


Transition mass
versus **merging
rate**
for galaxies

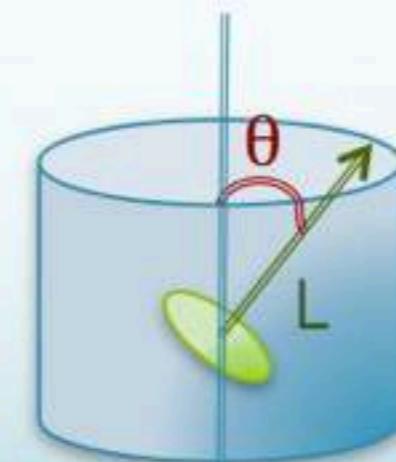
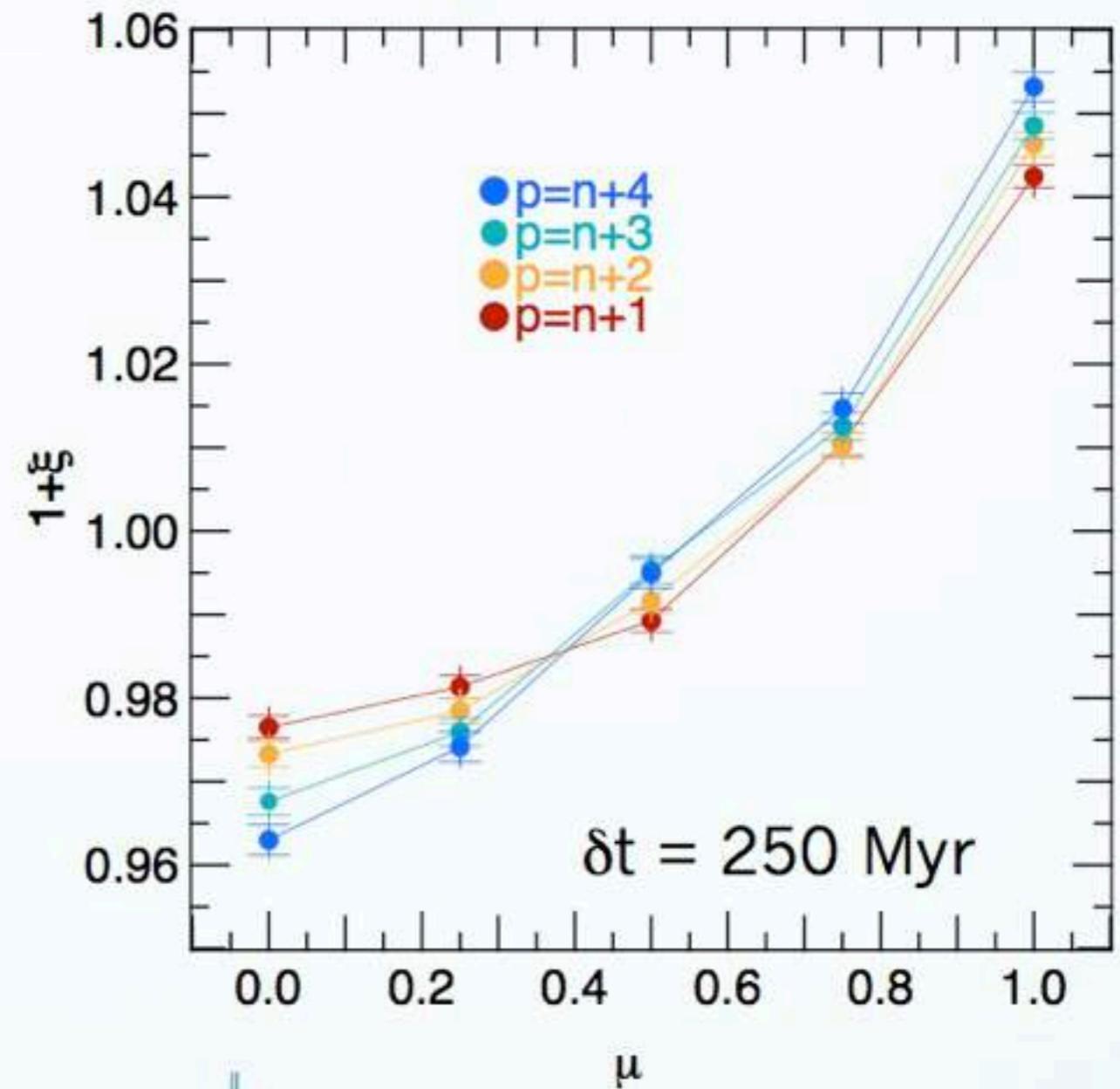
SMOOTH ACCRETION



- Gas inflows (re)-align galaxies with their filament



PDF of μ over 4 timesteps δt



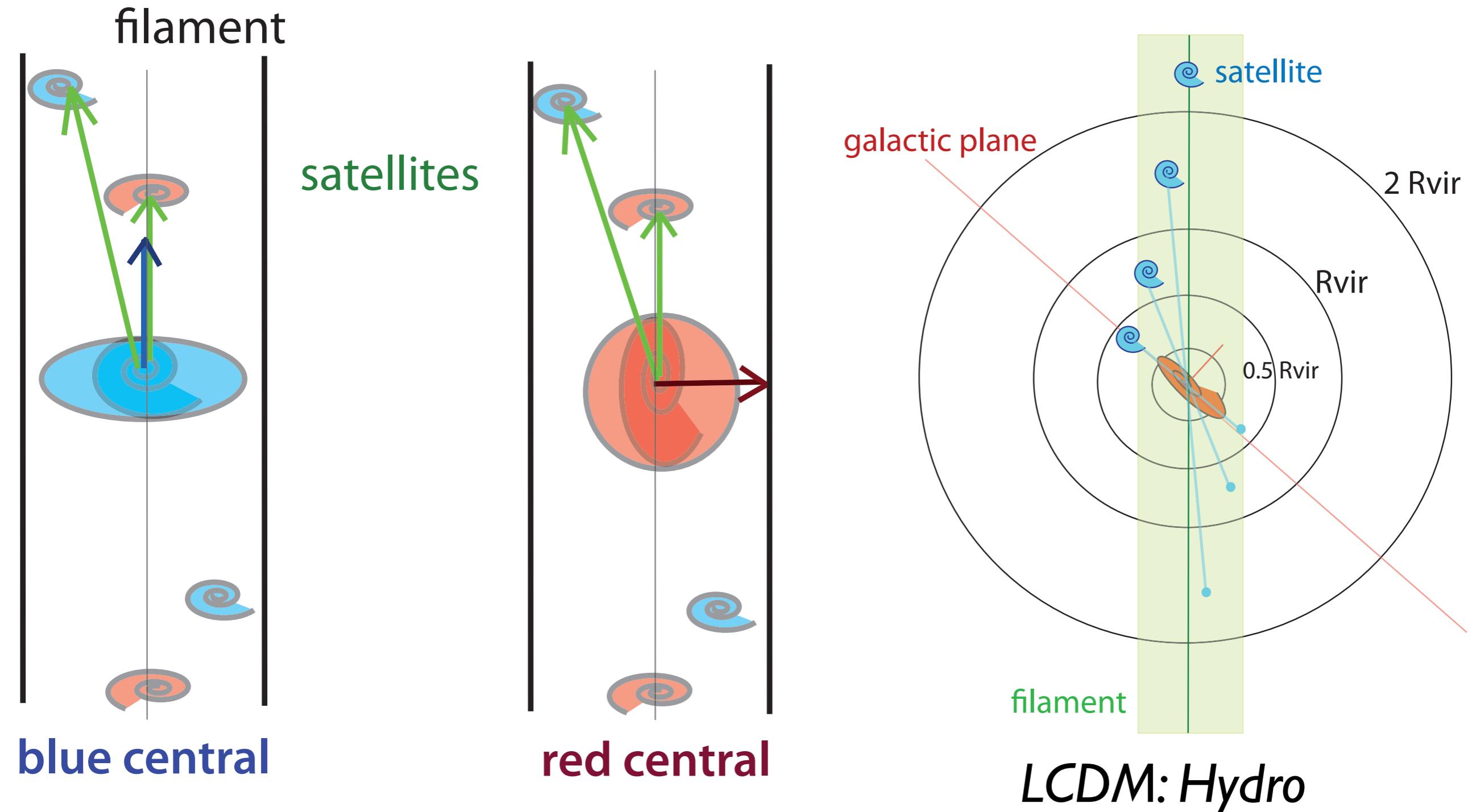
$$\mu = |\cos(\theta)|$$

- Spin-filament angle
- ξ : excess probability

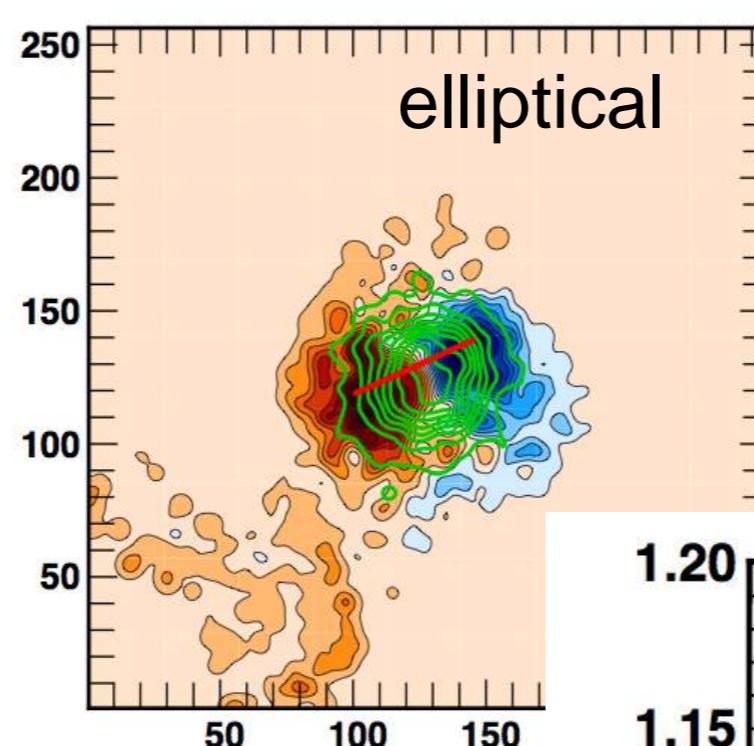
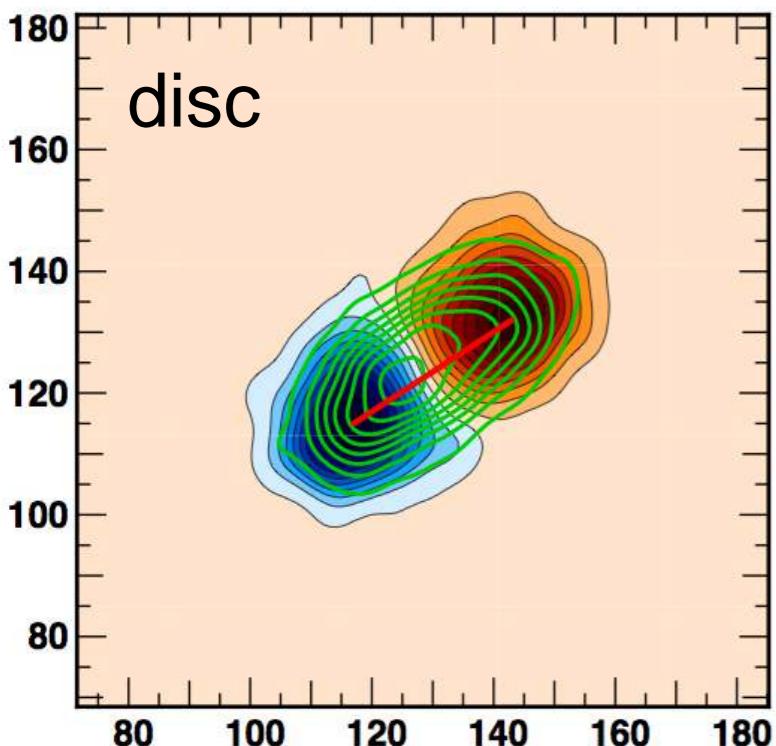
no merger : orientation versus look-back time

Caught in the rhythm: satellites in their galactic plane

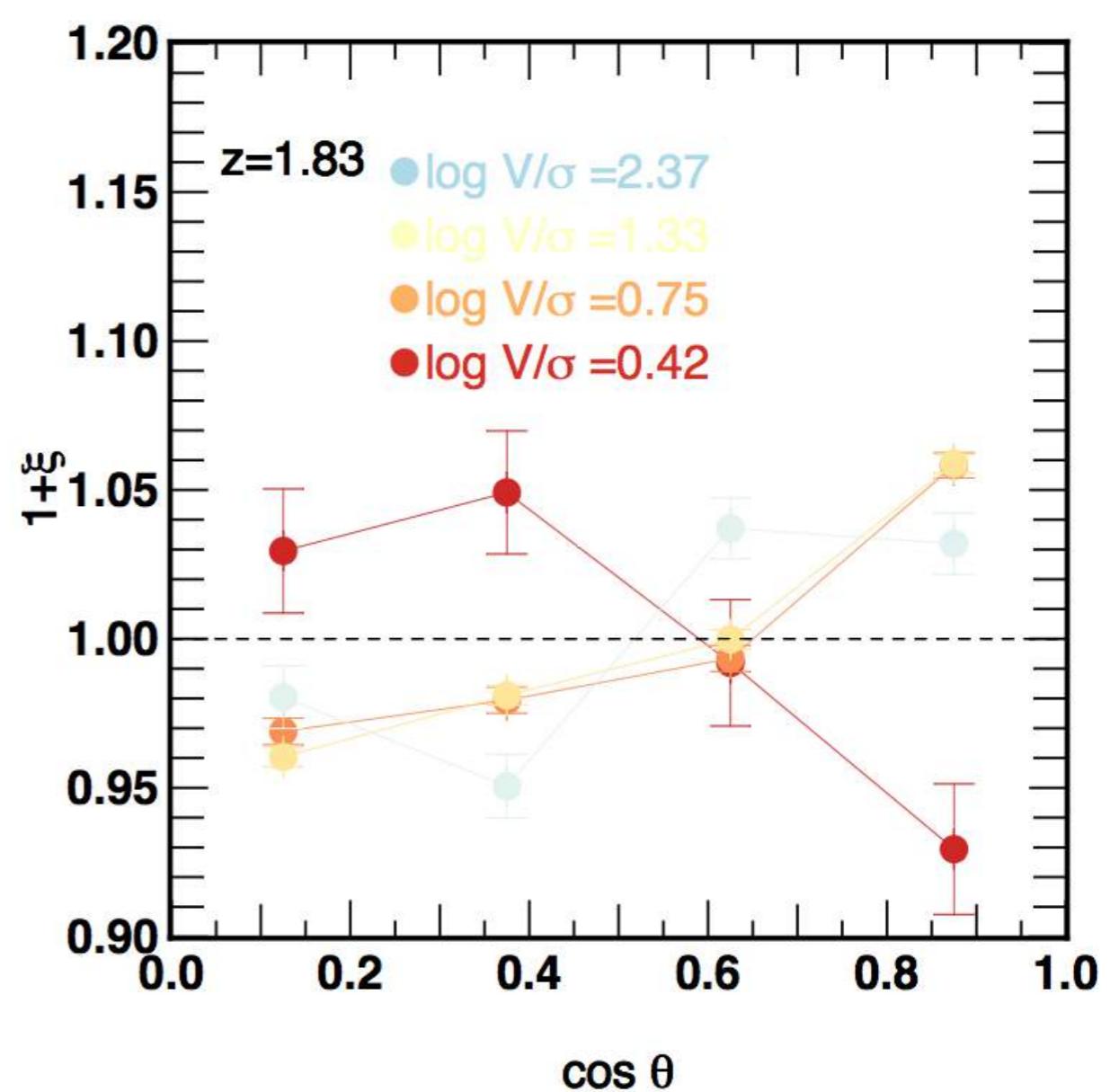
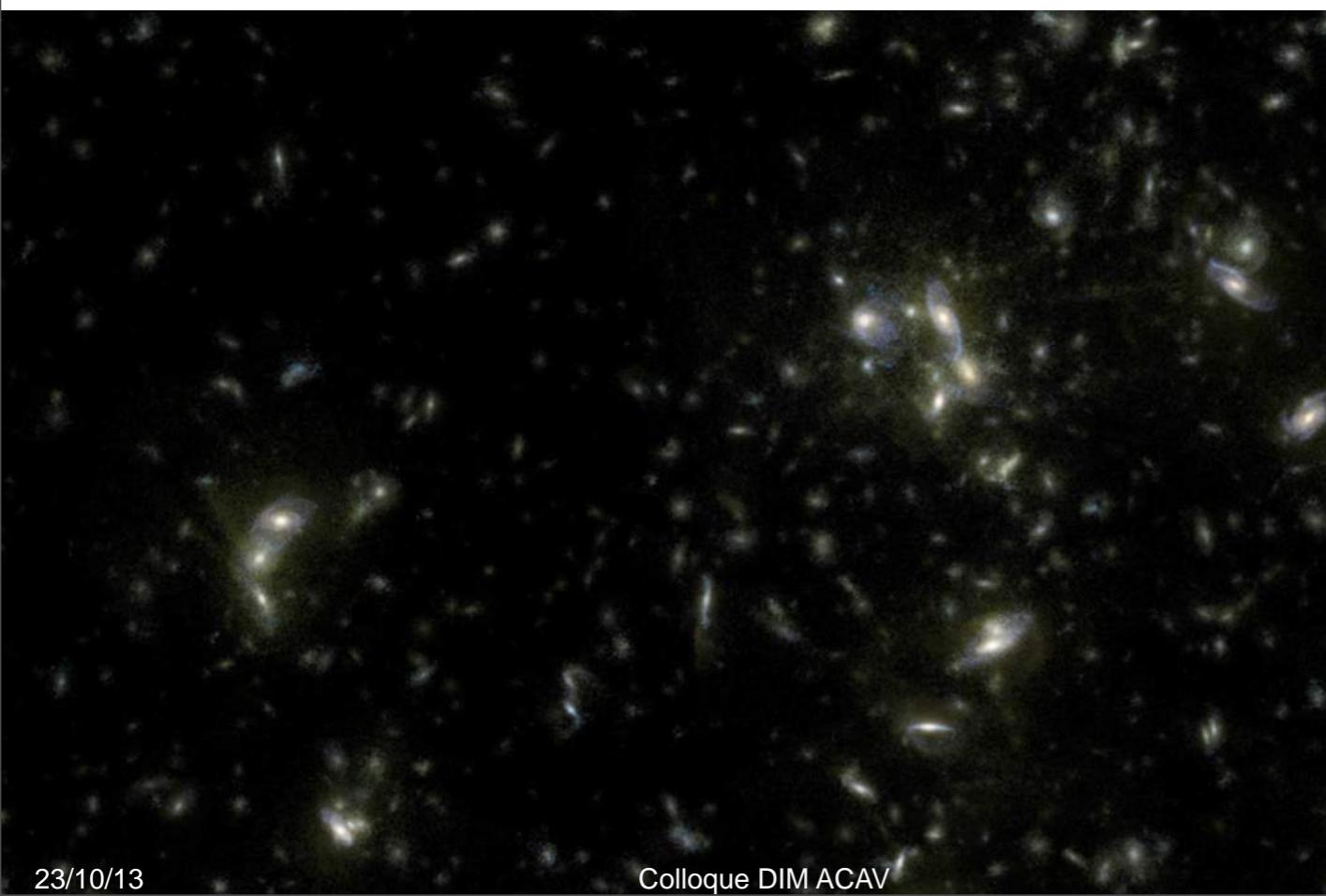
C. Welker^{1*}, Y. Dubois¹, C. Pichon^{1,2}, J. Devriendt^{3,4} and N. E. Chisari³



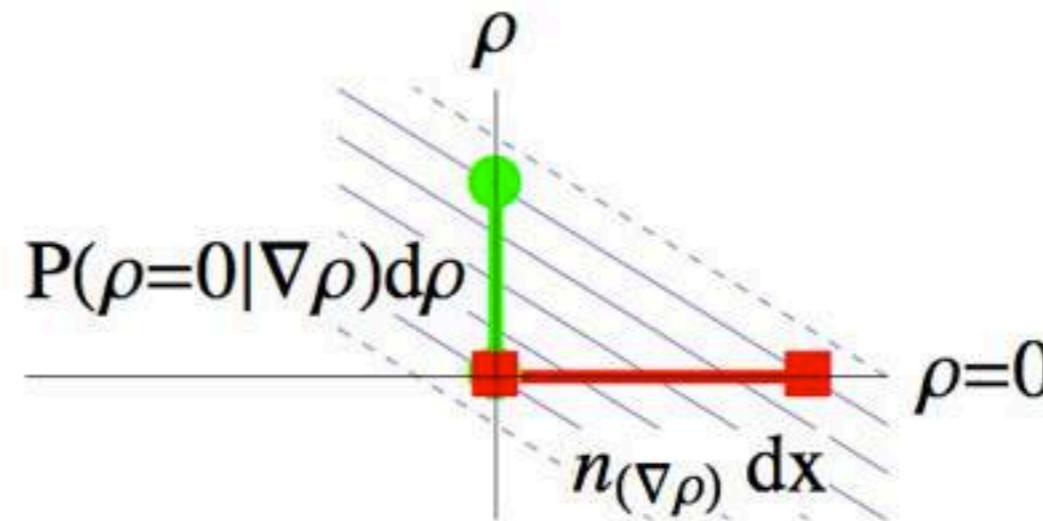
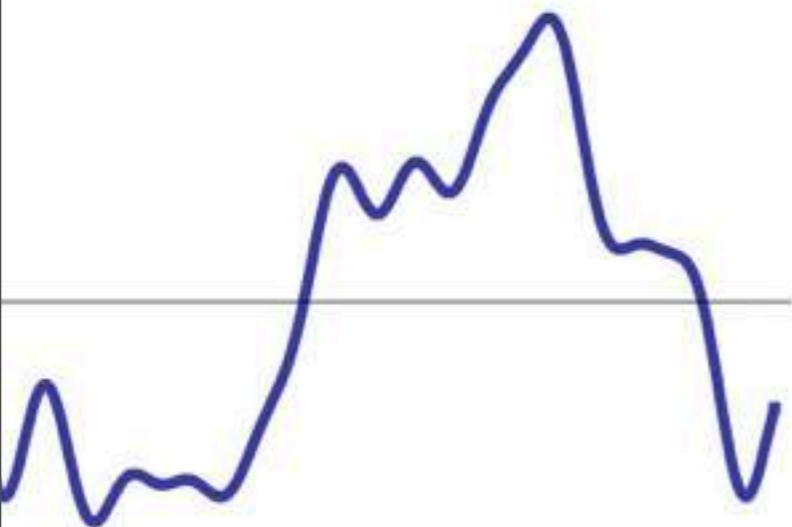
Are the imprints of LSS noticeable on galaxy properties ?



Horizon-AGN
>150 000 galaxies



1D level crossing primer



$$\left. \begin{array}{l} n_{\nabla\rho} dx = \mathcal{P}(\rho = 0 | \nabla\rho) d\rho \\ \frac{d\rho}{dx} = -\nabla\rho dx \end{array} \right\} \Rightarrow n_{\nabla\rho} = \mathcal{P}(\rho = 0 | \nabla\rho) |\nabla\rho|$$

$$n_{\text{zeroes}} = \int_{-\infty}^{\infty} d\nabla\rho |\nabla\rho| \mathcal{P}(\rho = 0, \nabla\rho) \propto \frac{\sigma_1}{\sigma_0}$$

$$n_{\text{extrem}} = \int_{-\infty}^{\infty} d\nabla\nabla\rho |\nabla\nabla\rho| \mathcal{P}(\nabla\rho = 0, \nabla\nabla\rho) \propto \frac{\sigma_2}{\sigma_1}$$

Summary of 1D Gaussian calculations

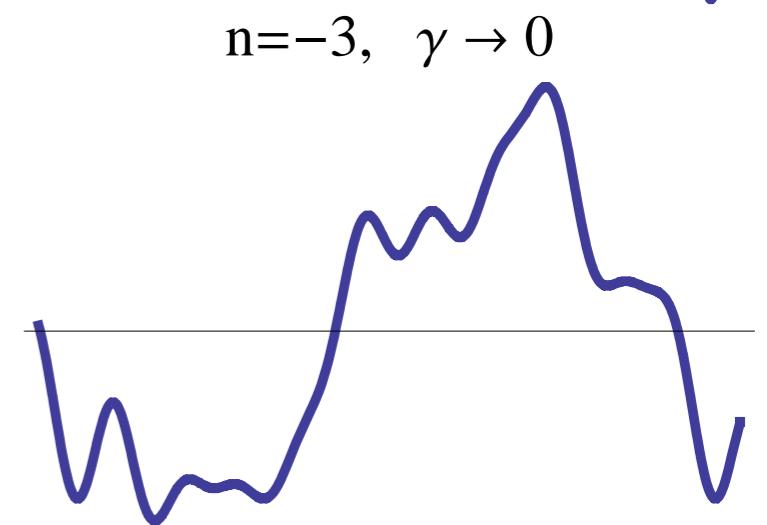
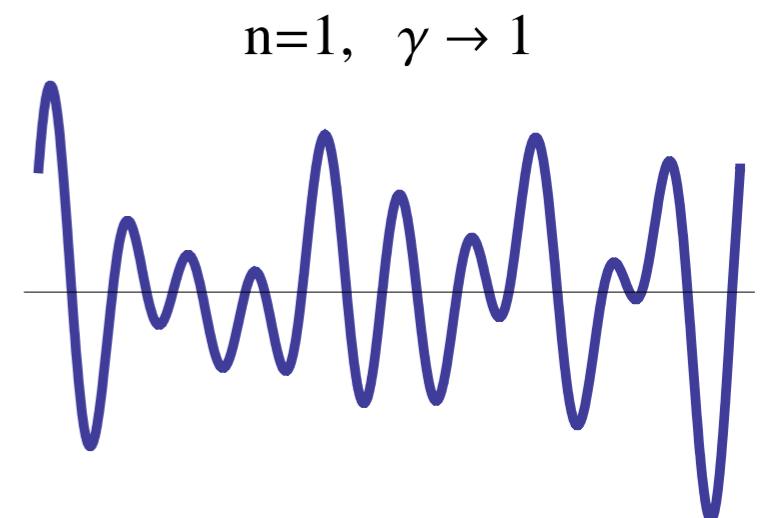
Fundamental scales R_0, R_*

$$n_{\text{zero}} \propto \frac{1}{R_0} = \frac{\sigma_1}{\sigma_0}, \quad n_{\text{maxima}} \propto \frac{1}{R_*} = \frac{\sigma_2}{\sigma_1}$$

$$\frac{n_{\text{maxima}}}{n_{\text{zero}}} = \gamma \equiv \frac{\sigma_1^2}{\sigma_0 \sigma_2} = \frac{R_*}{R_0}, \quad \gamma \in [0, 1]$$

More sophisticated result:
Maxima above threshold $\eta = \rho/\sigma_0$

$$n_{\text{max}}(\rho > \eta\sigma_0) = \begin{cases} n_{\text{max}} \mathcal{P}(\eta) & \gamma \rightarrow 0 \\ n_{\text{max}} \mathcal{P}(\eta)(\gamma\eta) & \gamma \rightarrow 1 \end{cases}$$



back to 2D Theory without Hessian approximation

$$L_k = \varepsilon_{ijk} I_{li} T_{lj}$$

$\approx \varepsilon_{ijk} \cancel{H}_{li} T_{lj}$

Tidal

Hessian

